

# **Neutrino detectors**

## **for Super-Beams, $\beta$ -Beams and $\nu$ Factories**

**Low-Z Tracking Calorimeter**  
**Water Čerenkov**  
**Liquid Argon TPC**  
**Magnetised Iron Calorimeter**  
**Emulsion Cloud Chamber**

### **For each technique:**

main issues on  
R&D  
and detector design

### **Physics potential:**

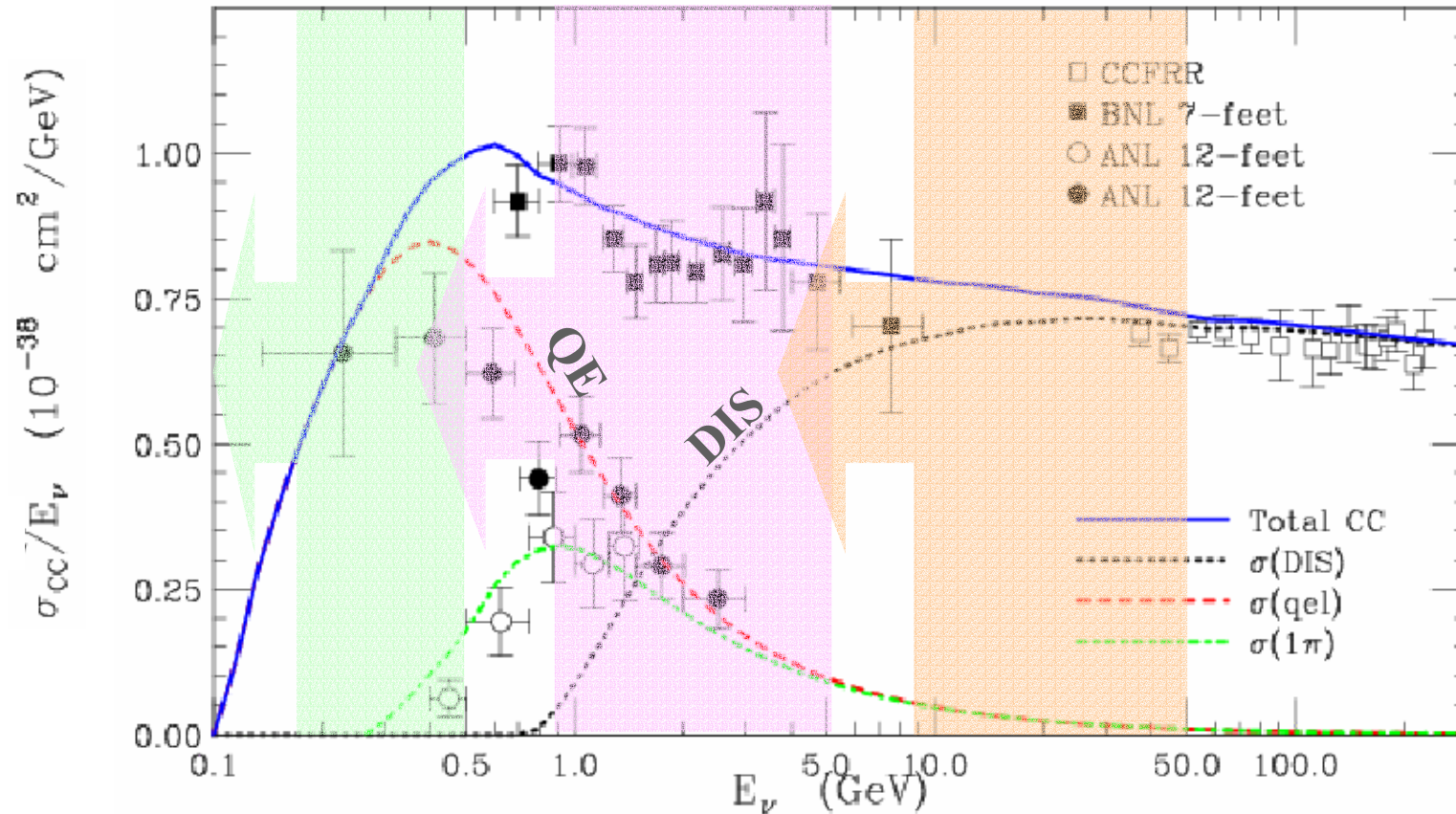
see specific talks  
(also depends on beam)

# Main detector tasks

- **Lepton identification** → identify  $\nu$  flavour
- **Measure  $\nu$  energy**
  - For Quasi-Elastics ( $\nu_l N \rightarrow l N'$ )  
Lepton energy →  $\nu$  energy  
with corrections due to Fermi motion in nucleus and to quark binding in nucleon
  - For DIS ( $\nu_l N \rightarrow l X$ ):  
measure energy of lepton and of  $X$
- **Measure muon charge (by magnetic field)**  
required at  $\nu$ -Factories → flavour of initial lepton ( $\nu_e$ - $\nu_\mu$  oscillation or  $\bar{\nu}_\mu$  ? )
- **Provide adequate target mass**  
50 kton – 1 Mton  
depending on physics aims ( $\theta_{13}$ , mass hierarchy,  $\delta$ ), exp. technique and beam

# $\nu$ energy ranges and related dominant reactions (event topologies)

P. Lipari, Nucl. Phys. Proc. Suppl. 112, 274 (2002) (NuInt01)



**$\beta$ -beam\***    **Super-beams**     **$\nu$ -Factory**

\*)  $E_\nu < \sim 10$  GeV if the radioactive ions are stored at LHC energy ( $\gamma \sim 1000$ )

# Envisaged baselines: 130 - 3000 km

(not discussed in this talk on detectors)

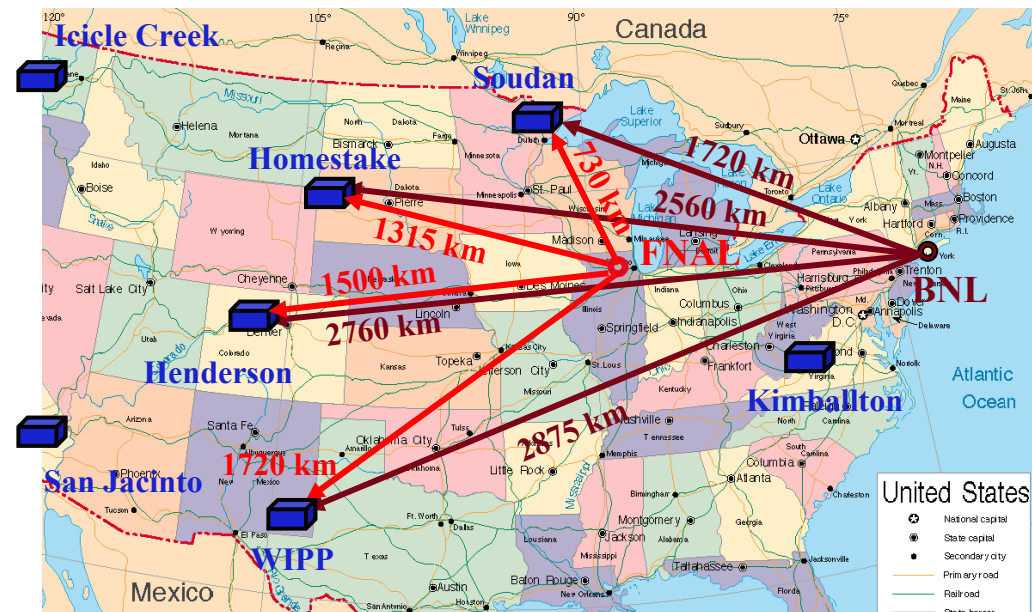
## Complexity of choice

L/E , matter effects, E



Detector design

Sensitivity to  $\theta_{13}$ , mass hierarchy,  $\delta$



# **Low-Z Tracking Calorimeter**

# A detector specialised for $\nu_\mu$ - $\nu_e$ oscillations

## Main backgrounds and tools for their reduction

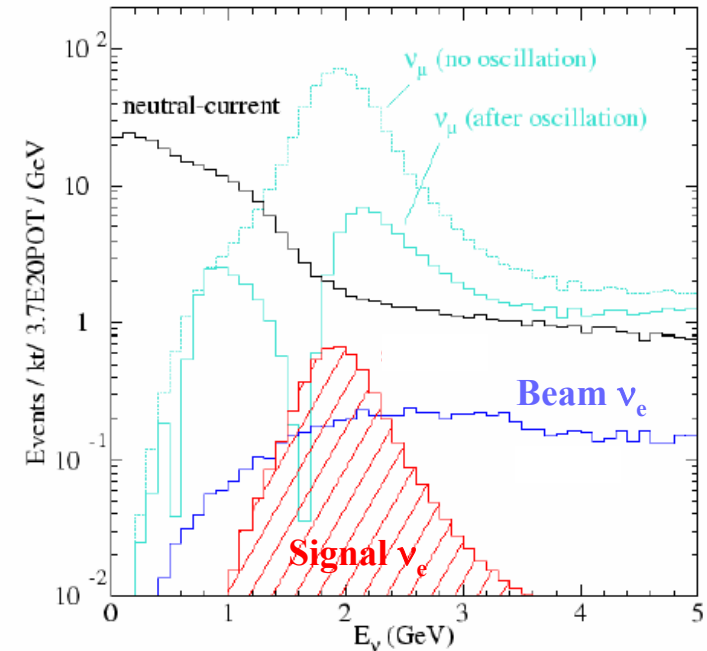
- Beam  $\nu_e$ 
  - $\Delta E/E$  combined with off-axis NBB
- $\pi^0$  in NC interactions
  - e/ $\pi^0$  separation by:
    - Tracking (e vs  $\gamma\gamma$ )
    - dE/dx at beginning of the shower (e vs  $e^+e^-$ )



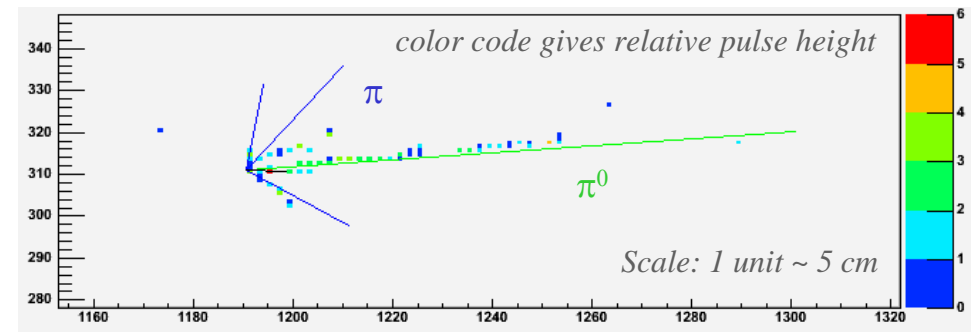
**Low Z** (small  $\Delta X^0$  sampling)  
**tracking calorimeter**  
 (CHARM II, ...NOE\* , NOvA at NuMI)

*e/h, e/ $\mu$  separation: turns out to be adequate*

\* 2.4 kton transition radiation detector proposed for CNGS



Neutrino spectra at NuMI off-axis



$\nu_\mu$  NC event in NOvA T ASD, with leading  $\pi^0$  and charged  $\pi$   
 (one projection)

# Main lines of NOvA design

*(Proposal P929 submitted on March 2004)*

## Evolution of proven experimental techniques

**Low Z, short sampling and mass 10 x MINOS → new technologies to reduce cost**

### Design features

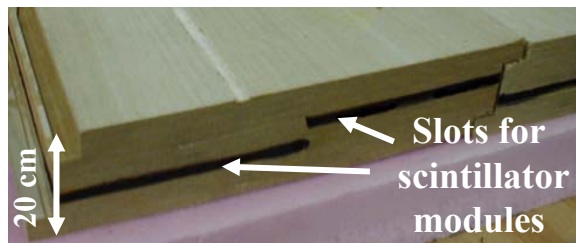
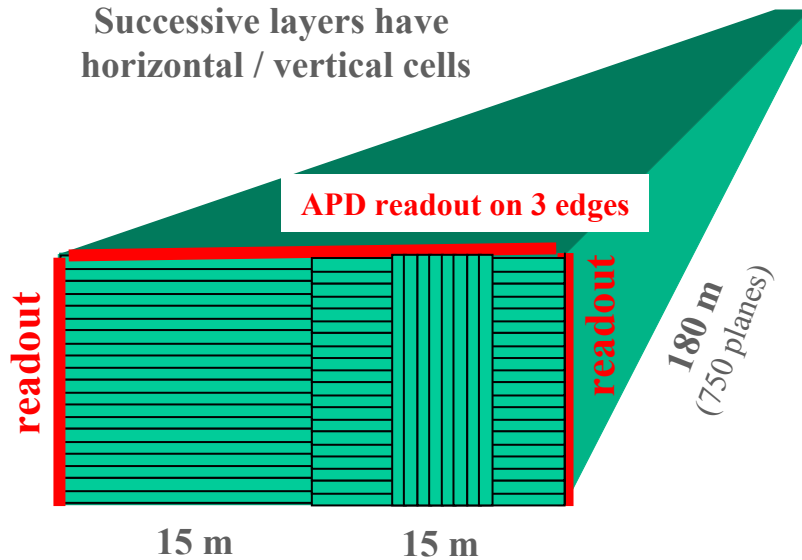
- Low Z and  $< 0.3 X_0$  sampling (Fe and  $1.5 X_0$  in MINOS)
- Liquid (plastic in MINOS) scintillator
- Avalanche Photo Diode (APD) readout (PMTs in MINOS) } Main new technologies wrt MINOS
- Two detector options
  - Baseline design:  $0.3 X_0$  sampling
  - Totally Active liquid scintillator Detector (TASD)  
(RPC as active detector option has been studied: two-dim. readout but no dE/dx)
- No need of underground location (live-time  $\sim 100$  s/year)  
Active shielding from cosmic rays foreseen (cheaper than passive overburden)

# Baseline Design

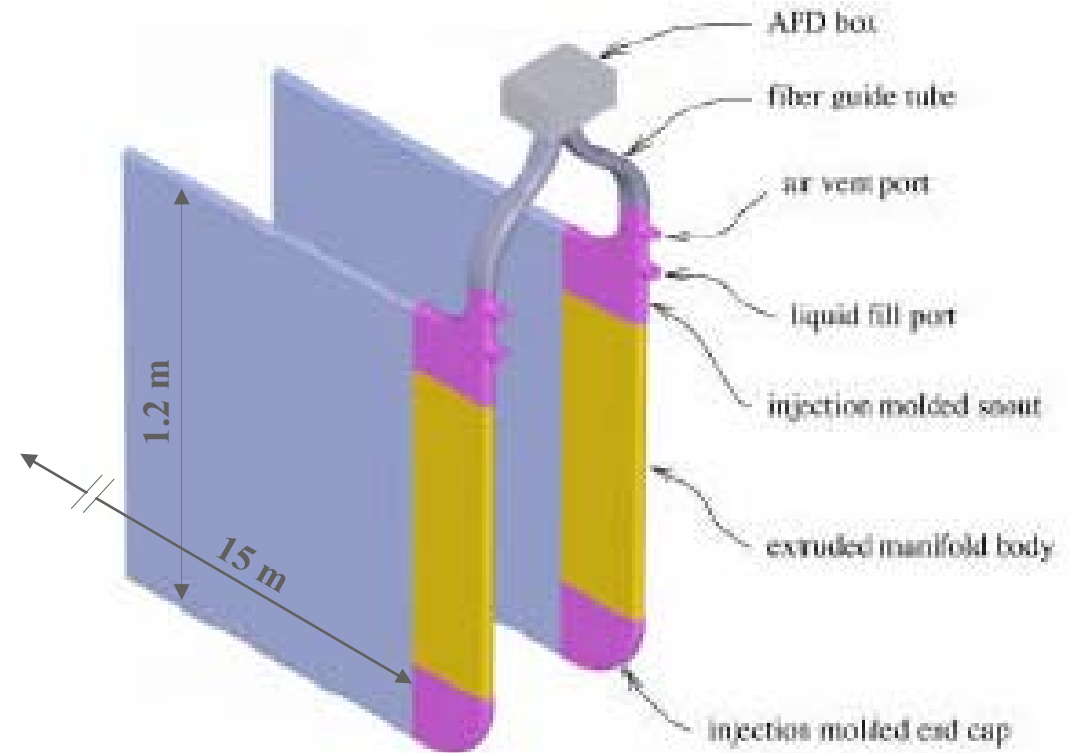
50 kton (7 kton scintillator, 0.3  $X_0$  sampling)

## View of the detector

Successive layers have horizontal / vertical cells

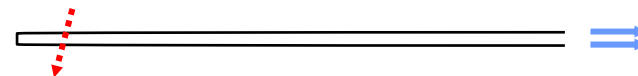


Particle-board as passive material (provides also the support structure)



## Liquid scintillator modules

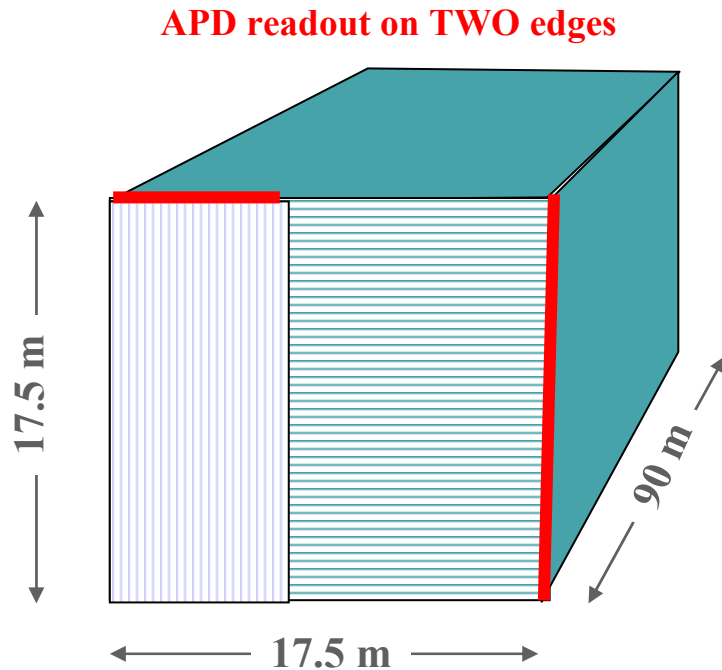
- 1.2 m wide, 15 m long PVC extrusions with 30 cells (3 cm wide)
- U-loop Wavelength Shifting Fiber ( $\phi = 0.8$  mm) in each cell: ~ doubled light collection from far end





# Totally Active liquid Scintillator Detector (TASD)

25 kton (21 kton scintillator)



The detector is wider and taller than the baseline detector but shorter along the beam

No crack down the center

With respect to the baseline design:

- **Similar scintillator modules**
  - thicker cells along the beam  
4.5 cm vs. 2.56 cm (more light)
  - Longer extrusions  
17.5 m long vs. 15 m (less light)
  - 32 cells wide vs. 30 cells: matches 16 ch. APD
- **Same U-Loop WLS fibres**
- **Same APD readout**  
but only on two detector edges
- **PVC must provide a self-supporting structure:** mechanics to be carefully studied (a PVC 5-story building)
- **85% scintillator, 15% PVC**  
~ **same cost as with baseline design**  
**implies ½ detector mass**

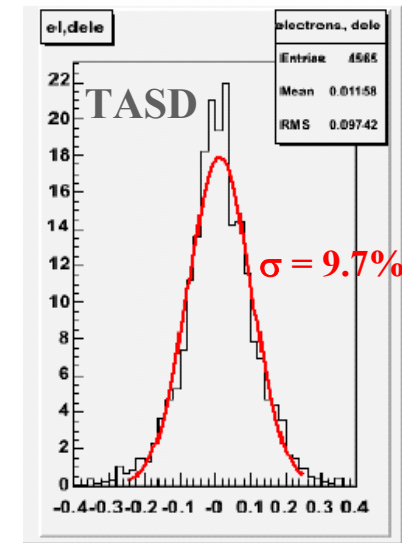
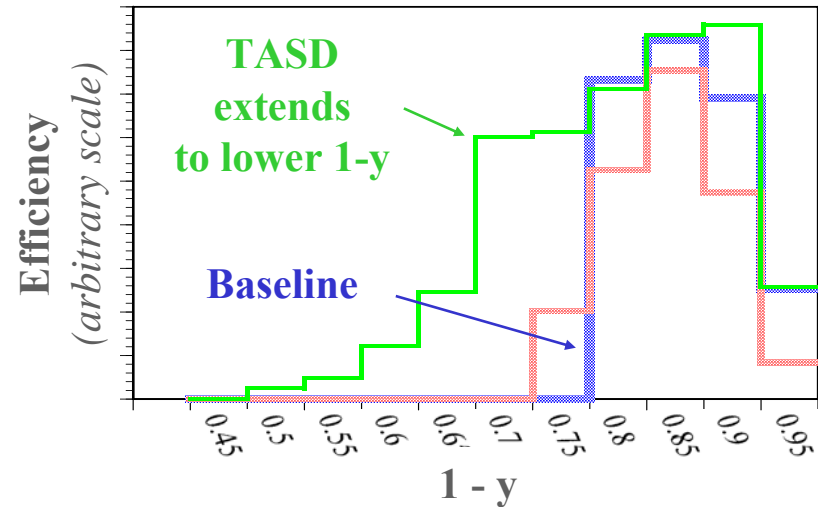


*see next slide*

# Baseline design (0.3 X<sub>0</sub> sampling) → T ASD

- 50 kton → 25 kton but higher oscillated  $\nu_e$  efficiency 18% → 32%
- T ASD has ~ 4 times as many hits / unit track length
  - 2.5cm liquid + 17.8 cm particle-board → 4.5cm liquid + 0.4cm PVC
  - pulse height information in every sample (favours  $e/\pi^0$ )
- T ASD has better energy resolution

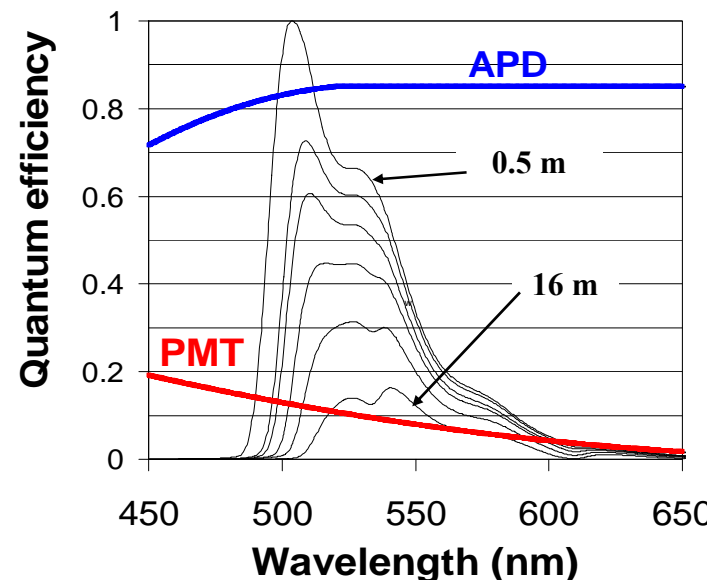
For  $e^-$   $\Delta E/E = 15\% / \sqrt{E} \rightarrow 10\% / \sqrt{E}$   
(23% /  $\sqrt{E}$  with RPCs)
- About the same cost and time scale (detector completed in late 2011 if funding begins in late > 2006)



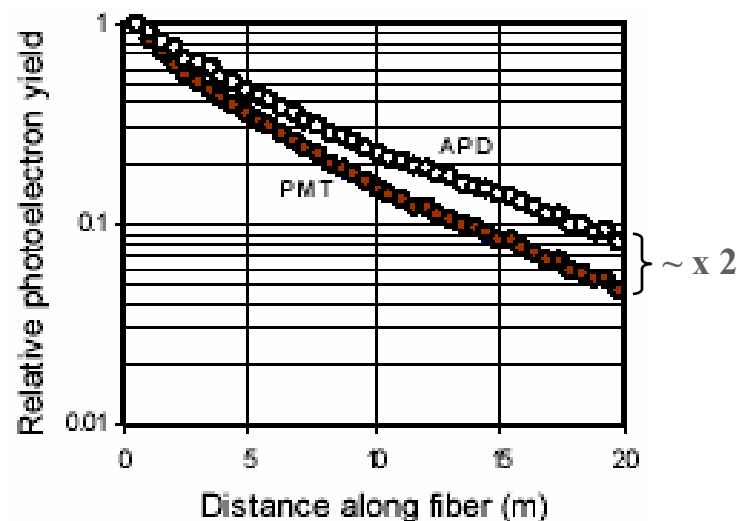
$\text{Meas-True}/\sqrt{\text{True}}$

# Avalanche Photo Diodes (APD)

- Cheaper than PMTs
- 2x16 pixel APDs commercially produced at large scale:  
CMS/LHC  $n^0$  pixels  $\sim 8 \times$  NOvA
- High QE ( $\sim 85\%$ )  
→ longer scintillator cells  
( $\sim 40$  photons/mip from far-end)
- Spectral shift at far-end improves the relative yield
- Dark noise from thermally generated electron-hole pairs: reduced to  $\sim 10$  e $^-/\mu$ s by cooling (Peltier effect) at  $-15$  °C
- Low gain (operated at  $\sim \times 100$ )  
→ need of stable and reliable amplifiers



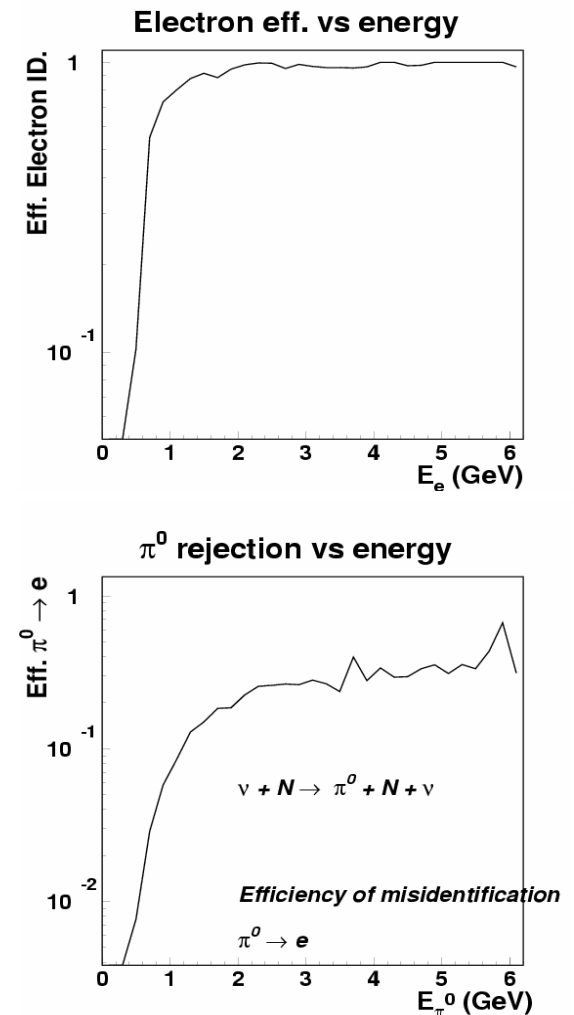
(spectra at various distances from photo-detector)



# Water Čerenkov

# Water Čerenkov

- **Sensitivity to low energy** (if PM density is adequate)  
→  $\nu$  oscillation,  $\nu$  astrophysics, proton decay
- **Difficulties at high energy**
  - Frequent multi-ring events, DIS dominates
  - $e/\pi^0$  separation:  $\sim 30\%$   $\pi^0$  mis-identification above a few GeV (because of  $\gamma\gamma$  collimation)
  - Hence not suitable at  $\nu$ -Factory (where, moreover, the muon charge measurement would be needed)
- **Proven technique**
  - A third generation of successful underground detectors:  
**IMB / Kam  $\rightarrow$  Super-Kam  $\rightarrow$  Hyper-Kam / UNO**  
In each generation one order of magnitude increase in mass
  - Performance as well as limitations  
known from SK and K2K, extrapolated by MC

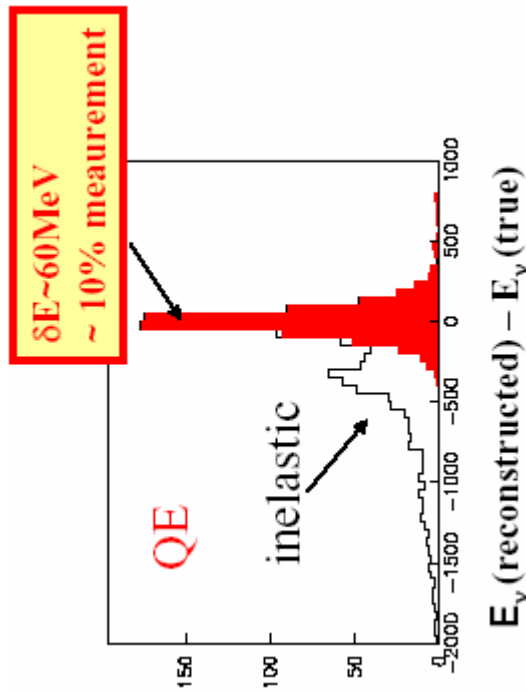


From Report BNL-69395

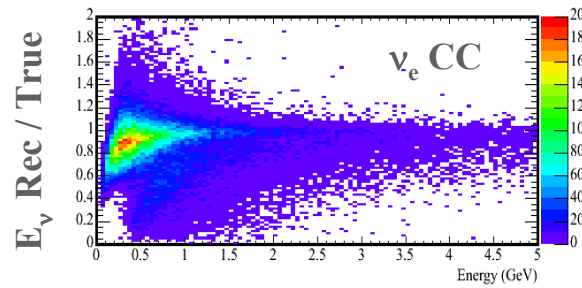
*Super-K: a large Water Čerenkov detector of which the performance has been simulated and observed*

# Energy resolution in Super-Kamiokande

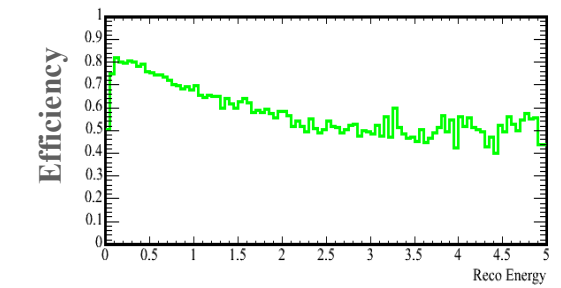
M. Messier



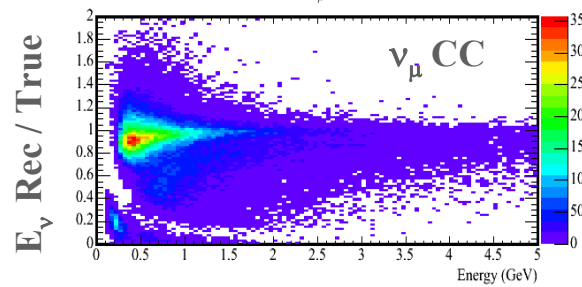
Reconstructed Energy vs True Energy for  $\nu_e$  CC Events



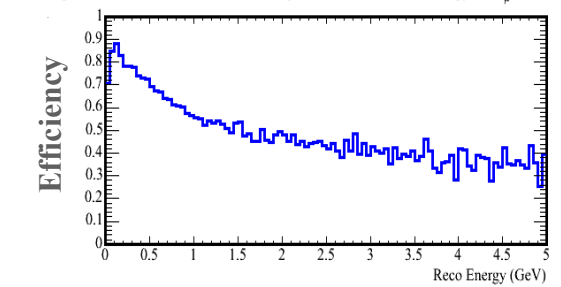
1-Ring, e-Like Reconstruction Efficiency vs Reconstructed Energy for  $\nu_e$  CC Events



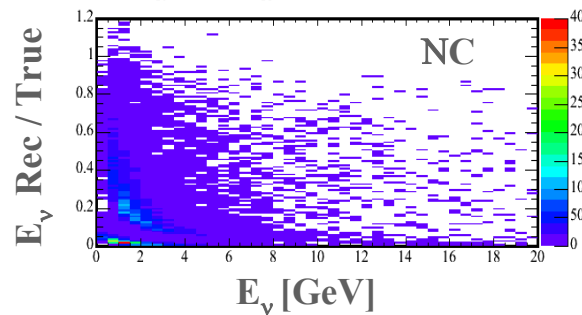
Reconstructed Energy vs True Energy for  $\nu_\mu$  CC Events



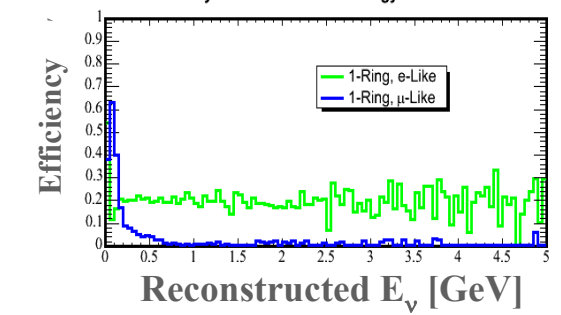
1-Ring,  $\mu$ -Like Reconstruction Efficiency vs Reconstructed Energy for  $\nu_\mu$  CC Events



Reconstructed Energy vs True Energy for NC Events



Reconstruction Efficiency vs Reconstructed Energy for NC Events



$E_\nu$  reconstructed accounts for  $p_\mu$ ,  $\theta_\mu$  and, in case of DIS, for  $E_{had}$

# Geometry of next Water Čerenkov generation

- Max 50 m water depth pressure, with current 20” Hamamatsu PMTs
- ~ 80 m light attenuation length in pure water
- Mining cost  $\propto$  total detector volume
- Instrumentation cost  $\propto$  detector surface area
- Cavern should have rounded edges, to reduce rock stresses



## Elongated shape

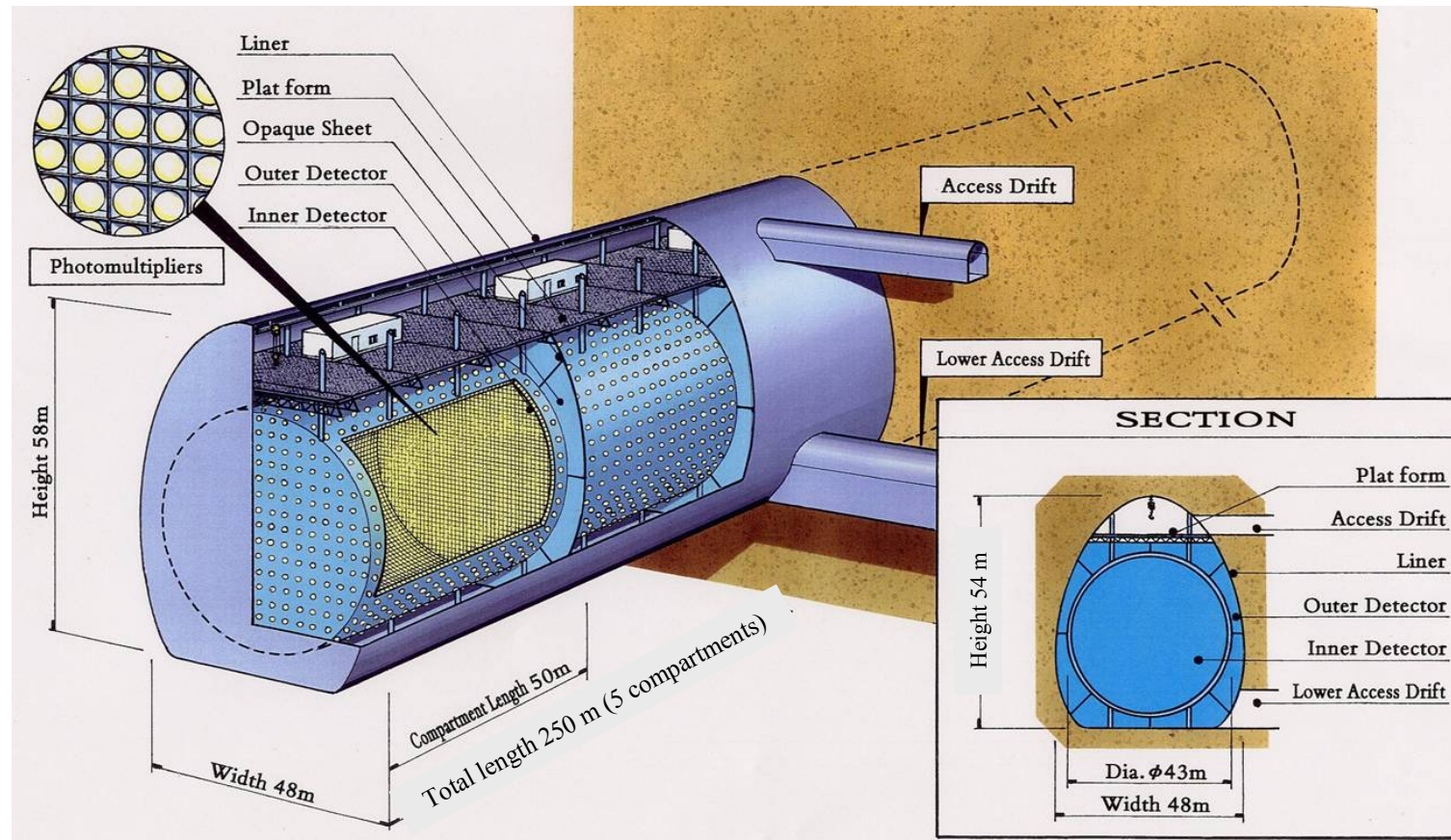
(transverse dimensions  $\sim$  60 m)

## Segmentation

- reduction of backgrounds from PMT discharges
- increased operational live-time due to independent module calibration (and maintenance, in case of complete separation)



# Hyper-Kamiokande

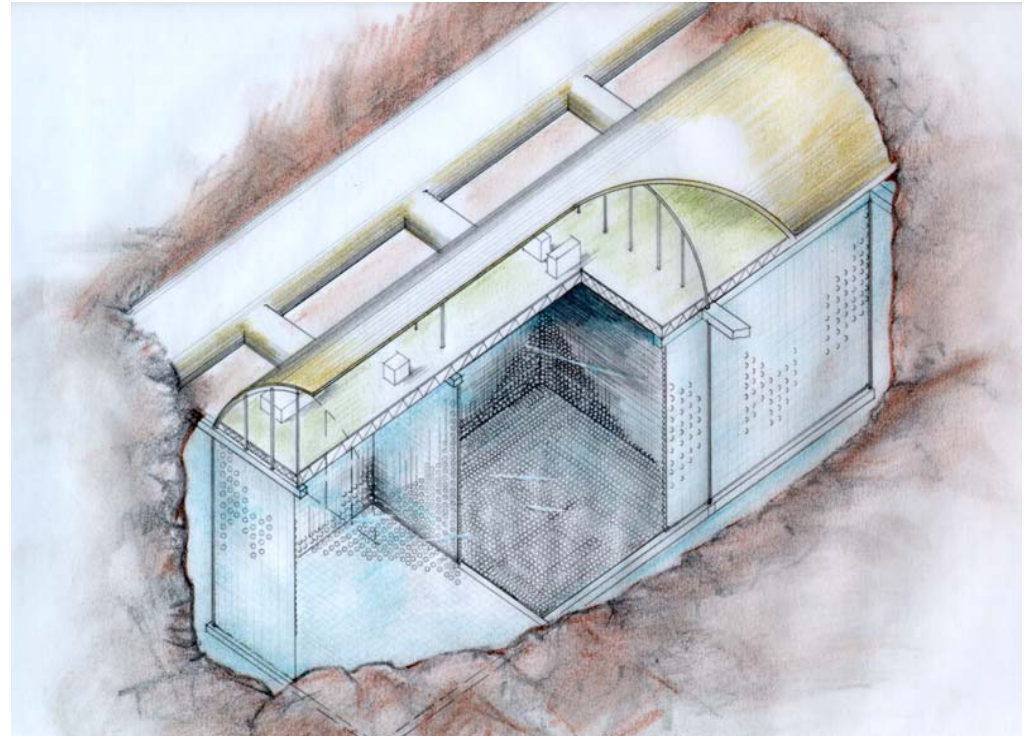
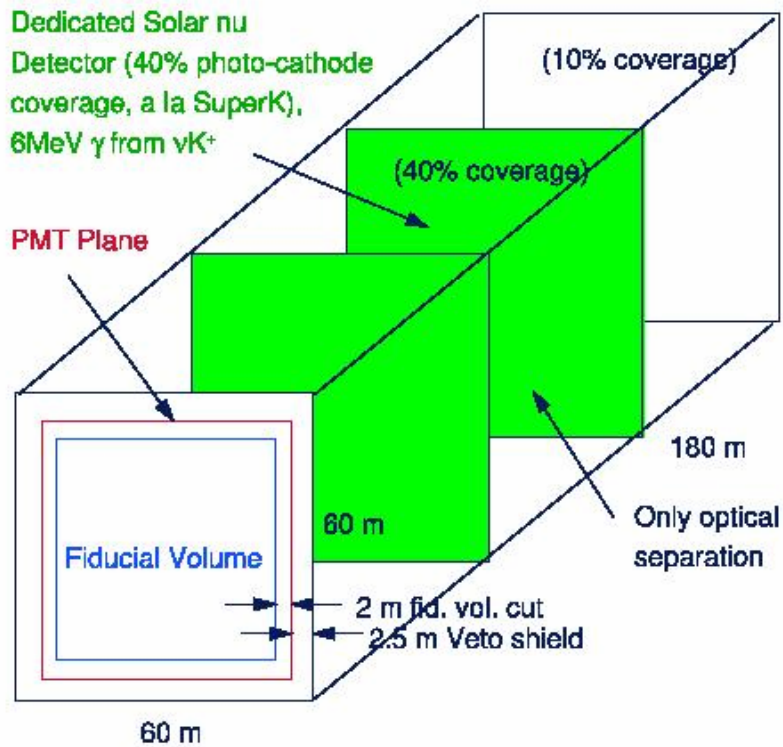


## One of the two 500 kton modules

- Two modules placed sideways, each with 5 compartments (50 m long)
- Higher cost than for a single module, but maintenance possible with one module always alive
- Both cavities should be excavated at the same time. But a staging scenario is possible.
- ~ 10 years construction time,  $t_0$  wished after a few years of operation of T2K



# UNO



- **Total (fiducial) mass 650 (445) kton**
- **Three zones with different photo-sensor density, to reduce cost**
  - middle zone high density (40% PMT coverage like Super-K) for N decay and solar  $\nu$
  - edge zones with 10% PMT coverage: suitable for atmospheric and beam  $\nu$
- **56,650 20" PMTs ( $\sim \frac{1}{2}$  wrt to full 40% coverage); 15,000 8" PMTs**

# Comparison to Super-Kamiokande

	<b>S-K</b>	<b>Hyper-K</b>	<b>UNO</b>
Total mass [kt]	50	2 x 500	650
Fiducial mass [kt]	22.5	2 x 270	440
Fiducial mass (solar $\nu$ ) [kt]	22.5	2 x 270	145
Size [m <sup>3</sup> ]	$\Phi$ 41 m x 39 m	2 x $\Phi$ 43 m x 250 m	60 m x 60 m x 180 m
photocathode coverage [%]	40	40	$\frac{1}{3}$ 40 (5 MeV threshold) $\frac{2}{3}$ 10 (10 MeV threshold)
PMTs	<b>11,146</b> (20")	<b>200,000</b> (20")	<b>56,650</b> (20") 15,000 (8")

**A large fraction (  $\sim \frac{1}{2}$  or more, in case of PMTs)  
of the total detector cost comes from the photo-sensors**

# Increasing the mass ...

- **Champagne:**

taste improves from Standard to “Magnum” to “Jeroboam” bottles  
(lower surface/volume)

- **Water Čerenkov:**

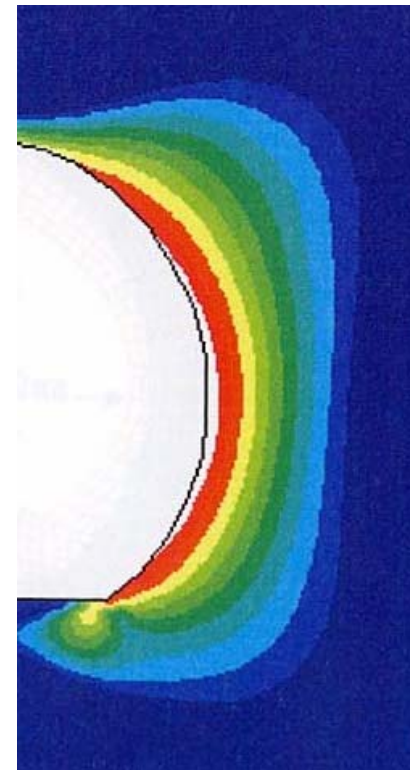
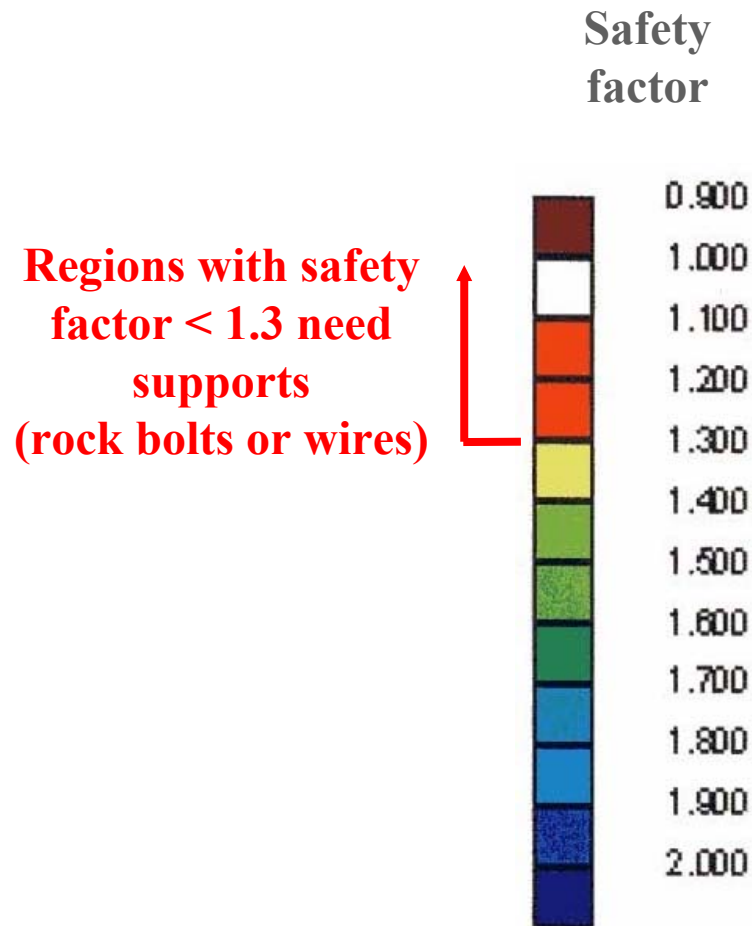
- better energy containment
- larger “effective” granularity of photo-sensors  
(due to larger average distance from event vertex)

## Main issues

- Design of a large cavern
- R&D on photo-sensors, in strong collaboration with industry\*, to improve:
  - cost
  - production rate: ~ 8 years for 100,000-200,000 PMTs (with related storage problems)

\* the development of 20” PMTs is at the basis of the success of Kamiokande and Super-Kamiokande

# Finite element analysis of Hyper-K cavern



*Pressure hor/vert = 0.45*

**Depth of region with safety factor < 1.3 similar to that in Super-K**



**Possible to excavate the Hyper-K cavity**

# R&D on photo-sensors for Water Čerenkov

- PMTs

R&D with industry to reduce cost, speed-up production rate, improve performance

- New photo-sensors

- Spherical hybrid photo-sensor (HPD)

ICRR Tokyo - Hamamatsu

- “ReFERENCE” tube: photo-cathodes operating in reflection mode

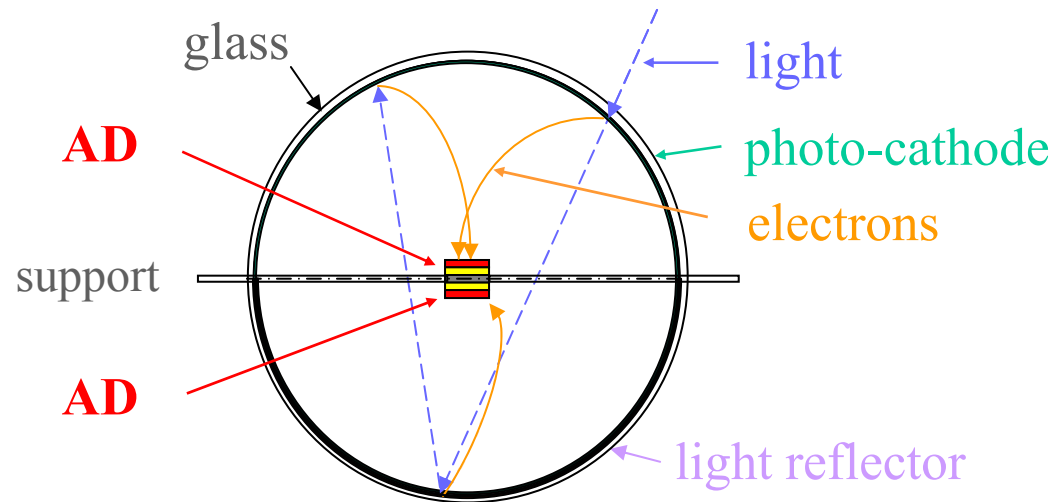
U.C. Davis - ITT Night vision, at an early stage of development

**General comment:**

**Long term stability and reliability are a must**

**Proven for PMTs**

# Spherical HPD

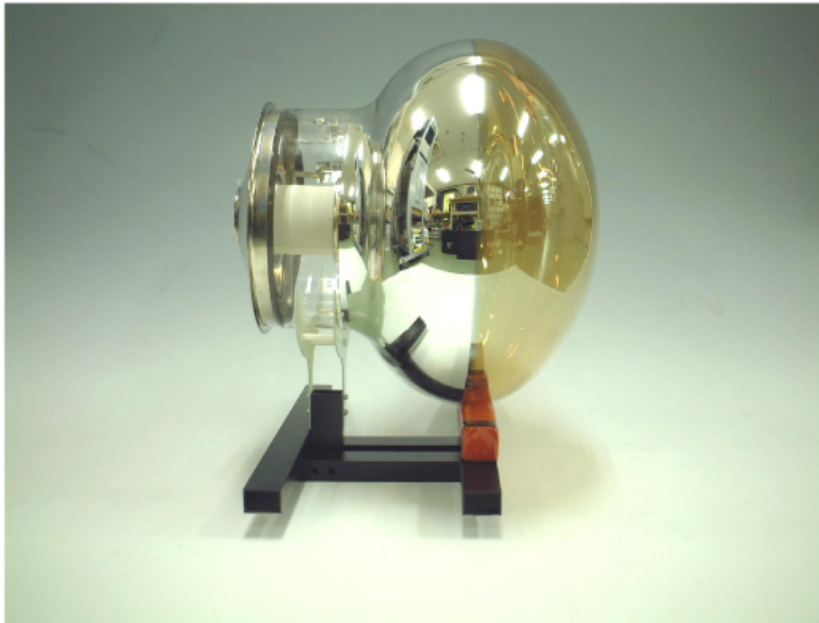


- Spherical glass envelope, coated with photo-cathode and light reflector
- **Electrons accelerated by 20 KV between photocathode and AD**
- **Avalanche Diodes (AD)**
  - strong amplification in the electron bombardment of AD
    - single photon sensitivity from the high gain ( $\sim 4000$ ) in this 1st amplification stage
    - the noise thermally generated in the AD itself becomes ineffective
  - resolve 2-3 events per 50 ns
  - overall gain ( $\sim 10^5$ ) lower than with PMTs: need of stable and highly reliable amplifiers

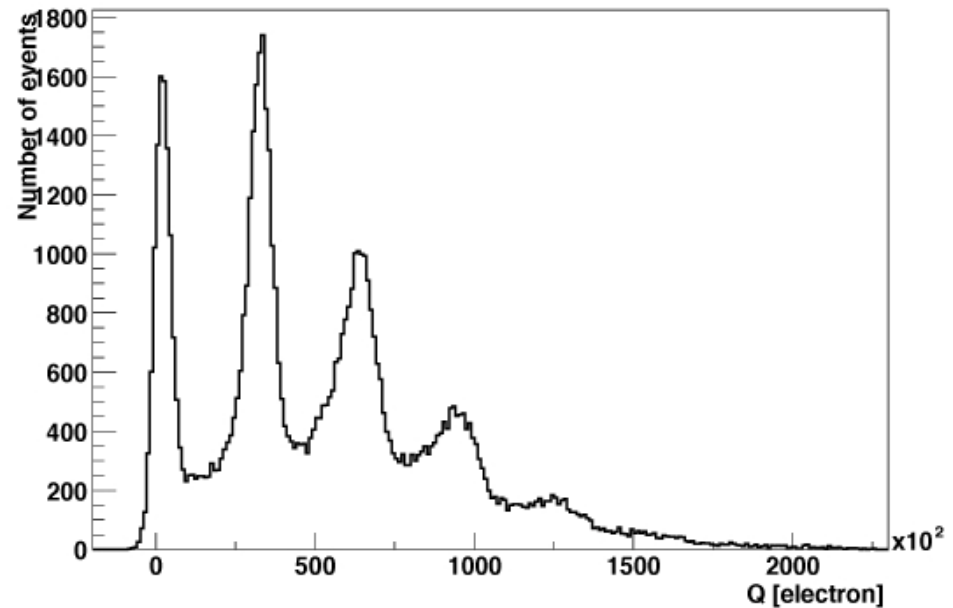
➡ **Cost reduction from AD instead of PMT dynode structure**

# R&D on HPDs

13 Inch-Dia. HPD



**HAMAMATSU**  
HAMAMATSU PHOTONICS K.K., Electron Tube Center

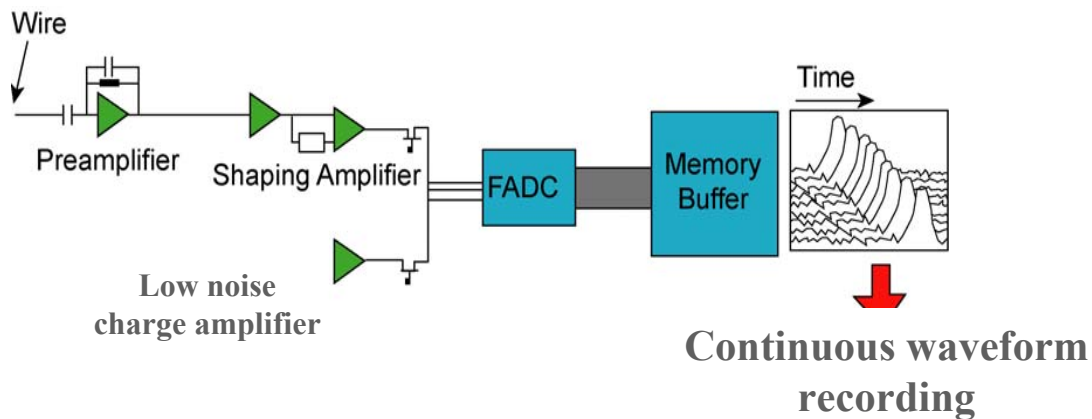
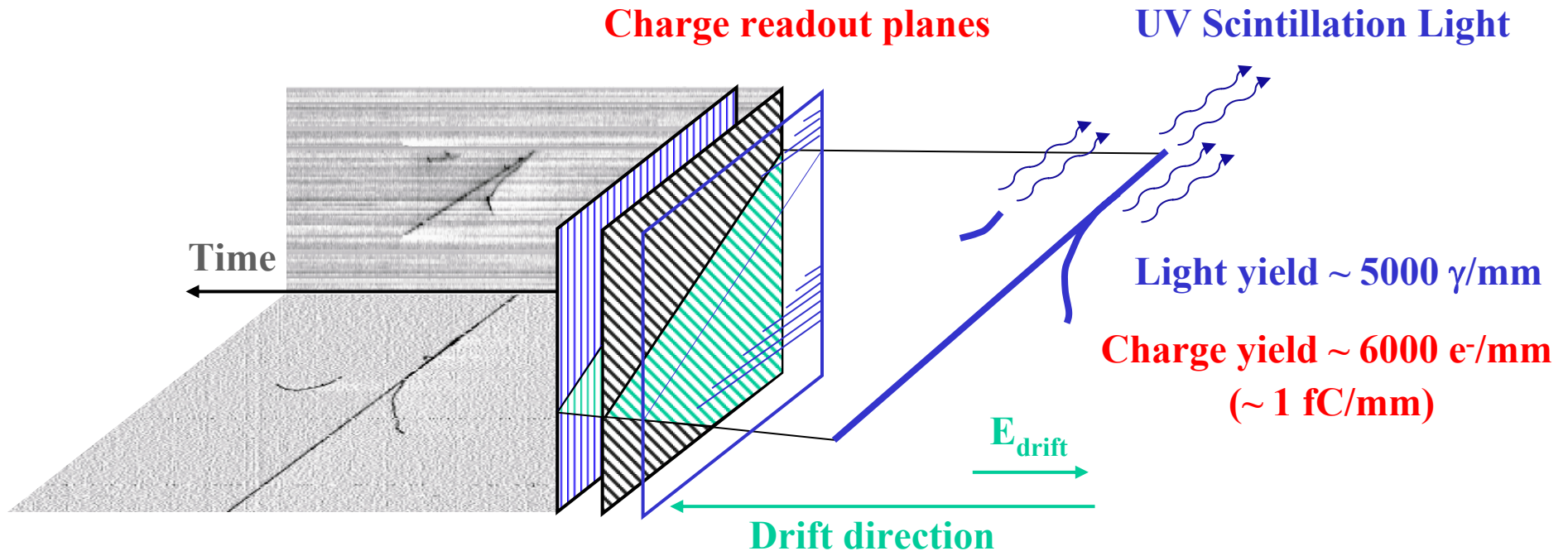


- 5" prototype tested
- Tests of 13" prototype in progress:  
gain  $>10^5$ , Transit Time Spread  $\sim 1$  ns, single photon sensitivity
- Amplifier, Digital Filter, Analog Memory Cell: in progress
- 13" production model by spring 2005
- Next: design of Spherical 20" HPDs (requires higher field or development of larger AD)

**Liquid Argon  
Track Projection Chamber**



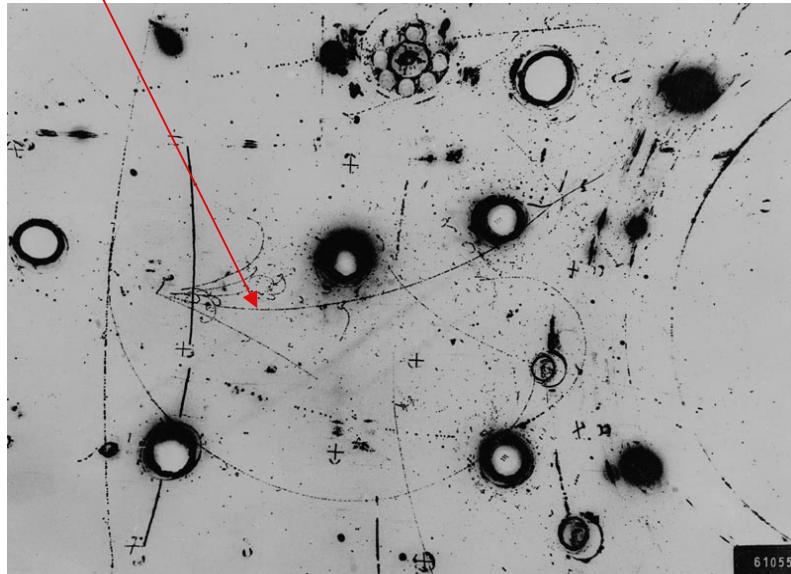
# The Liquid Argon TPC principle



**High density**  
**Non-destructive readout**  
**Continuously sensitive**  
**Self-triggering**  
 **$t_0$  from scintillation**

# ...an “electronic bubble chamber” with broad physics potential and energy range

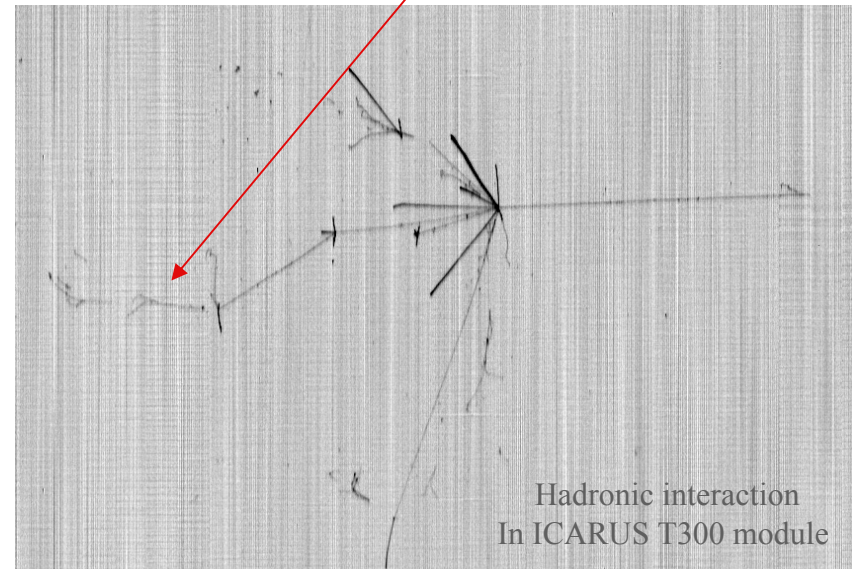
Bubble diameter  $\approx 3$  mm (diffraction limited)



**Gargamelle**

Medium	Heavy Freon
Sensitive mass	<b>3 ton</b>
Density (g/cm <sup>3</sup> )	1.5
Radiation length (cm)	11.0
Collision length (cm)	49.5
dE/dx (MeV/cm)	2.3

Bubble size  $\approx 3 \times 3 \times 0.4$  mm<sup>3</sup>



**Liquid Ar TPC**

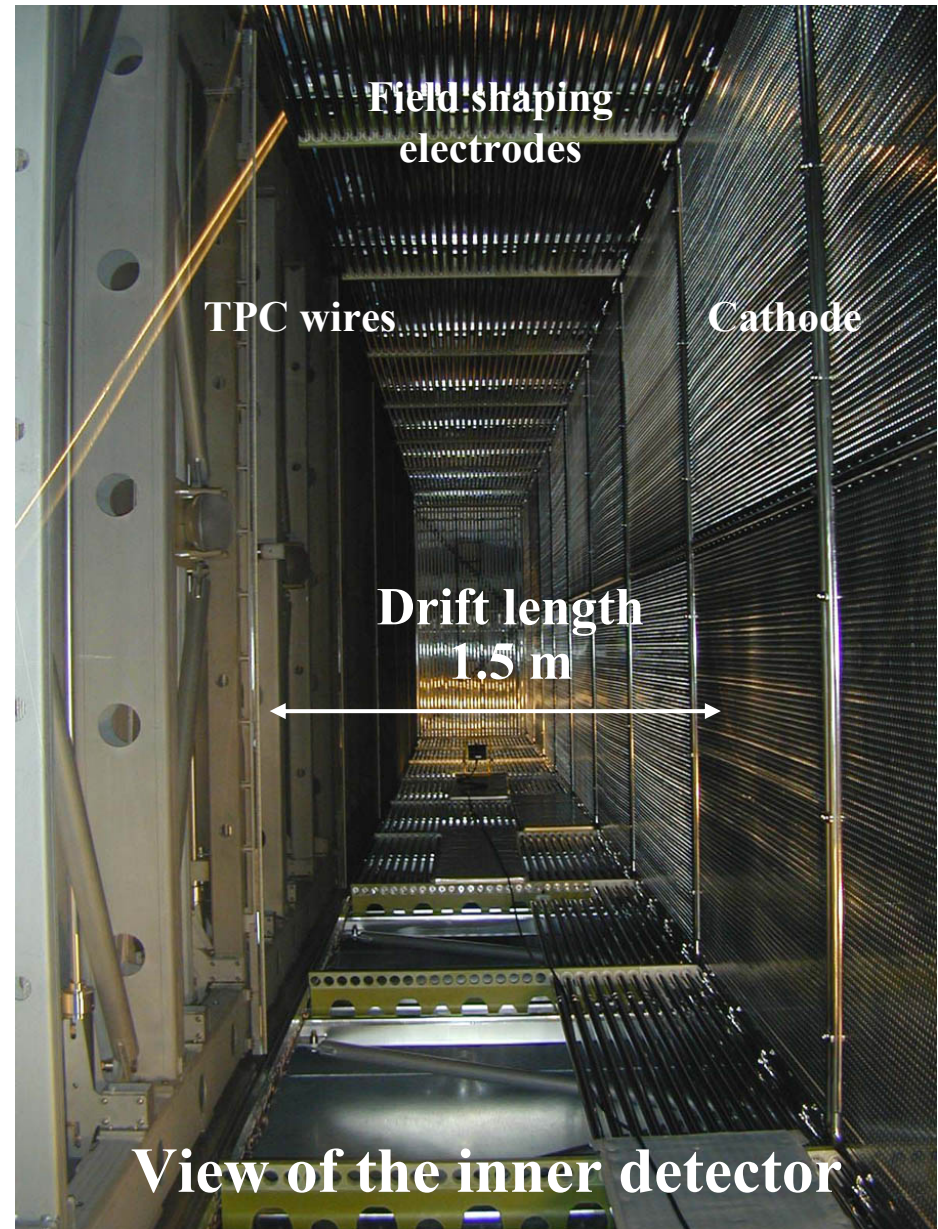
Liquid Argon
<b>300 ton - .... kton</b>
1.4
14.0
54.8
2.1
<b>continuously sensitive</b>



# ICARUS T300 module (0.3 kton)



- Operated in a ground level laboratory (Pavia)
- “The present status of the art”



**T300: 300 ton, with 1.5 m drift length**  
**ICARUS design: a ~ 2.5 kton modular detector**

## **To reach a 50-100 kton mass:**

- **Module size and geometry**

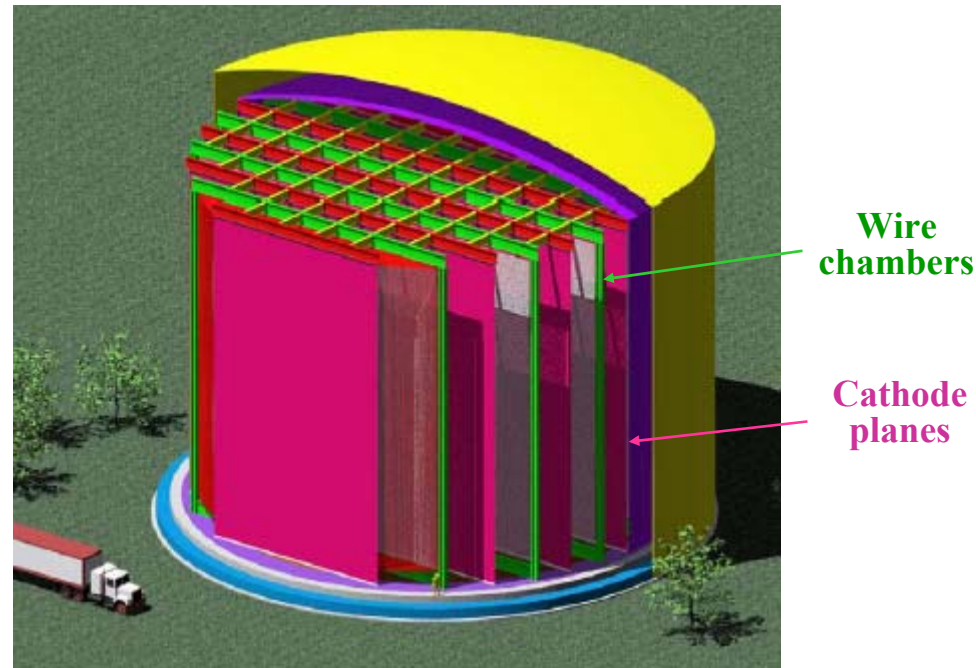
- Cryogenic insulation imposes constraints
- Do not pursue the ICARUS multi-module approach, design a single large module
- Aspect ratio ~ 1:1 for larger volume/surface ratio

- **Drift length: larger to limit the number of readout channels**

- Drift lengths ~ 5 m and readout in Liquid Argon as in ICARUS (FLARE and LANND)
- Very long drift lengths (~ 20 m) in Liquid Argon, with amplification and (“Double Phase”) readout in Gas Argon
- Purity and signal attenuation to be kept under control

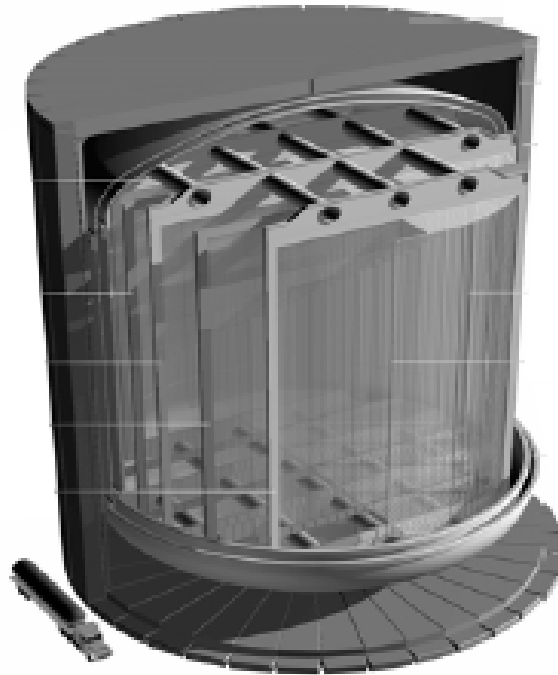
# An off-axis Liquid Ar TPC for the NuMI beam

(FLARE, Lol in preparation)



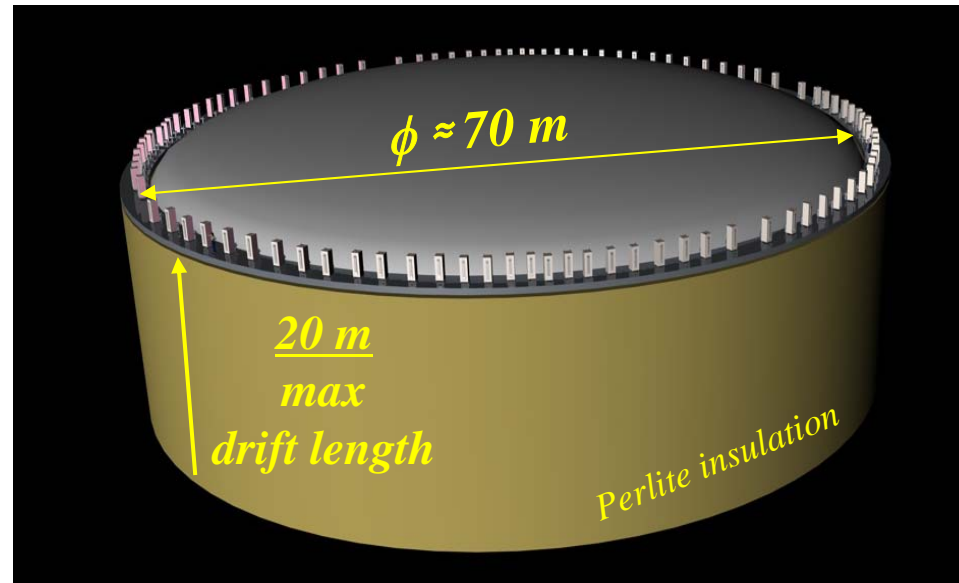
- Readout in liquid phase, as in ICARUS
- 50 kton:  $\sim 1.5 \times 10^2$  extrapolation in mass wrt ICARUS T300
- 3 m max drift length (1.5 m in ICARUS T300)
- Surface location (operated only with beam, for  $\nu$  oscillation studies)

# LANNDD



- **A detector for  $\nu$  oscillation, as well as (if located underground) for  $\nu$  astrophysics and proton decay**
- **Readout in liquid phase, as in ICARUS**
- **50-100 kton : 1.5- 3 x 10<sup>2</sup> extrapolation in mass wrt ICARUS T300**
- **4-8 m max drift length (1.5 m in ICARUS T300)**
- **Tests foreseen at 5 m drift length**

# A 100 kton Liquid Argon TPC detector with “Double-Phase” readout



*A giant Liquid Argon scintillation, Čerenkov and charge imaging experiment*  
A.Rubbia, Proc. II Int. Workshop on Neutrinos in Venice, 2003

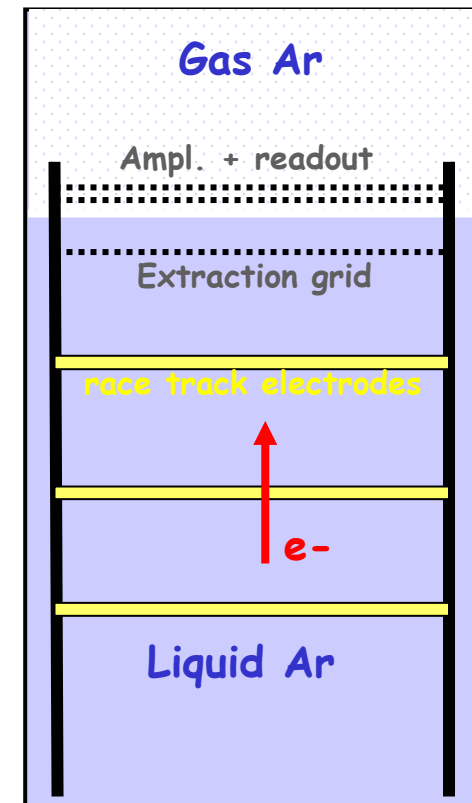
- **A detector for  $\nu$  oscillation,  $\nu$  astrophysics, proton decay**
- **20 m max drift length**
- **Use of Liquid Natural Gas storage techniques**  
(technical study with industry)
- **$\sim 3 \times 10^2$  extrapolation in mass wrt ICARUS T300**



# Double-phase (Liquid – Gas) readout

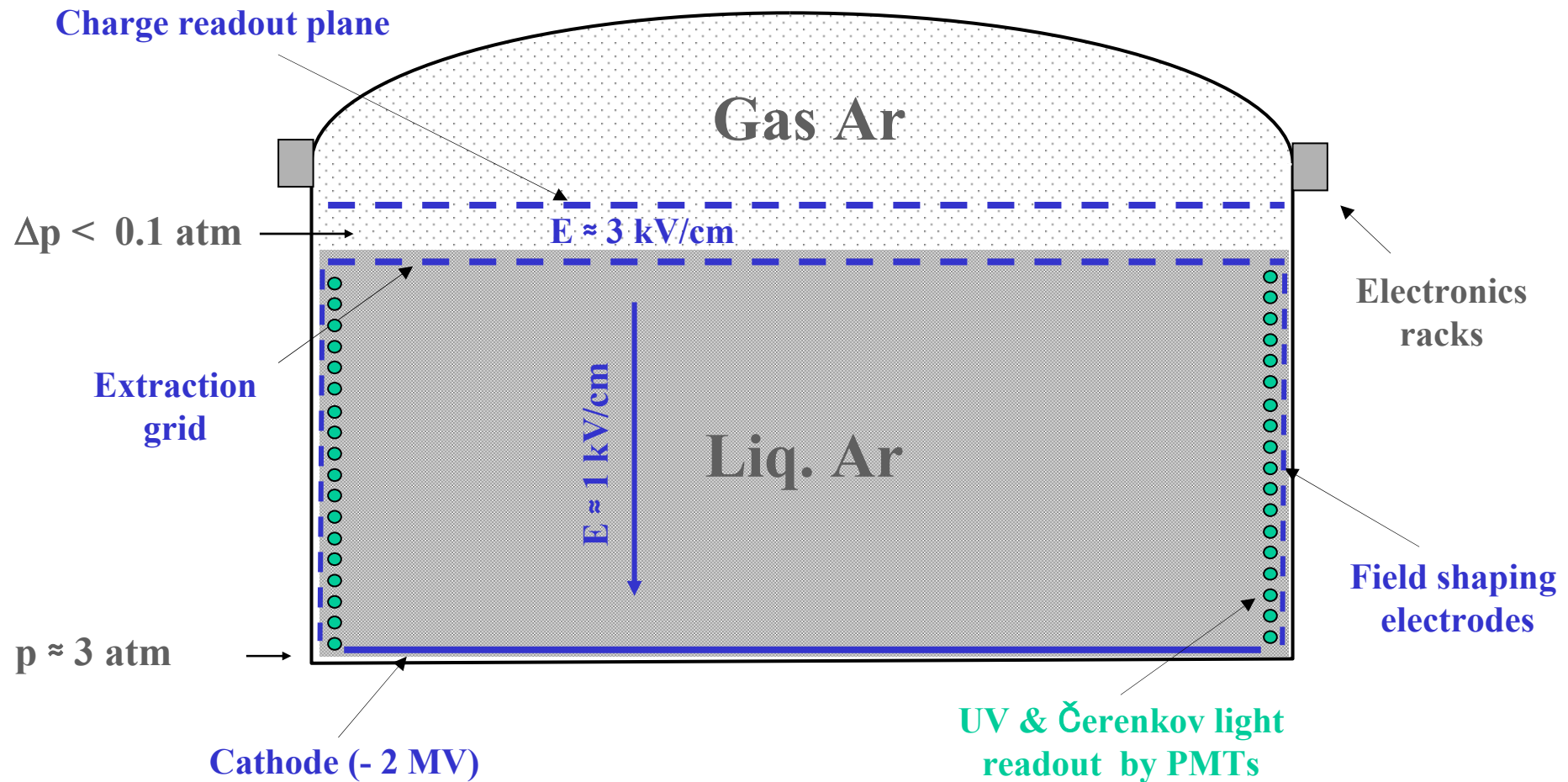
## Charge attenuation after very long drift in liquid compensated by charge amplification near anodes in gas phase

- Basic references: Dolgoshein et al. (1973); Cline, ... Picchi ... et al. (2000)
- Tested on the ICARUS 50 litres chamber
- $v_d \approx 2 \text{ mm}/\mu\text{s}$  at 1 kV/cm: drift time  $\approx \underline{10 \text{ ms}}$   
over 20 m max drift length, high purity required
- With 2 ms e-lifetime, max charge attenuation is  $e^{-t/\tau} \approx 1/150$   
(allowed because 6000 e- / mm for a MIP in LAr)
- Electron extraction to gas phase
- Amplification in proportional mode (x 100-1000) and readout:  
thin wires ( $\phi \approx 30 \text{ }\mu\text{m}$ ) with pad readout, .....
- Diffusion after 20 m drift:  
 $\sigma \approx \underline{3 \text{ mm}}$  gives a limit to the practical readout granularity





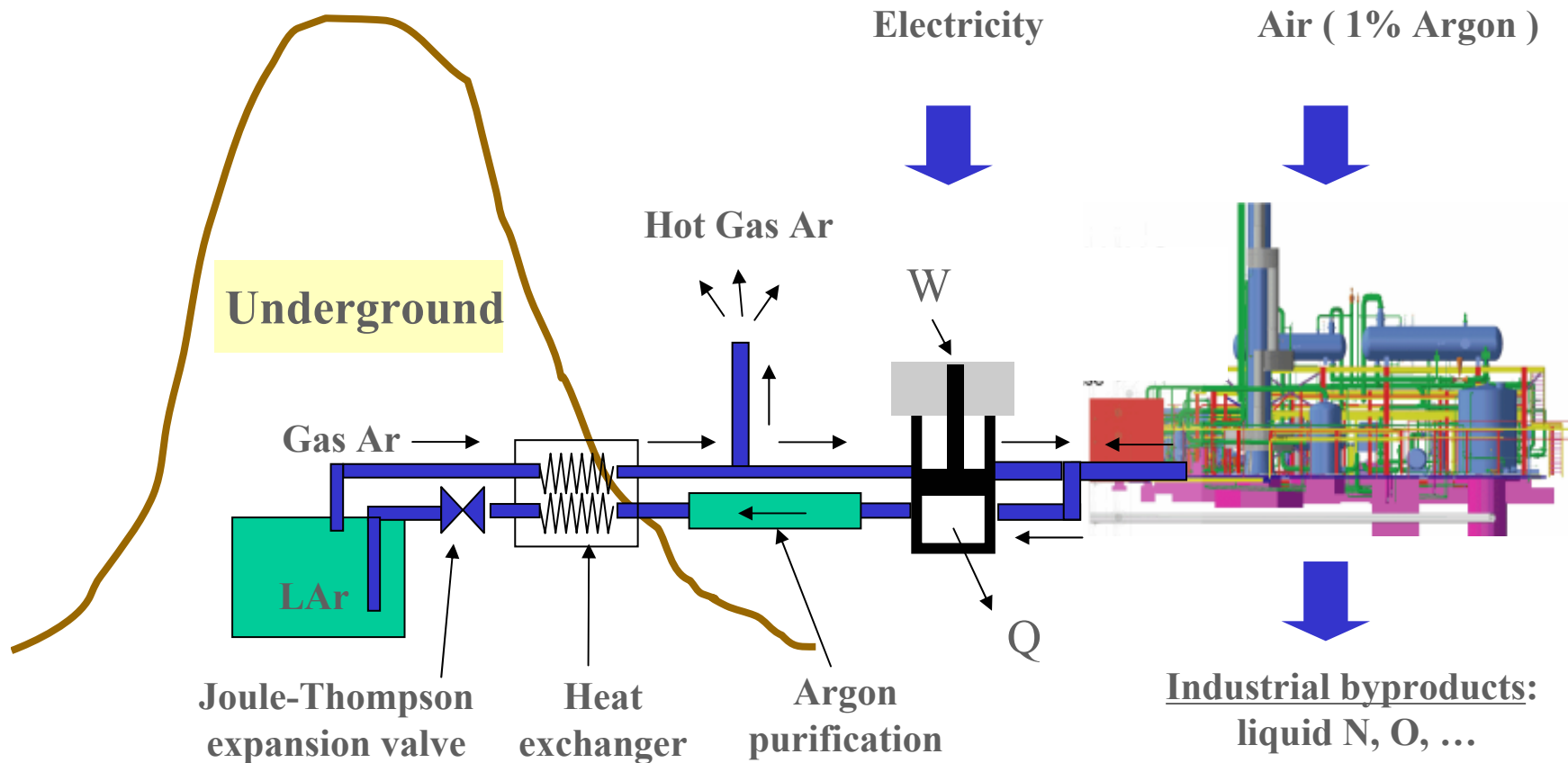
# A tentative detector layout



## Liquid Ar at boiling temperature

*If a gas exhaust is provided (feature not foreseen for ICARUS) the temperature remains constant  
Same technique is used for transportation and storage of Liquefied Natural Gas*

# A large cryogenic plant is needed



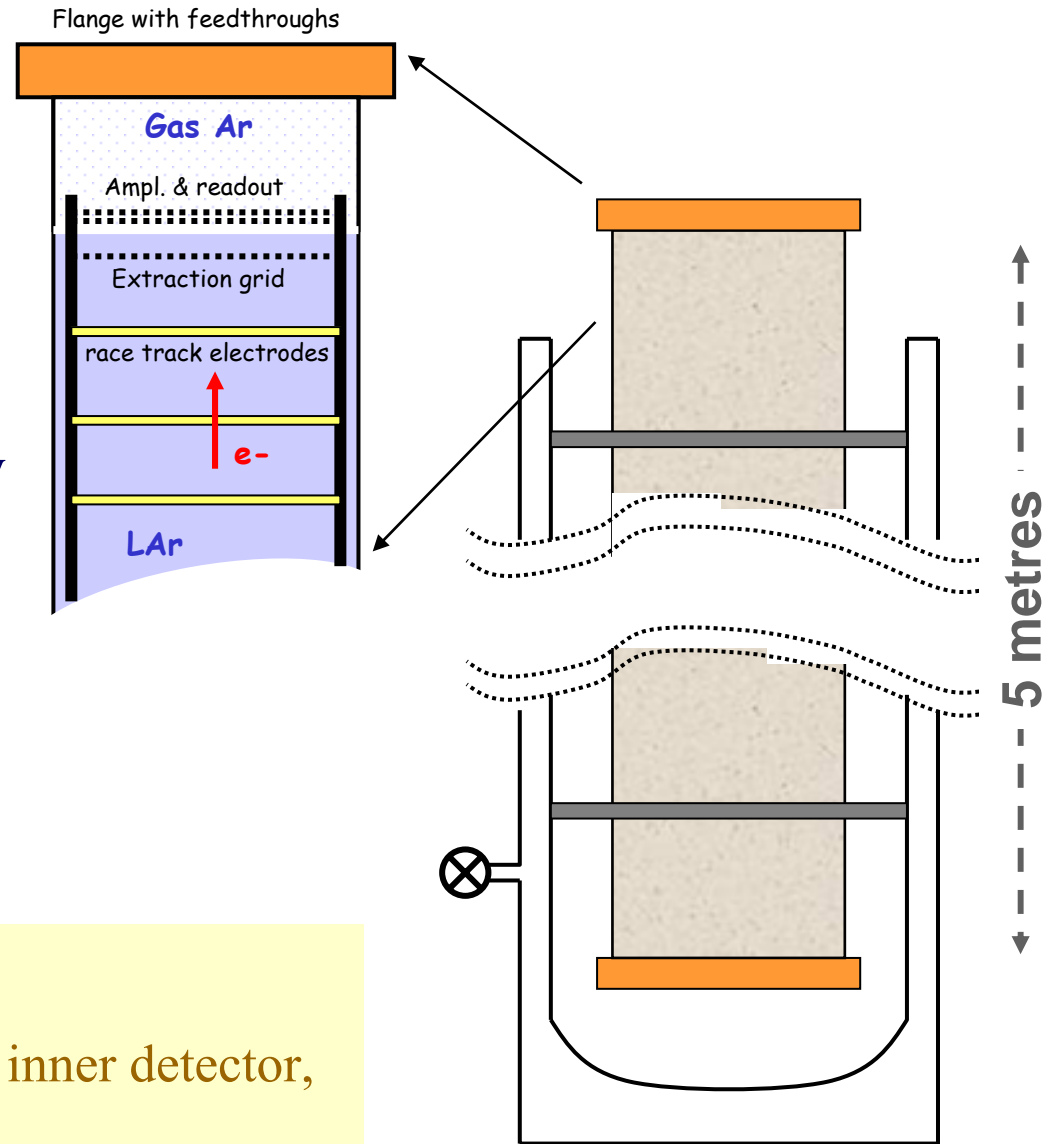
- Initial filling: transport LAr or in situ cryogenic plant  
Filling speed 150 ton/day → 2 years to fill
- 5 W/m<sup>2</sup> heat input: 30 ton/day refilling needed
- Continuous re-circulation (purity)

# Ongoing studies and initial R&D strategy

- Study of electron drift under high pressure ( $p \sim 3$  atm at tank bottom)
- Charge extraction, amplification and imaging devices
- Cavern design: large volume with aspect ratio  $\sim 1 : 1$
- Cryostat design, in collaboration with industry
- Logistics, infrastructure and safety issues for underground sites
- Study of LAr TPC prototypes in a magnetic field (for  $\nu$  Factory)
- Realization and test of a column-like detector prototype:  
5 m long drift and double-phase readout

# Test module for long drift, extraction, amplification

- Measurement of 5 m long drift, signal attenuation and multiplication
- Simulate 'very long' drift (20 m) by reduced E field and LAr purity

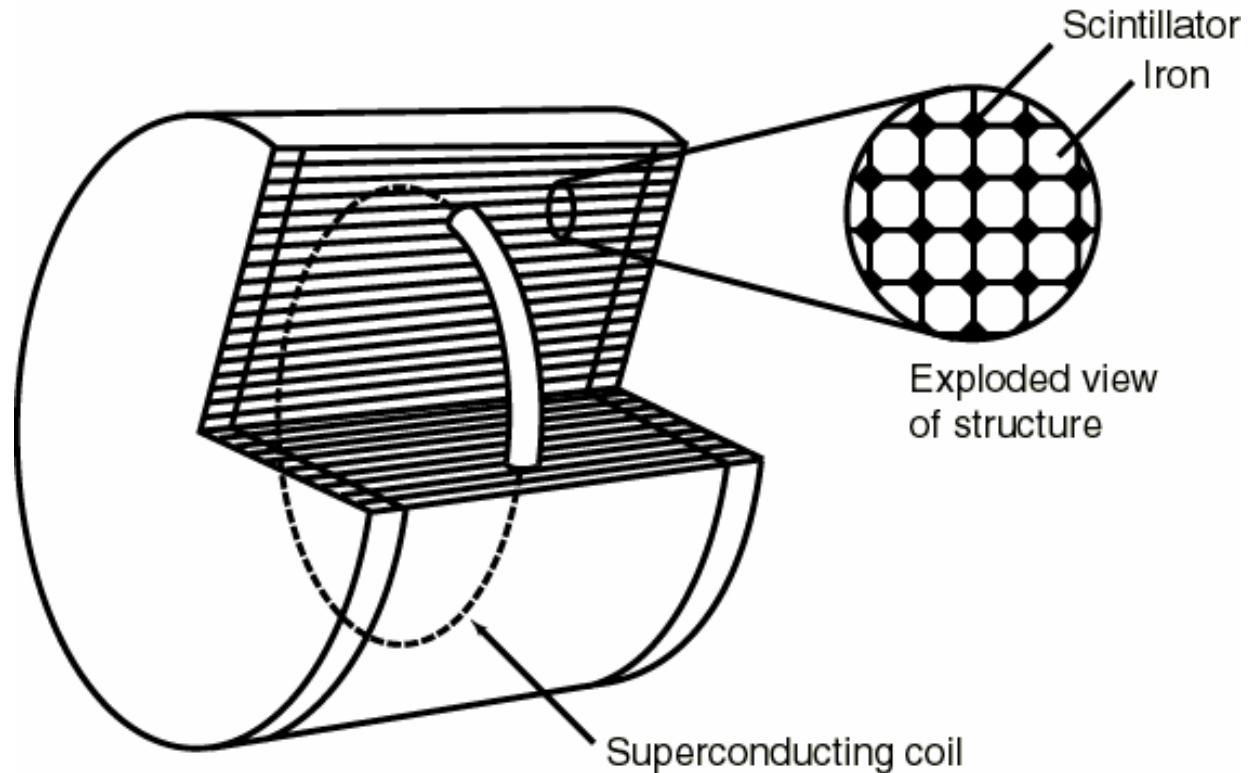


## Design in progress:

external dewar, detector container, inner detector, readout system, ...

# A Magnetised Iron Detector for a $\nu$ Factory

- Iron calorimeter, with scintillator rods as active detector
- Magnetised at  $B = 1$  T  
→ “wrong sign” muons from  $\nu_e$ - $\nu_\mu$  oscillations or  $\bar{\nu}_\mu$  CC?
- Conventional technique, but mass one order of magnitude  $>$  MINOS
- Only conceptual drawing available: practical problems (mechanics, magnet design.... ) must be addressed to assess the feasibility

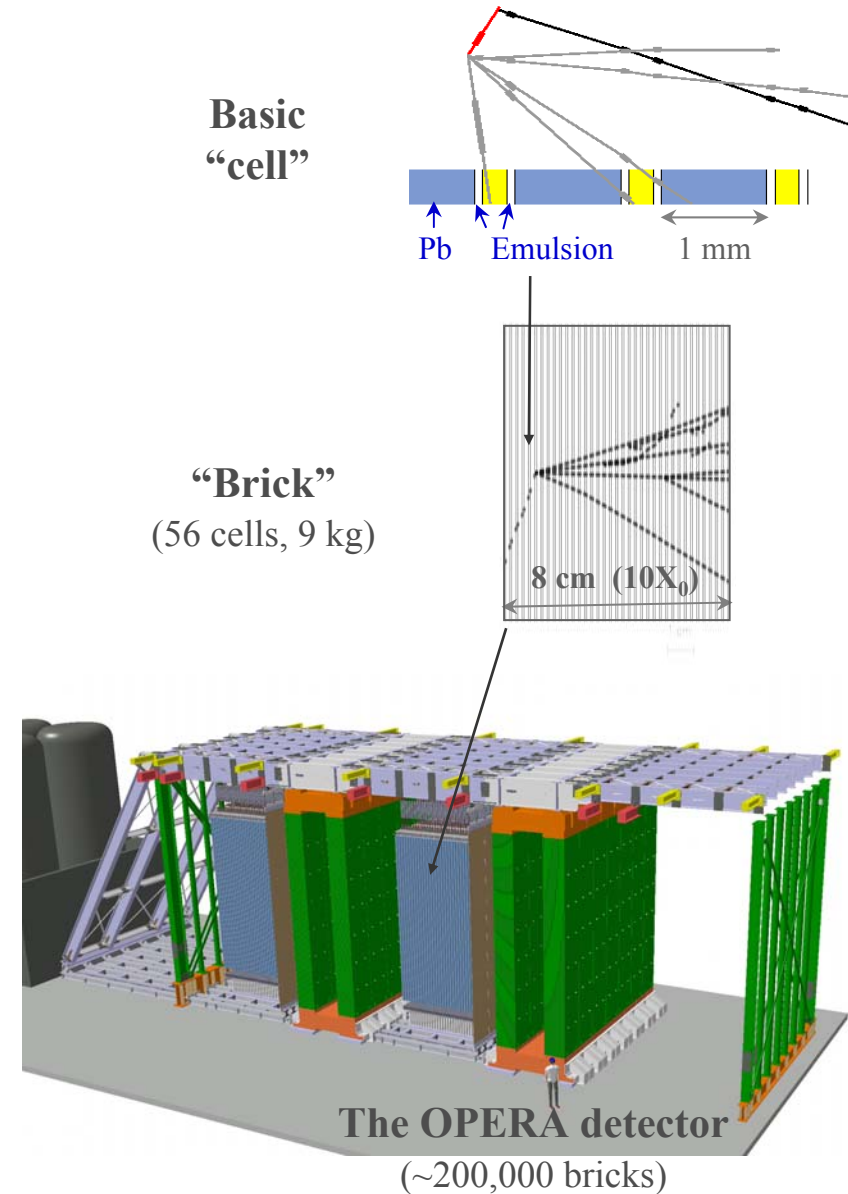


Dimension: radius 10 m, length 20 m  
Mass: 40 kt iron, 500 t scintillator

**Another possible design:  
a dipole magnet equipped with RPCs, à la MONOLITH**

# The Emulsion Cloud Chamber (ECC) for $\nu_e \rightarrow \nu_\tau$ appearance at $\nu$ Factories

- $\nu_e \rightarrow \nu_\mu$  (golden events) and  $\nu_e \rightarrow \nu_\tau$  (silver events) to reduce  $\theta_{13} - \delta$  ambiguities
- Pb as passive material, emulsion as sub- $\mu\text{m}$  precision tracker: unique to observe  $\tau$  production and decay
- 1.8 kton OPERA target mass  
→  $\sim 4$  kton at  $\nu$  Factory
- Search for  $\tau$ -decay only in events with a “wrong sign” muon: x 2 increase of scanning power required
- Hybrid experiment: emulsion + electronic detectors
- OPERA as a “milestone” for the technique



# Summary (1)

## Low-Z Calorimeter

- $\nu_{\mu}$ - $\nu_e$  oscillations  $\rightarrow \theta_{13}$  in off-axis NuMI beam: NOvA
- Follow-up of a proven technique
- NOvA mass  $\sim 10$  x MINOS
- Main issue: improve performance and reduce cost of trackers
- Main new technologies with respect to MINOS
  - plastic  $\rightarrow$  liquid scintillator (sampling or totally active detector)
  - PMTs  $\rightarrow$  APDs

# Summary (2)

## Water Čerenkov

- $\nu$  oscillation ,  $\nu$  astrophysics, proton decay
- Proven and very successful technique, well known also in its limitations
- Hyper-K, UNO and Frejus detector: mass  $\sim 10$  x Super-K
- Main issues:
  - cost and production of PMTs
    - strong collaboration with industry
    - can one develop other photo-detectors with adequate long-term reliability?
  - design of a large cavern



# Summary (3)

## Liquid Argon TPC

- A beautiful detector for  $\nu$  oscillation ,  $\nu$  astrophysics, proton decay
- Broad energy range
- Tested at the scale of the 300 ton ICARUS module:  
to be extrapolated in mass by  $>$  two orders of magnitudes
- New features envisaged to reach 100 kton: longer or much longer drifts, double-phase amplification and readout
- Very substantial R&D required on various aspects, depending on the design features and on the (underground) location: signal propagation and readout; electric field shaping; cryogenics, civil engineering, safety and logistics issues; .....

## Summary (4)

### Magnetised Fe Sampling Calorimeter

- “wrong sign”  $\mu$  from  $\nu_e \rightarrow \nu_\mu$  oscillations at  $\nu$ -Factories
- Proven technique, at smaller scale
- Mass  $\sim 10$  x MINOS
- Only conceptual drawing available

### Emulsion Cloud Chamber

- $\nu_e \rightarrow \nu_\tau$  at  $\nu$ -Factories, to complement  $\nu_e \rightarrow \nu_\mu$  oscillation in order to resolve  $\theta_{13}$  -  $\delta$  ambiguities
- Technique used in DONUT and being implemented in OPERA
- Mass  $\sim 2$  x OPERA
- OPERA as “milestone” for the technique

# General conclusions

**A physics program to be planned over decades, aiming at discoveries and precision measurements on neutrino masses and mixing matrix as well as on neutrino astrophysics, nucleon decay, unexpected physics ....**

**Requires large and reliable detectors with appropriate performance and acceptable cost**

**Depending on physics, detector developments and funding:**

- **Initiate what is now possible and advisable, carry out the related R&D**
- **Lay down the bases of further future:**
  - **R&D must look far ahead**
  - **Stimulate new ideas**