



***Present and future of neutrino (beam) oscillation experiments***

*Heavy Quarks and Leptons Workshop  
Puerto Rico, June 2004*

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**(special thanks to André Rubbia)**

Where are we now

## What do we know about neutrino masses and mixing ?

- there exist 3 'light' neutrinos (LEP)
- masses from direct measurements are small (limits from tritium & cosmology)
- neutrino mix (oscillations) → they are massive; PMNS matrix (3 x 3)
- oscillation parameters: 2 large mixing angles  $\theta_{\text{sol}} \sim \theta_{12}$ ,  $\theta_{\text{atm}} \sim \theta_{23}$   
2 independent mass splittings:  $\Delta m^2_{\text{sol}} \sim \Delta m^2_{12}$   
(masses are small, indeed)  $\Delta m^2_{\text{atm}} \sim \Delta m^2_{23}$

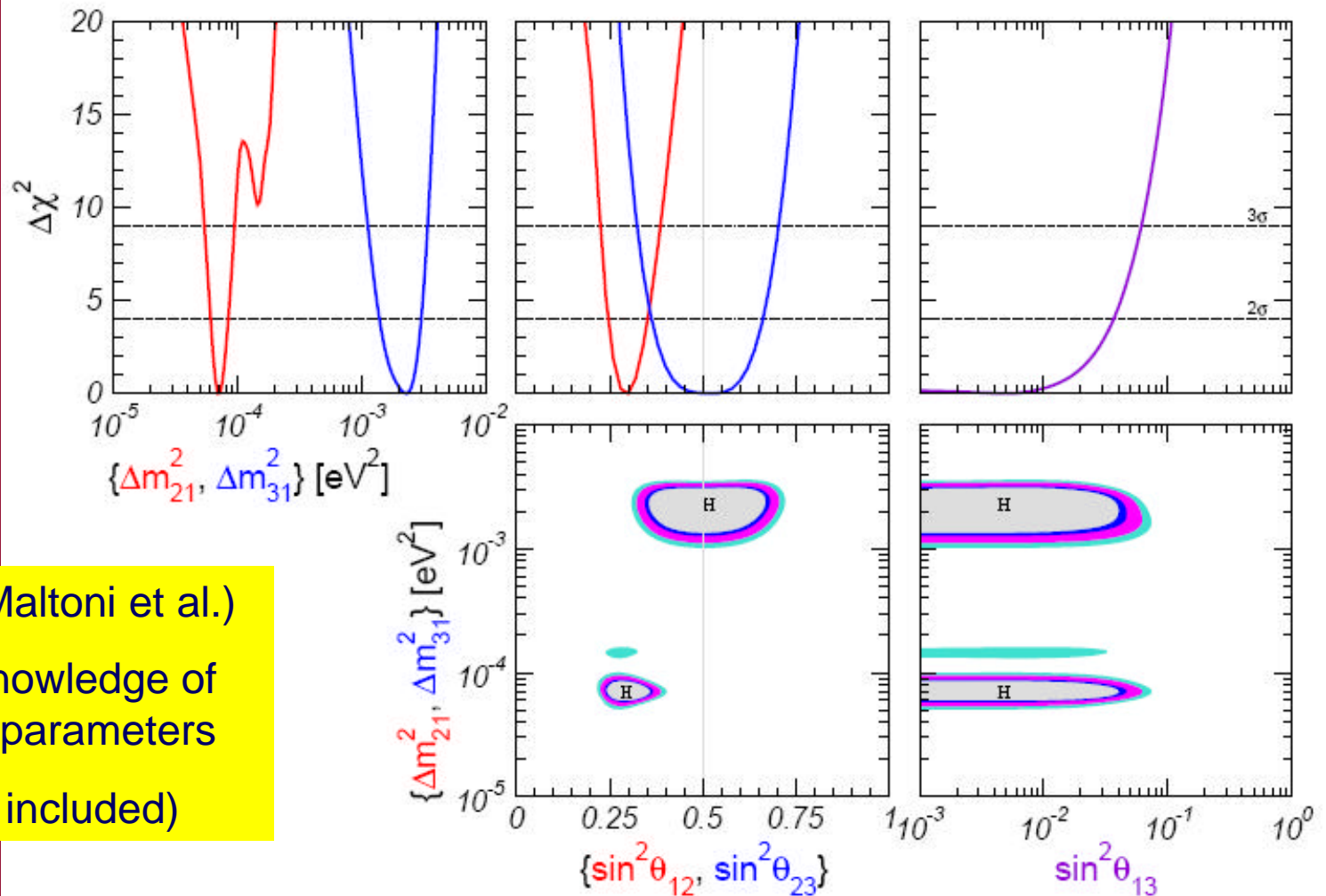
## What we do not know...

- absolute mass values (why are they small ?)
- why  $\theta_{12}$  and  $\theta_{32}$  angles are large and  $\theta_{13}$  seems very small or null ?
- Is mass hierarchy the same as for charged leptons (sign of  $\Delta m^2_{23}$ )
- Is there any CP violating phase in the mixing matrix ?

IMPORTANT NOTE: in all this I assume that there is no LSND effect !

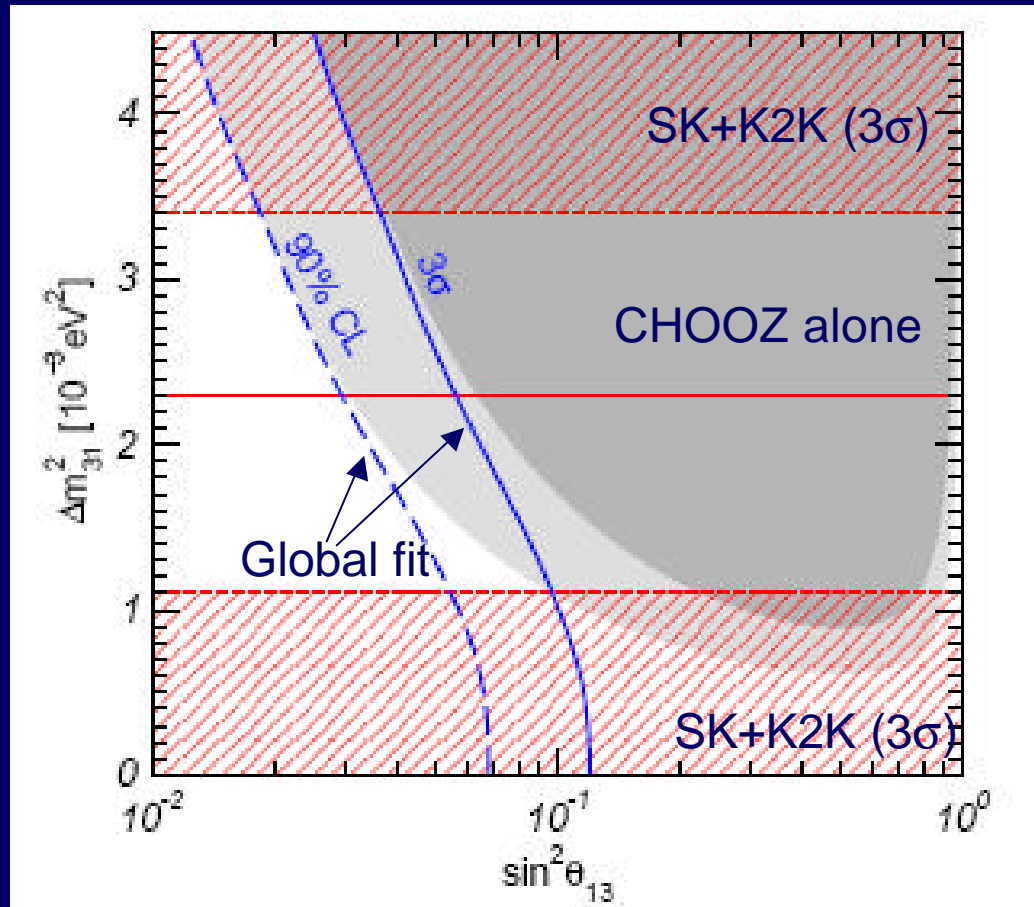
Wait for MiniBoone...

Global fit (Maltoni et al.)  
 our best knowledge of  
 oscillation parameters  
 (all data included)



parameter	best fit	2σ	3σ	5σ
$\Delta m_{21}^2 [10^{-5} \text{eV}^2]$	6.9	6.1–8.4	5.4–9.4	2.1–29
$\Delta m_{31}^2 [10^{-3} \text{eV}^2]$	2.3	1.4–3.0	1.1–3.4	0.68–4.4
$\sin^2 \theta_{12}$	0.30	0.25–0.35	0.23–0.39	0.16–0.47
$\sin^2 \theta_{23}$	0.52	0.36–0.66	0.32–0.70	0.26–0.78
$\sin^2 \theta_{13}$	0.005	$\leq 0.037$	$\leq 0.061$	$\leq 0.13$

In particular, estimate of  $\theta_{13}$  upper bound



Global fit (Maltoni et al.)  
CHOOZ, solar, Kamland  
+ atmospheric and K2K

## Objective of planned and future neutrino beam experiments:

- accurately measure the two  $\Delta m^2$ ,  $\theta_{12}$  and  $\theta_{23}$
- find the value of  $\theta_{13}$  from  $P(n_\mu - n_e)$
- show CP violating effects (without matter effects)
- show matter effects (without CP violation effects)  $\rightarrow$  hierarchy

# Neutrino mixing matrix and general 3 neutrino oscillation probability

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} e^{i\alpha_1/2} \nu_1 \\ e^{i\alpha_2/2} \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\begin{aligned} P(\nu_\ell \rightarrow \nu_{\ell'}) &= \left| \sum_i U_{\ell i} U_{\ell' i}^* e^{-i(m_i^2/2E)L} \right|^2 \\ &= \sum_i |U_{\ell i} U_{\ell' i}^*|^2 + \Re \sum_i \sum_{j \neq i} U_{\ell i} U_{\ell' i}^* U_{\ell j}^* U_{\ell' j} e^{i \frac{|m_i^2 - m_j^2|L}{2E}} \end{aligned}$$

For the important case of  $P(\nu_\mu \rightarrow \nu_e)$  oscillations, we have... 

$$P(\nu_\mu \rightarrow \nu_e) = \sum_{i=1,4} P_i$$

$$P_1 = \sin^2 \theta_{23} \sin^2 2\theta_{13} \left( \frac{\Delta_{13}}{B_\pm} \right)^2 \sin^2 \frac{B_\pm L}{2}$$

atmospheric part

$$P_2 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \left( \frac{\Delta_{12}}{A} \right)^2 \sin^2 \frac{AL}{2}$$

solar part

$$P_3 = J \cos \delta \left( \frac{\Delta_{12}}{A} \right) \left( \frac{\Delta_{13}}{B_\pm} \right) \cos \frac{\Delta_{13}L}{2} \sin \frac{AL}{2} \sin \frac{B_\pm L}{2}$$

interference

$$P_4 = \mp J \sin \delta \left( \frac{\Delta_{12}}{A} \right) \left( \frac{\Delta_{13}}{B_\pm} \right) \sin \frac{\Delta_{13}L}{2} \sin \frac{AL}{2} \sin \frac{B_\pm L}{2}$$

interference

where

$$\Delta_{ij} = \frac{\Delta m_{ij}^2}{2E_\nu}$$

$$A = \sqrt{2} G_F n_e$$

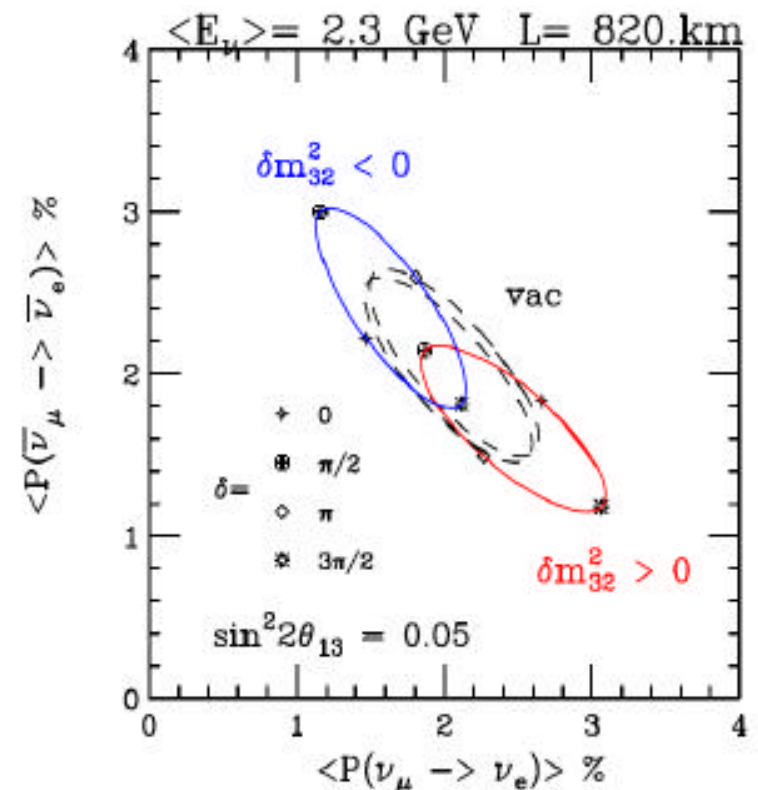
$$B_\pm = |A \pm \Delta_{13}|$$

$$J = \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$$

and the  $\pm$  signifies neutrinos or antineutrinos

In vacuum, at leading order:

$$P(\nu_\mu \rightarrow \nu_e) \propto \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \frac{\Delta m_{23}^2 L}{4E}$$





# Measuring CP violating effects

Best method:

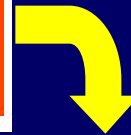
$$A_{CP} = \frac{P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu) - P(\nu_e \rightarrow \nu_\mu)}{P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu) + P(\nu_e \rightarrow \nu_\mu)} \simeq \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \cdot \sin \delta \cdot \sin \frac{\Delta m_{12}^2 L}{4E}$$

it requires:  $\Delta m_{12}^2$  and  $\sin 2\theta_{12}$  large (LMA solar): OK !

larger effects for long L: matter and 2<sup>nd</sup> osc.max

however...

$$P(\nu_\mu \rightarrow \nu_e) \propto \sin^2 2\theta_{13} \quad A_{CP} \propto \frac{1}{\sin \theta_{13}}$$



$\sin^2 2\theta_{13}$  small: low statistics and large asymmetry

$\sin^2 2\theta_{13}$  large: high statistics and small asymmetry

impact on the detector design

...and:

$$P(\nu_\mu \rightarrow \nu_e) \propto \sin^2 \frac{\Delta m_{23}^2 L}{4E}$$

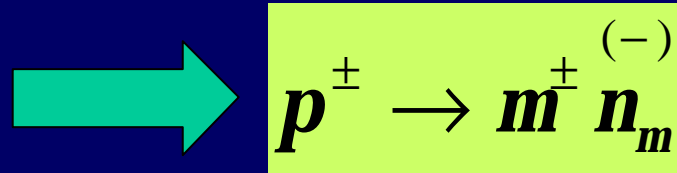
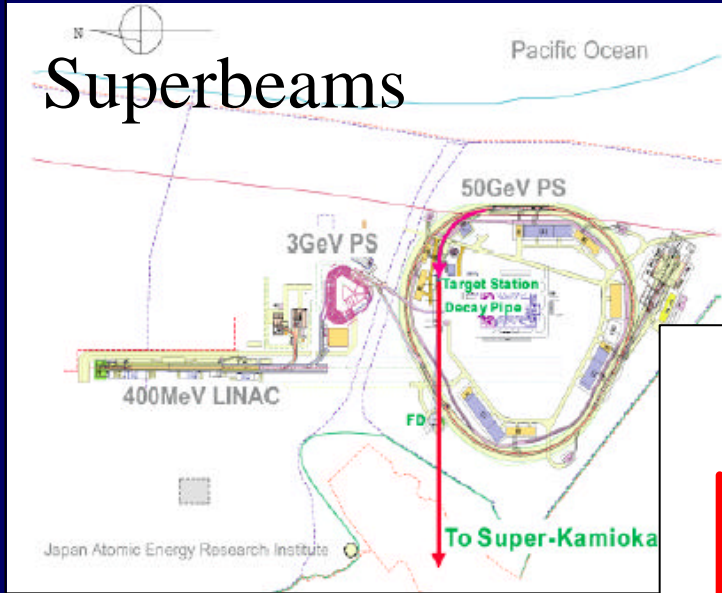


oscillations are governed by  $\Delta m_{atm}^2$ , L and E:

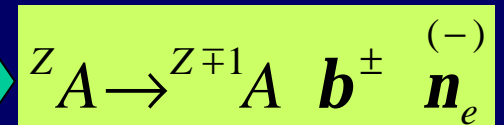
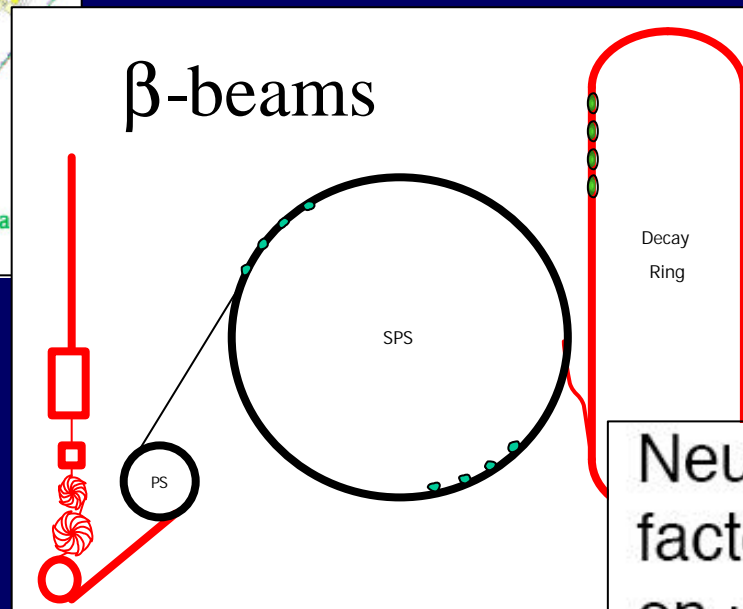
$E \gg 5 \text{ GeV} \rightarrow L \gg 3000 \text{ km}$

flux too low with a conventional LBL beam

# High intensity neutrinos facilities



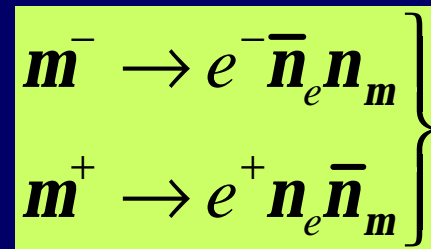
Select focusing sign



Select ion

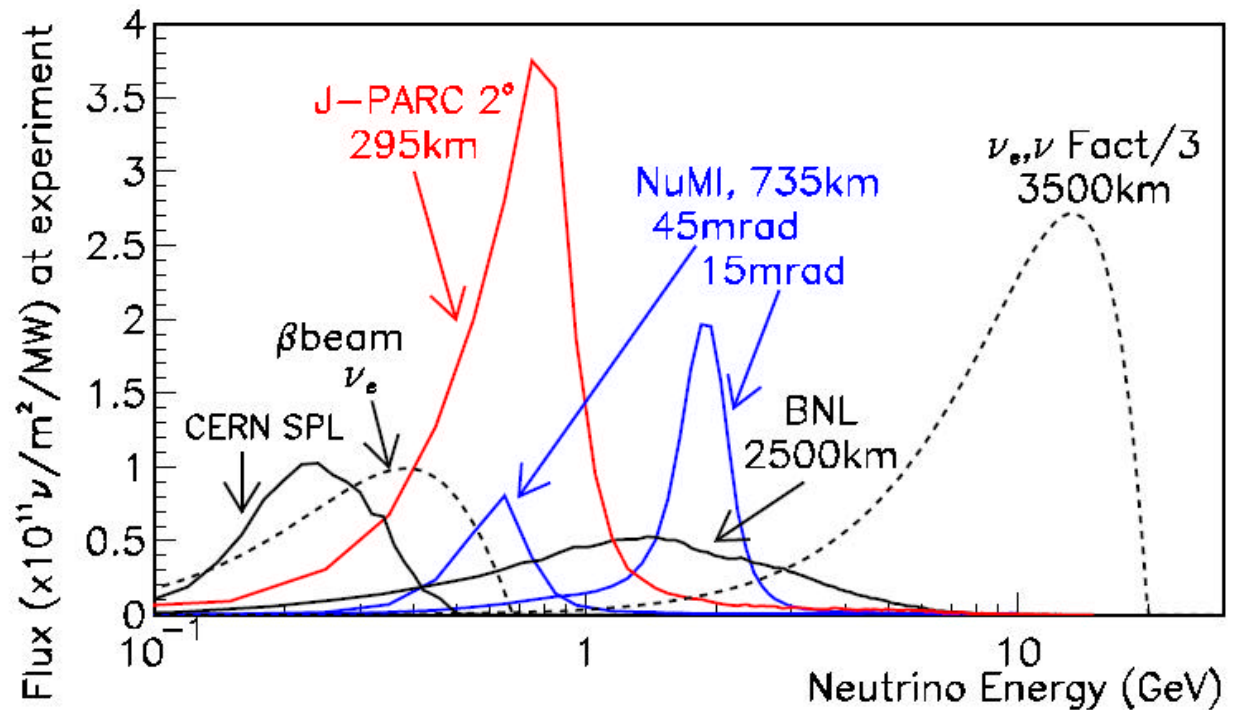


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Future neutrino beams

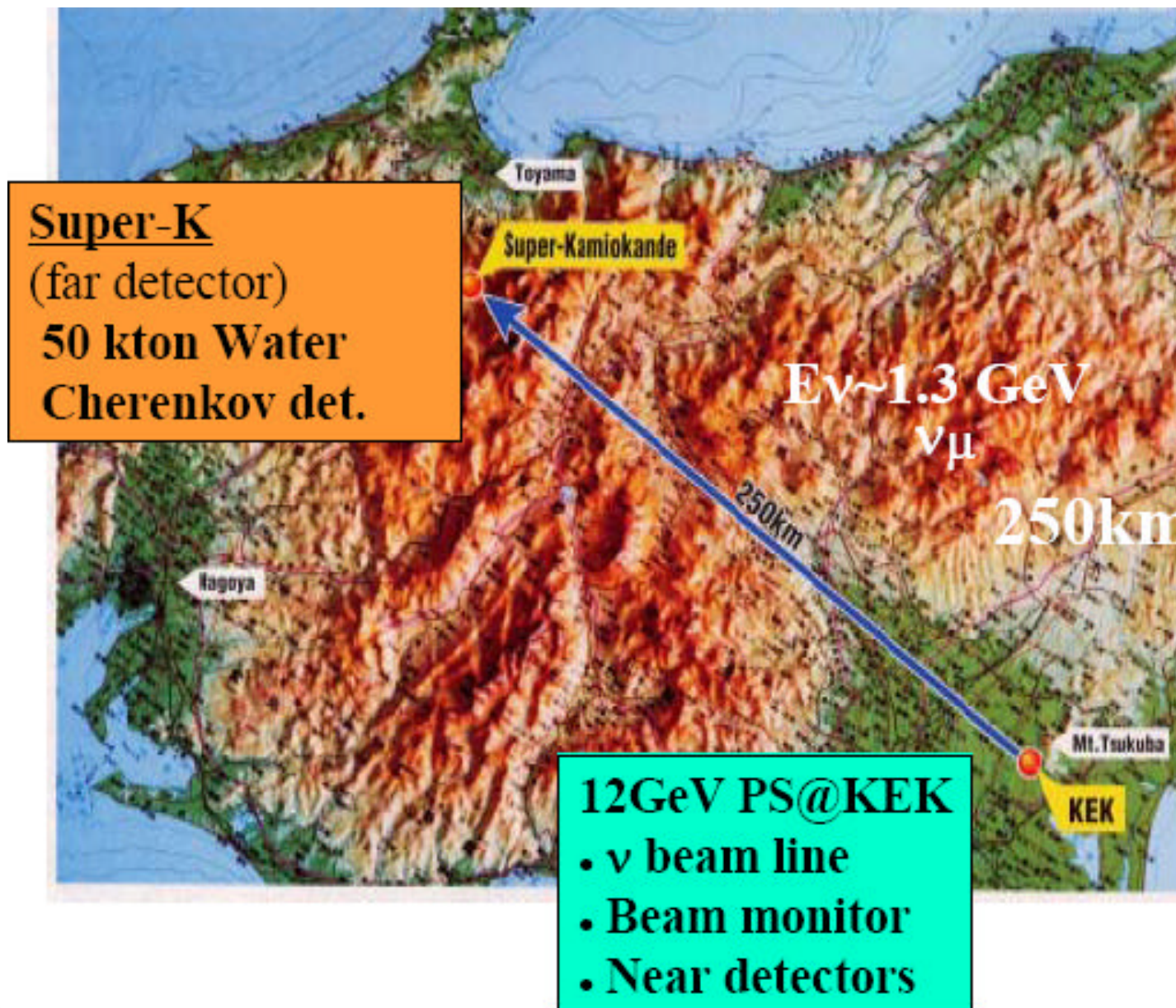
Outstanding goals



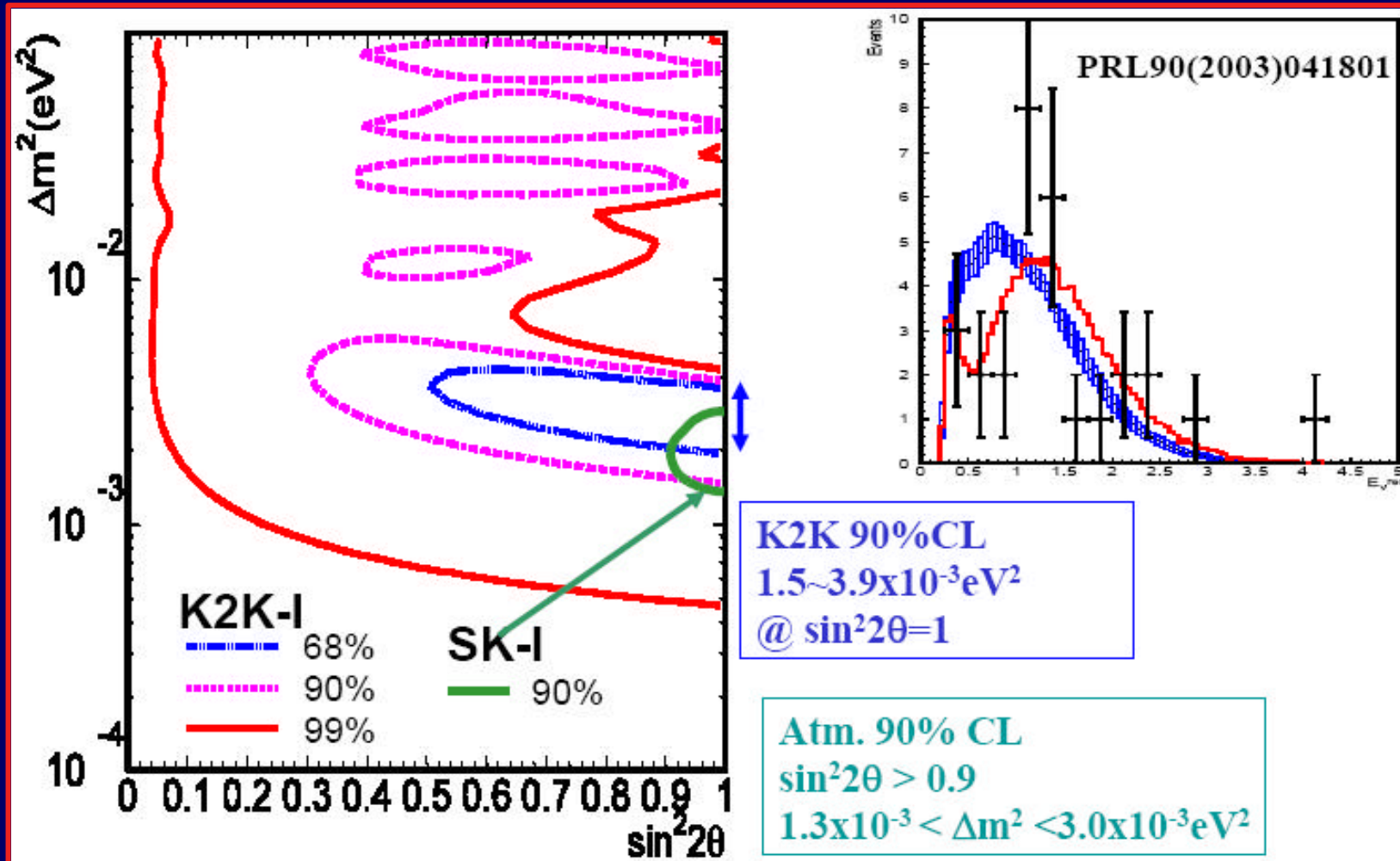
Physics	Value of $\sin^2 2\theta_{13}$			
	$> 4 \times 10^{-2}$	$> 1 \times 10^{-2}$	$> 10^{-3}$	$> 10^{-4}$
Seeing $\theta_{13} \neq 0$	MINOS CNGS	Conventional Superbeams Phase I	Conventional Superbeams Phase II	$\nu$ Factory $L \geq 3500$ km
Mass Hierarchy	Combinations of Phase I Superbeams	Combinations of Phase II Super/ $\beta$ -beams	Combinations of $\nu$ Factory and Super/ $\beta$ -beams	$\nu$ Factory $L \sim 7700$ km
Evidence for CP-violation	Combinations of Phase I Superbeams	Combinations of Phase II Super/ $\beta$ -beams	Combinations of $\nu$ Factory and Super/ $\beta$ -beams	Combinations of $\nu$ Factory 2 baselines

Back from the future: work in progress...

## K2K: the mother of all LBL experiments

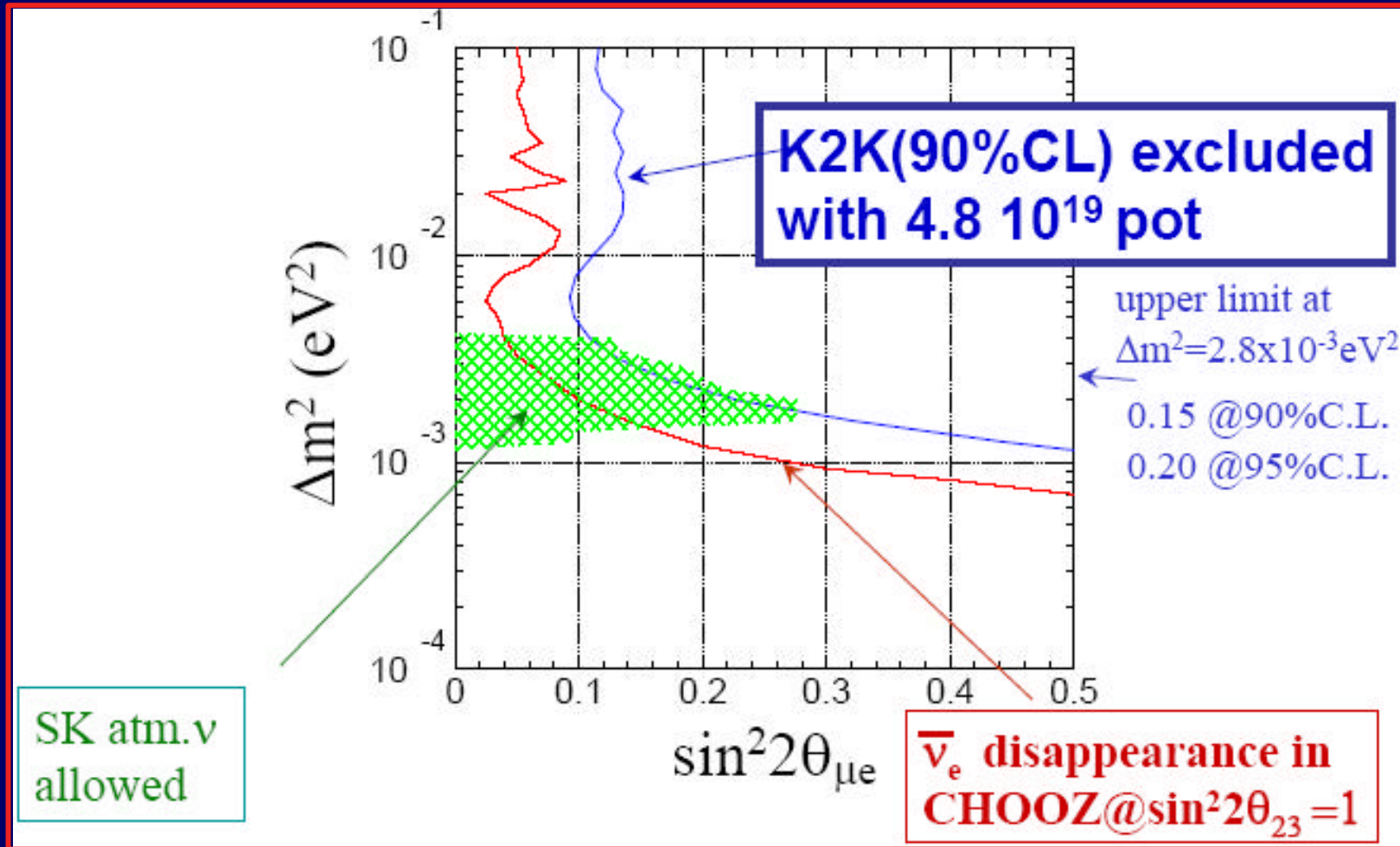


# SK plus K2K



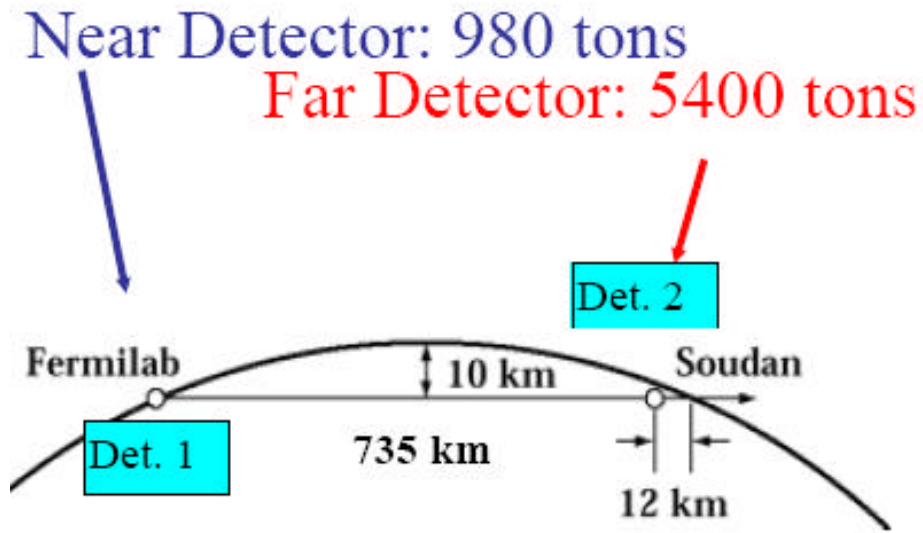
The two experiments agree well: full mixing and  $\Delta m^2$  in the range of 1-3  $\text{eV}^2$

## K2K looking for electron appearance



# Next to come on duty: MINOS in the NuMi neutrino beam

Start 2005

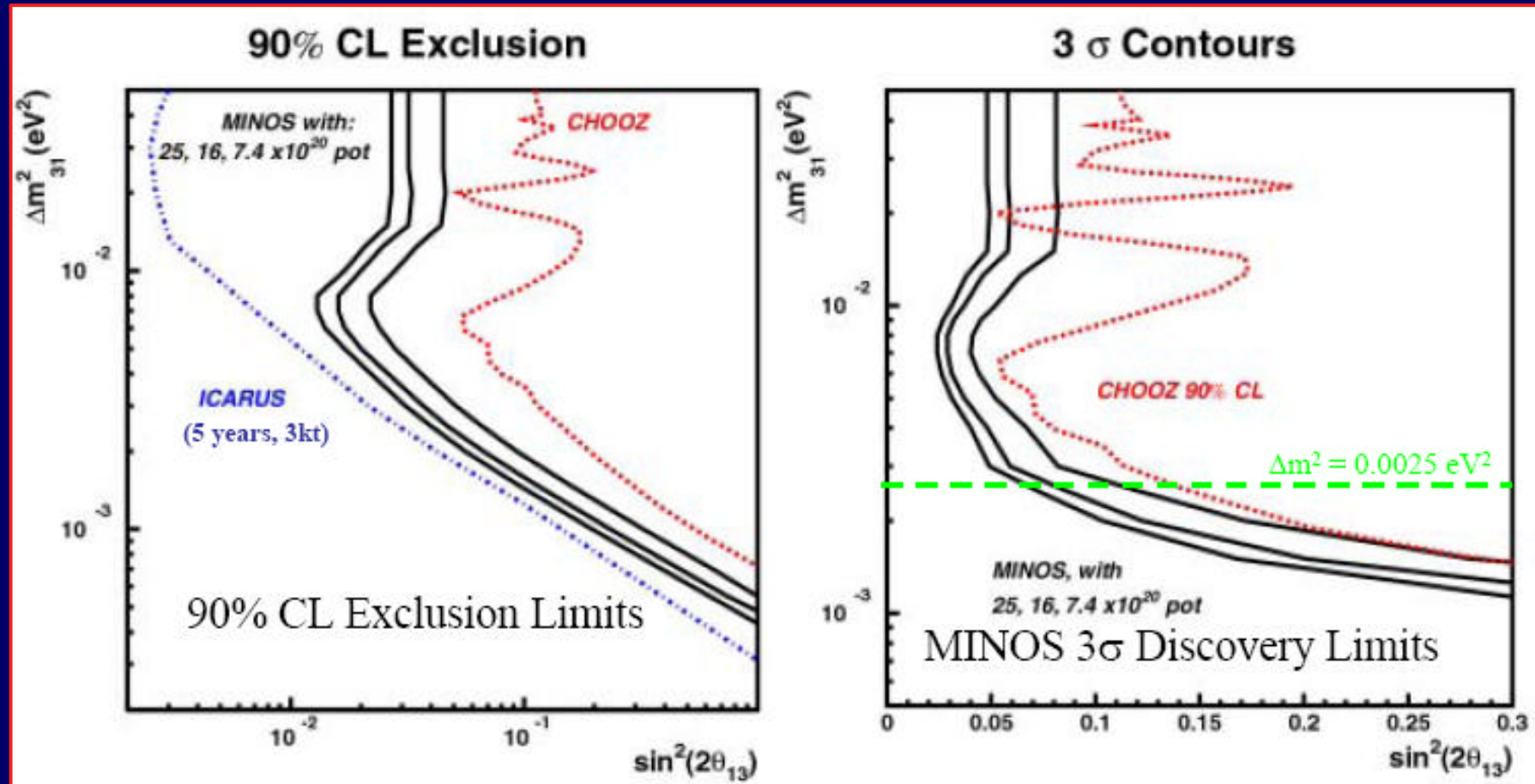


Magnetized steel/scintillator calorimeter

- low E neutrinos (few GeV):  $\nu_\mu$  disappearance experiment
- $4 \times 10^{20}$  pot/y  $\rightarrow$  2500  $\nu_\mu$  CC/year
- compare Det1-Det2 response vs E  $\rightarrow$  in 2-6 years sensitivity to  $\Delta m^2_{\text{atm}}$

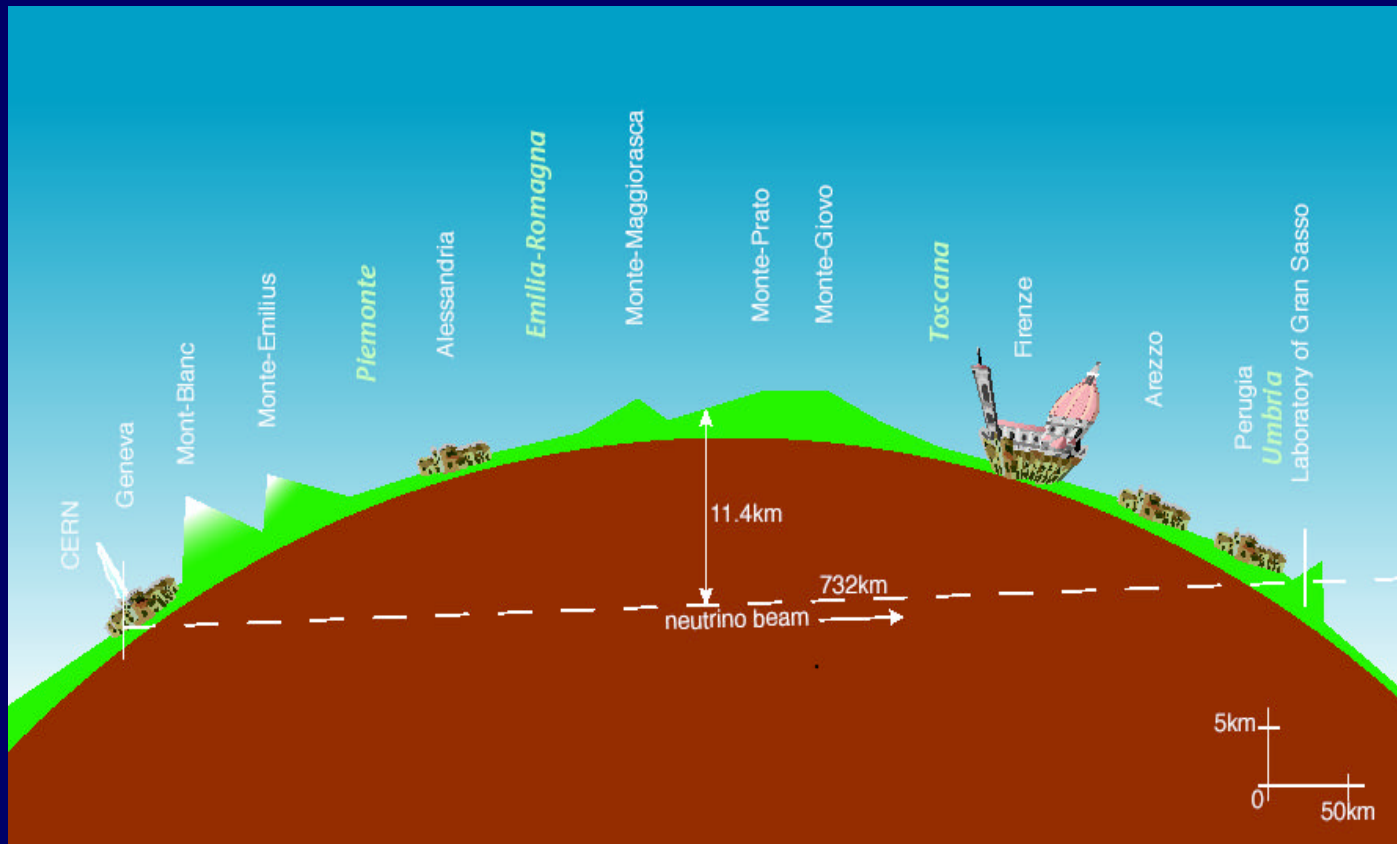


# electron appearance in MINOS



# $\tau$ appearance at LNGS in the CNGS beam

Start 2006



- High energy beam:  $\langle E \rangle$  of about **20 GeV**: tau appearance search
- **$4.5 \times 10^{19}$**  pot/year from the CNGS. In the hypothesis of no oscillation:
- **2600**  $n_m$  CC/year per kton detector mass
- Assuming  $n_m - n_t$  oscillation, with parameters  **$\sin^2 2q = 1$**  and  **$Dm^2 = 2.5 \times 10^{-3} \text{ eV}^2$** :  
**15  $n_t$**  CC/year per kton
- construction well advanced: in schedule. Two experiments: **OPERA** and **ICARUS**

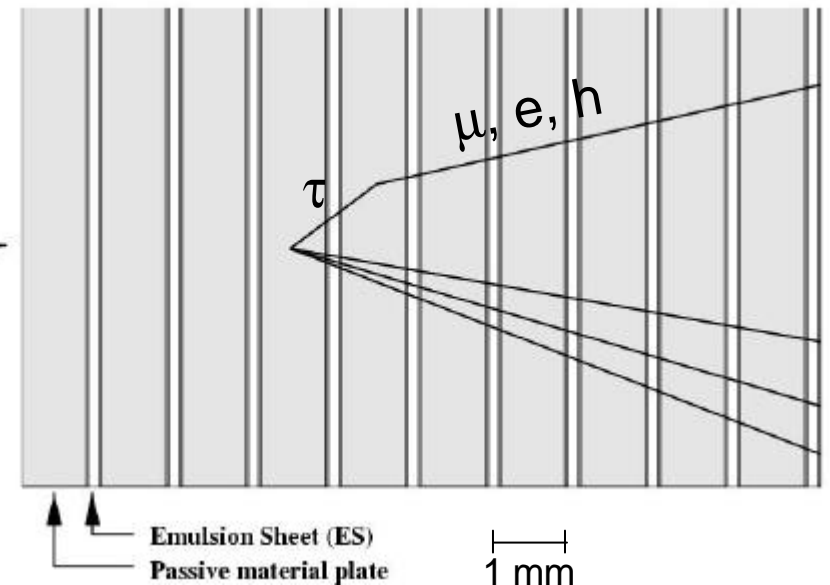
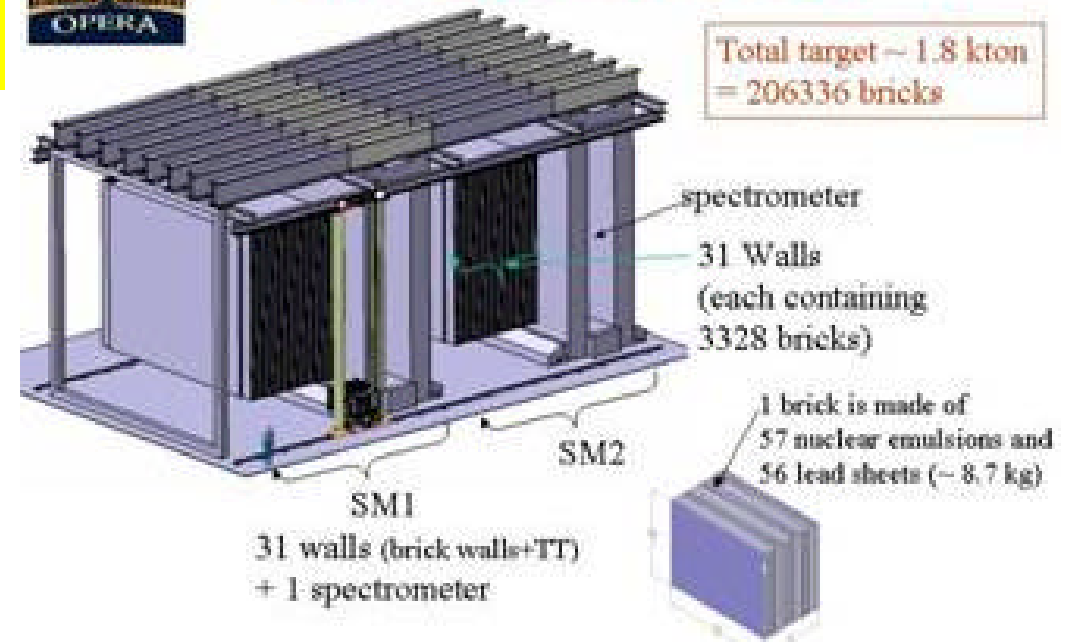
## The OPERA experiment at LNGS: the rebirth of the emulsion technique

- detector: 1800 ton emulsion/lead bricks (ECC technique) complemented by tracking scintillator planes and two muon spectrometers
- industrial emulsion production and handling
- need huge scanning power/speed: > tens of automatic microscope running in parallel @ 10 cm<sup>2</sup>/hour (advances of the technique)

- **Low BG experiment:** (<1 ev.) t track reconstruction
- **Low statistics:** about 10 events/5 years at nominal CNGS intensity @ SK parameter values: statistics goes like  $(Dm^2)^2$
- **Aim at beam intensity increase**
- **Installation in progress**



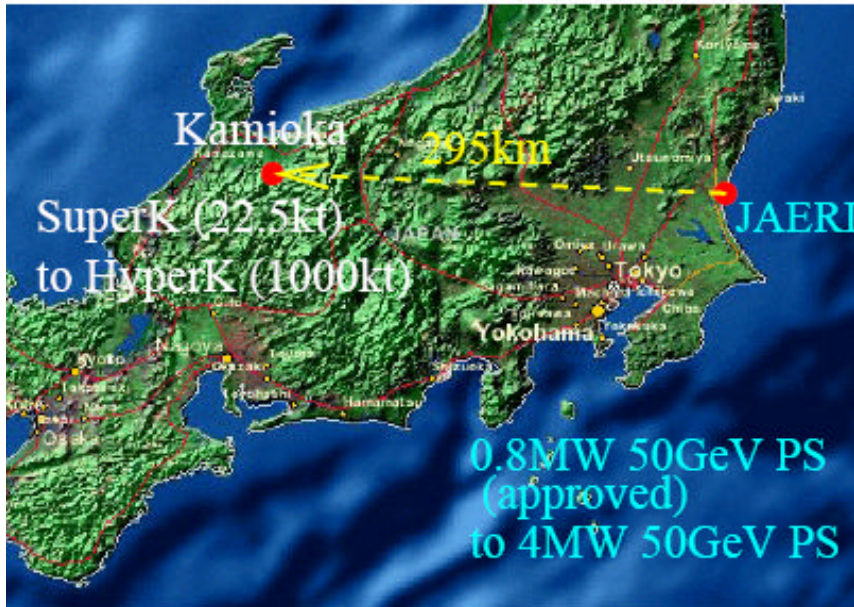
## The OPERA Detector



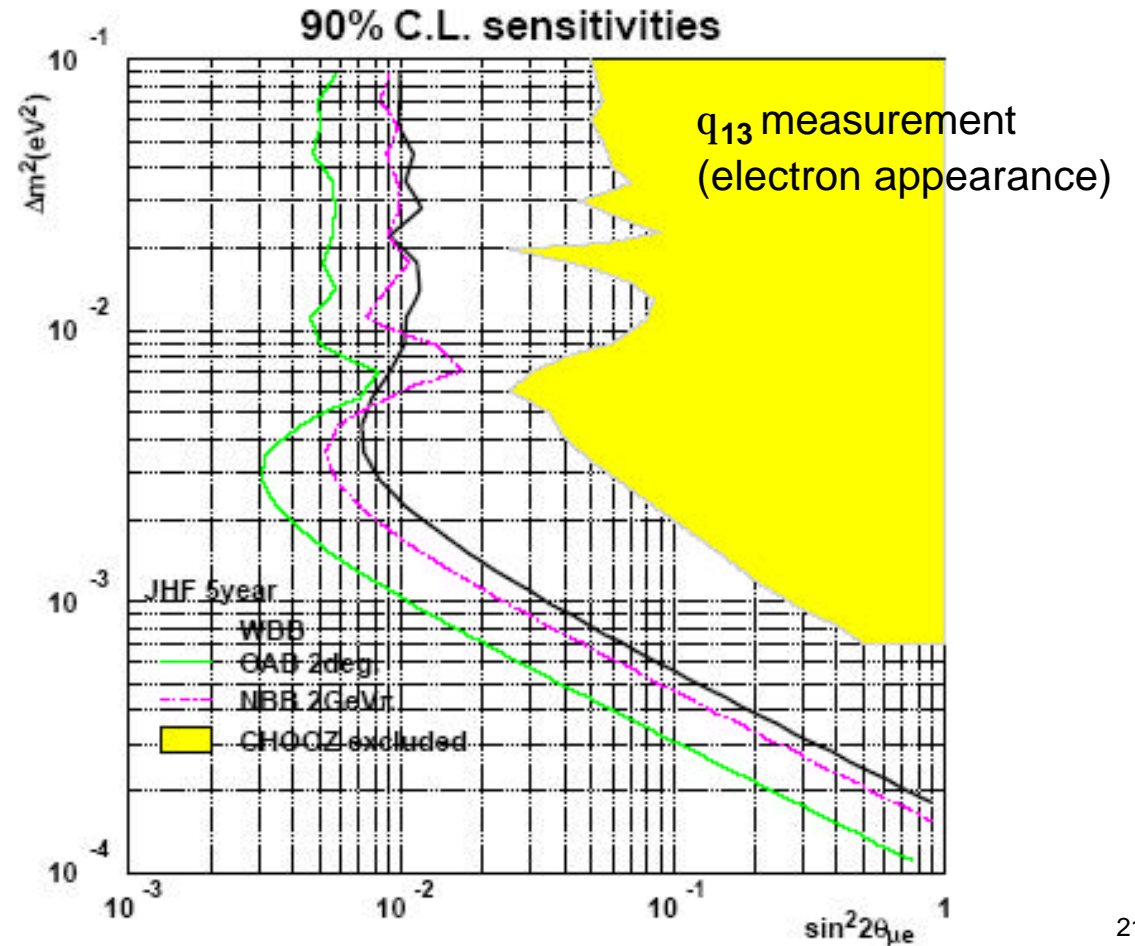
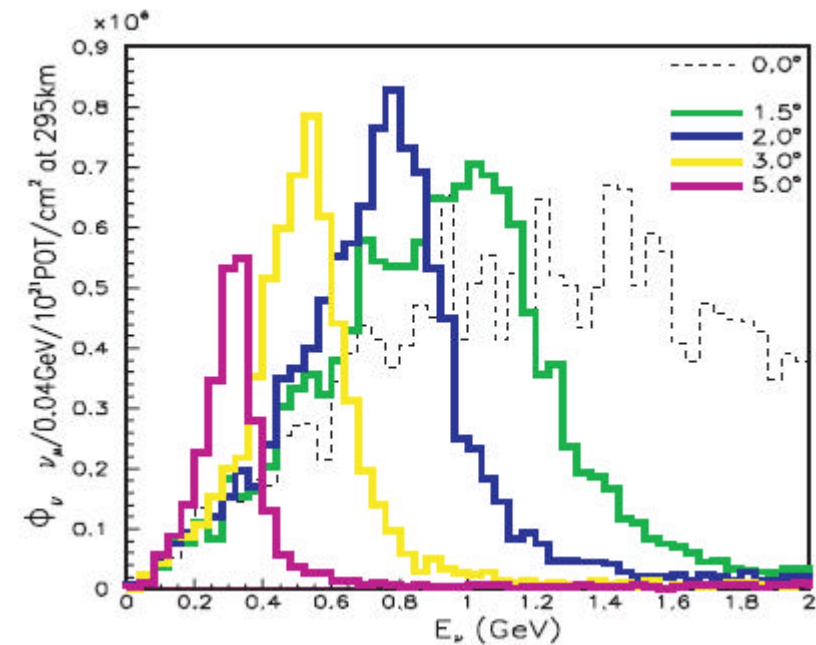
Measure  $\theta_{13}$

# Off-axis beam in Japan, another experiment with SK: T2K

Start 2009



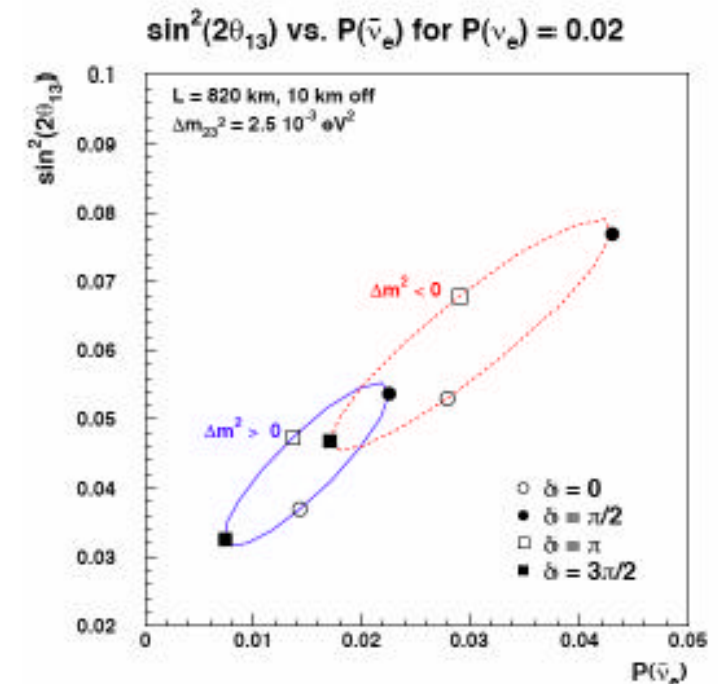
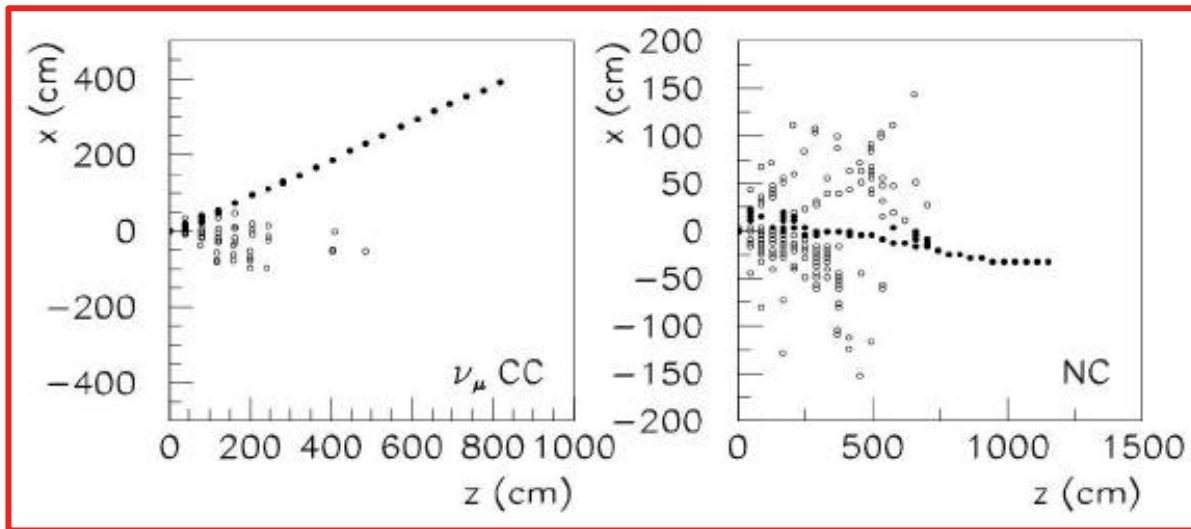
- low  $E_\nu$  (<1 GeV) Super-Beam:  $10^{21}$  pot/year
- @  $2^\circ \rightarrow 3000 \nu_\mu$  CC/year (x10 w.r.t. K2K)
- SK plus two near detectors (280 m and 2 km)



# An off-axis experiment in the NuMI beam: $\text{NonA}$

- recent proposal (March 04); nominal NuMI beam: 0.4 MW
- far detector: 50 kton @ Ash River (MN) 810 km from Fermilab (12 km, 14 mrad off-axis)
- technique: particleboard/liquid scintillator with fiber/APD R/O
- near detector: same as far, 1 ton fid. mass; also use MINERVA ?

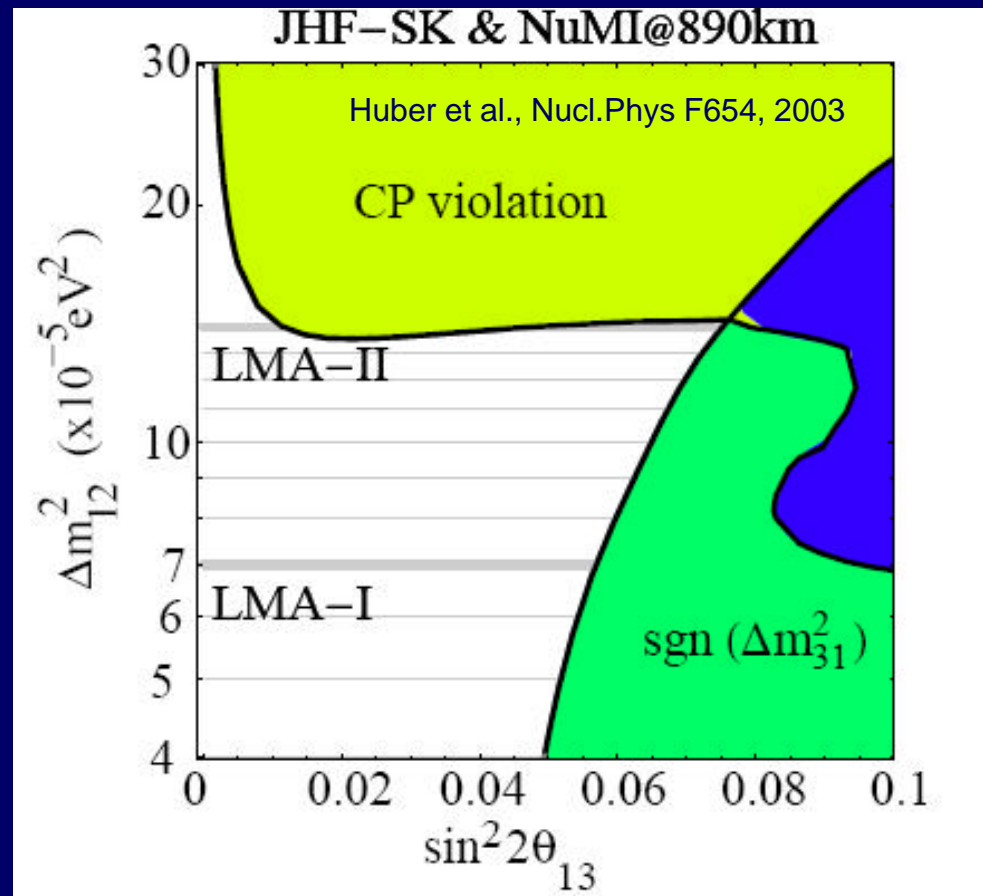
Start 2009 -2010



Conventional detector design: well known technique of low density fine grained calorimeters (CHARM II at CERN)

**cost of about \$150 M**

With some chance, next generation experiments on  $\theta_{13}$  could measure mass hierarchy and CP effects



# Pin down CP phase and mass hierarchy



# More distant future: Super-Beams, Beta-Beams, n-fact

Start 2015 -2020

Outstanding goal: detect CP violation (if  $q_{13}$  not zero!)

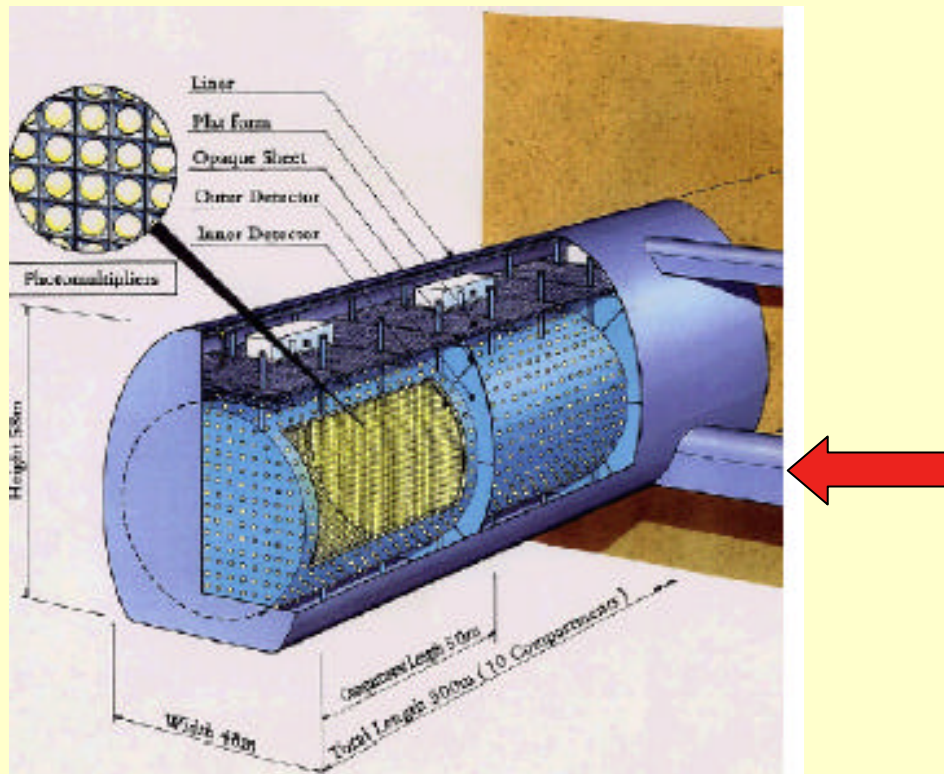
high intensity is a must; two approaches on **L/E** :

**long/high** (e.g. BNL-Fermilab projects):

- matter effects increase signal ( $E_{\max 1}/E_{\max 2}$ )
- CP effects increase with L ( $3\pi/2$  vs  $\pi/2$ )

**short/low** (e.g. CERN-SPL to Frejus):

- below threshold for BG (? ...Fermi motion)
- atmospheric neutrino BG
- antineutrino x-section small



**DETECTORS**

500-1000 kton Water Cerenkov 'a la SK' (Hyper-K, UNO) are considered as baseline

Rationale: exploit a well known technique  
aim at a 'reasonable' cost

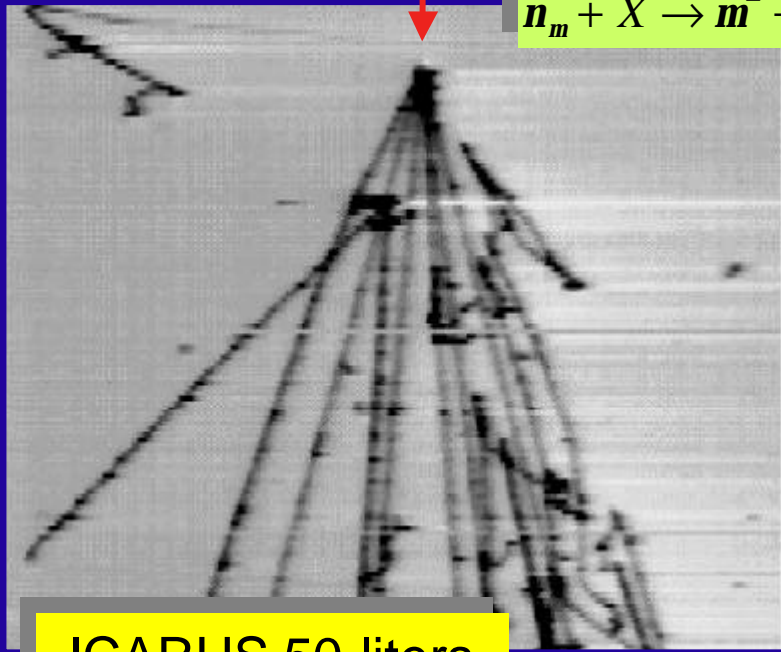
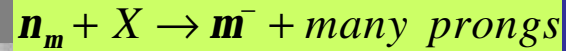
**HOWEVER....**

...liquid Argon instead of liquid water ?

See A.E. and A. Rubbia

Contribution to the Workshop on Physics with a Multi-MW Proton Source,  
CERN, 25-27 May 2004

# Real neutrino events observed by LAr TPC and water Cerenkov



K2K

## Super-Kamiokande

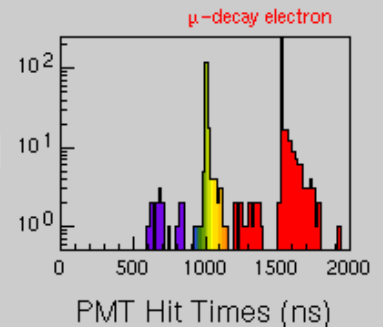
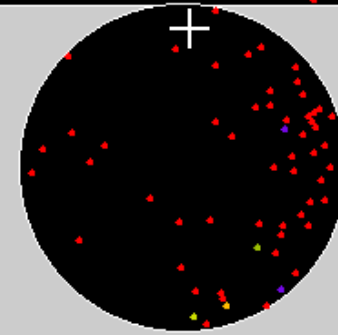
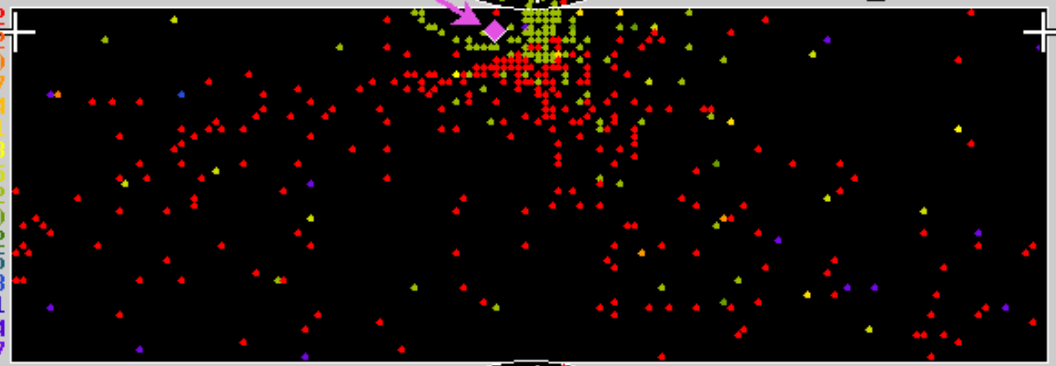
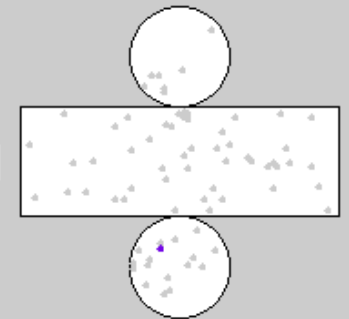
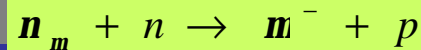
Run 7436 Event 1405412  
 99-06-19:18:42:4  
 Inner: 516 hits, 1018 pE  
 Outer: 2 hits, 2 pE (in-time)  
 Trigger ID: 0x0  
 D wall: 240.4cm

Neutrino Beam  
 Direction  
 from KEK

Resid(ns)

- > 182
- 160- 182
- 137- 160
- 114- 137
- 91- 114
- 68- 91
- 45- 68
- 22- 45
- 0- 22
- -22- 0
- -45- -22
- -68- -45
- -91- -68
- -114- -91
- -137--114
- <-137

FIRST K2K EVENT  
 RECORDED BY SUPER-K

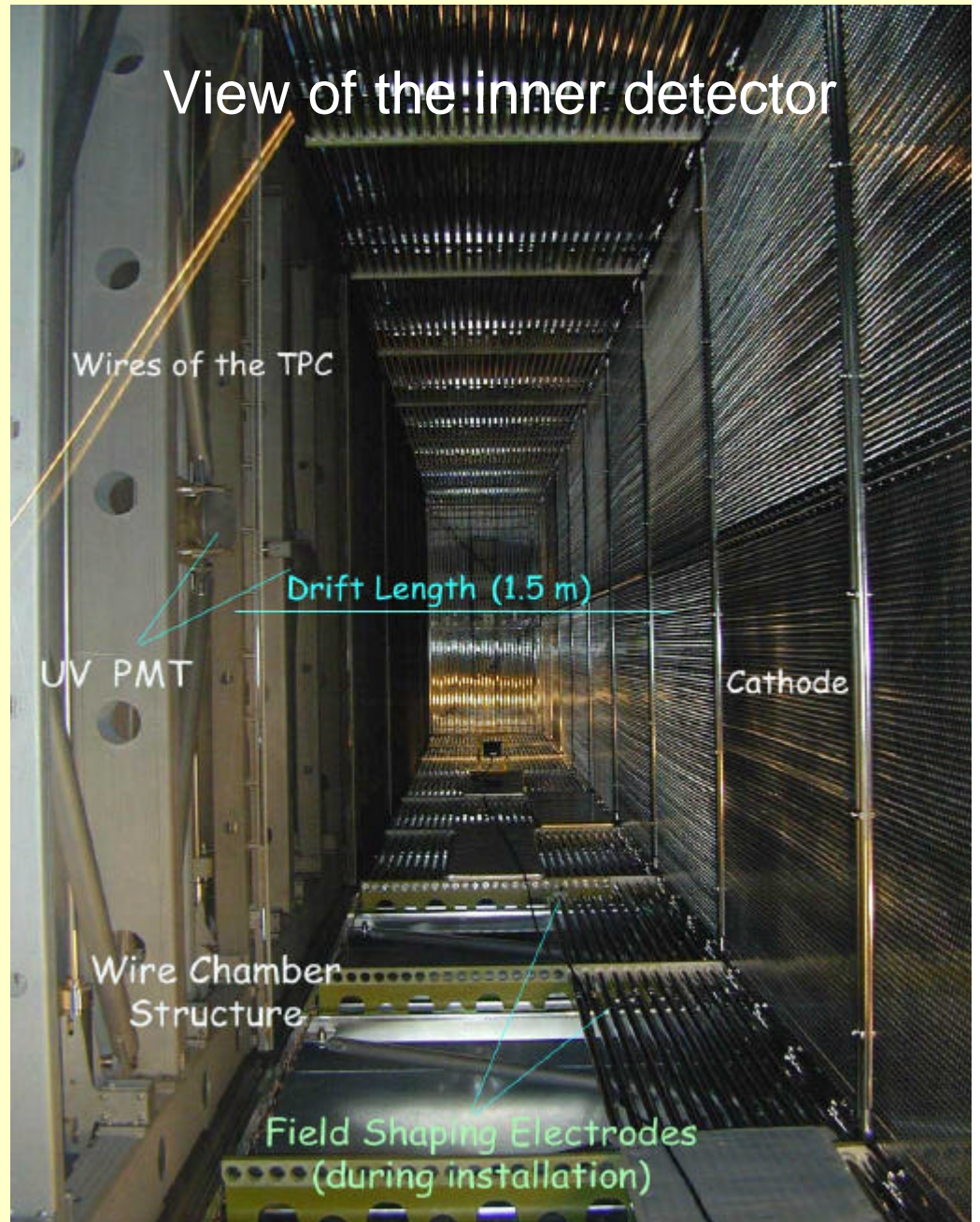
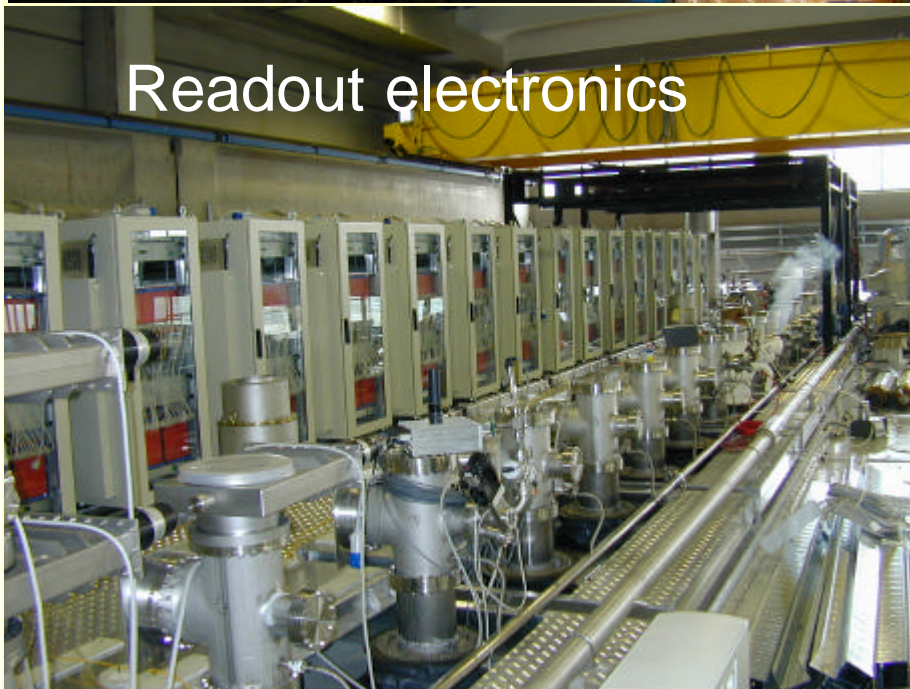
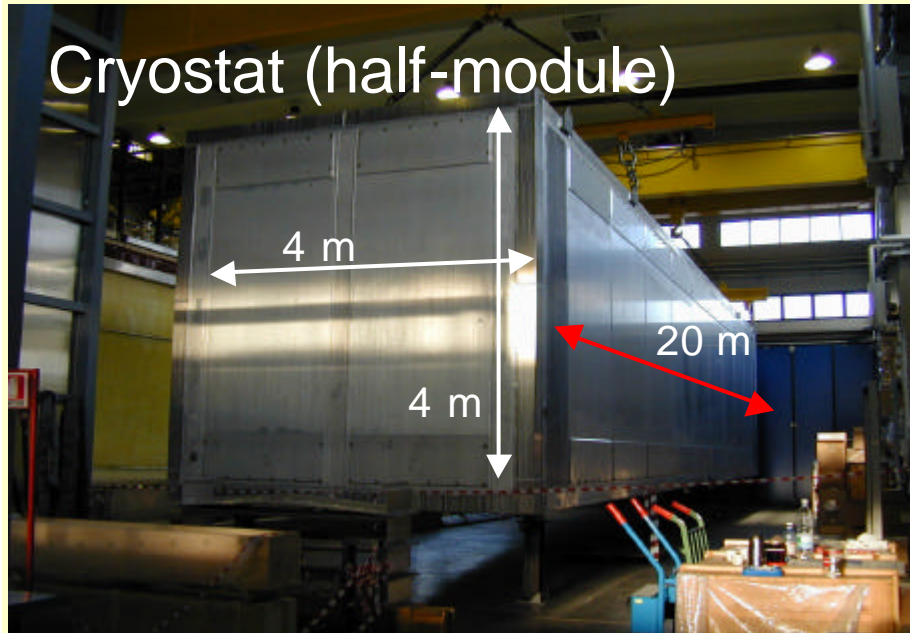


# *LAr TPC story...*

- L.W.Alvarez (late 60'): noble liquids for position sensitive detectors
- T.Doke (late 60'): systematic studies of noble liquids properties
- W.J.Willis & V.Radeka (70'): large calorimeters for HEP experiments
- C.Rubbia (1977): LAr TPC conceived and proposed
- E.Aprile, C.Giboni, C.Rubbia (1985): high purity → long drift distances
- ICARUS Coll. (1993-1994): 3 ton LAr TPC prototype
- ICARUS Coll. (1998): Neutrino detection at CERN with a 50 l LAr TPC
- ICARUS Coll. (2001): cosmic-ray test of the 300 ton industrial module
- ICARUS Coll. (2003-2004): detector/physics papers from the T300 test
- ICARUS Coll. (2004-2005): T600 installation and commissioning at LNGS

● ...

# ICARUS T300 detector



# Liquid Argon medium properties

	Water	Liquid Argon
Density (g/cm <sup>3</sup> )	1	1.4
Radiation length (cm)	36.1	14.0
Interaction length (cm)	83.6	83.6
dE/dx (MeV/cm)	1.9	2.1
Refractive index (visible)	1.33	1.24
Cerenkov angle	42°	36°
Cerenkov d <sup>2</sup> N/dEdx (b=1)	~ 160 eV <sup>-1</sup> cm <sup>-1</sup>	~ 130 eV <sup>-1</sup> cm <sup>-1</sup>
Muon Cerenkov threshold (p in MeV/c)	120	140
Scintillation (E=0 V/cm)	No	Yes (~ 50000 g/MeV @ l =128nm)
Long electron drift	Not possible	Possible (μ = 500 cm <sup>2</sup> /Vs)
Boiling point @ 1 bar	373 K	87 K

When a charged particle traverses LAr:

## 1) Ionization process

$$W_e = 23.6 \pm 0.3 \text{ eV}$$

## 2) Scintillation (luminescence)

$$W_\gamma = 19.5 \text{ eV}$$

UV "line" ( $\lambda=128 \text{ nm} \Leftrightarrow 9.7 \text{ eV}$ )

No more ionization: Argon is transparent

Only Rayleigh-scattering

## 3) Cerenkov light (if relativistic particle)

☞ Charge

☞ Scintillation light (VUV)

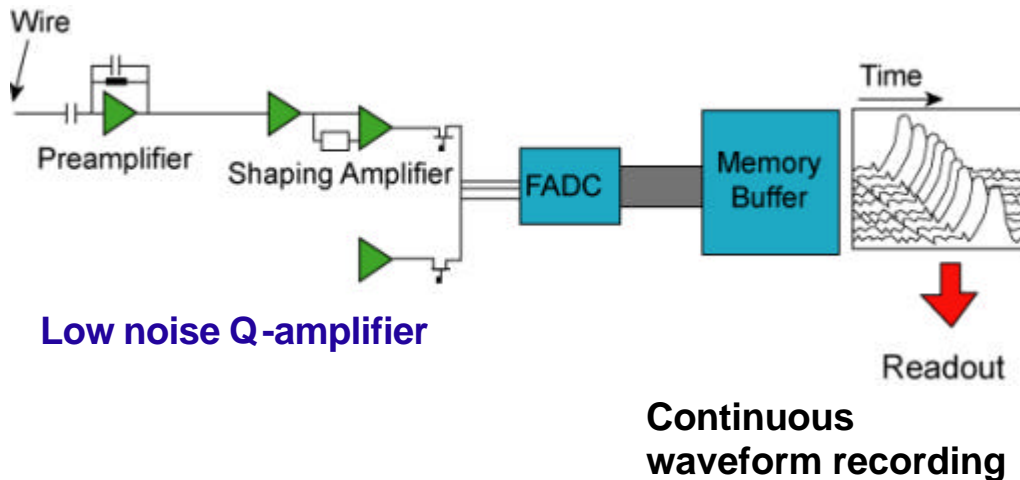
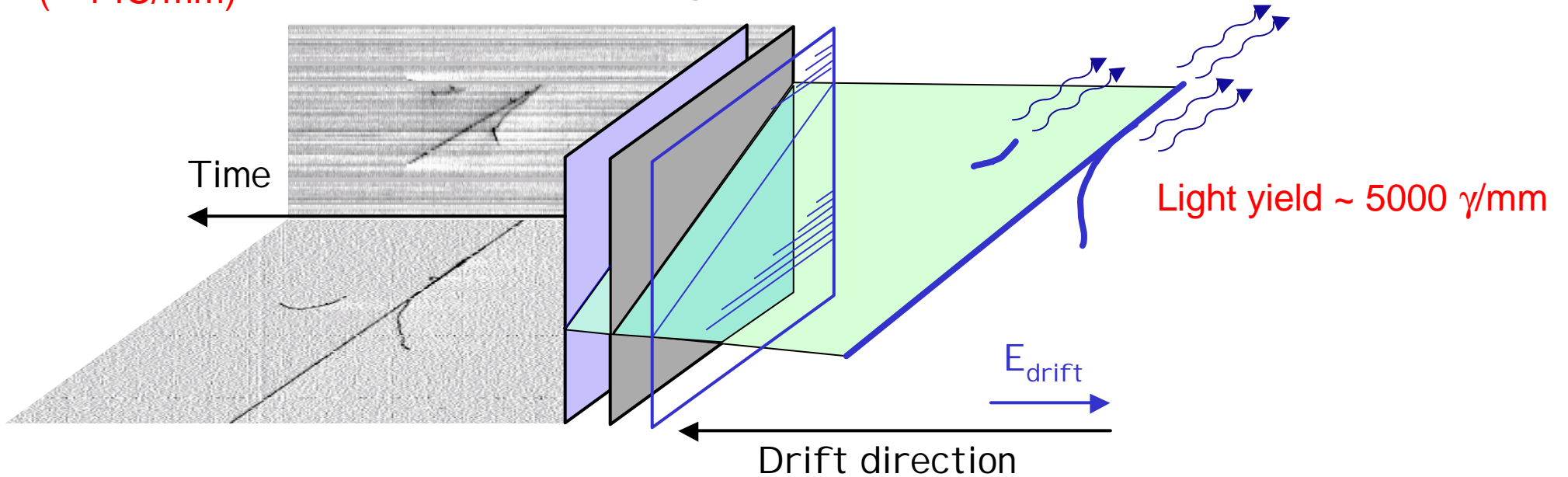
☞ Cerenkov light (if  $b > 1/n$ )

# The Liquid Argon TPC principle

Charge yield  $\sim 6000$  electrons/mm  
( $\sim 1$  fC/mm)

Charge readout planes: Q

UV Scintillation Light: L

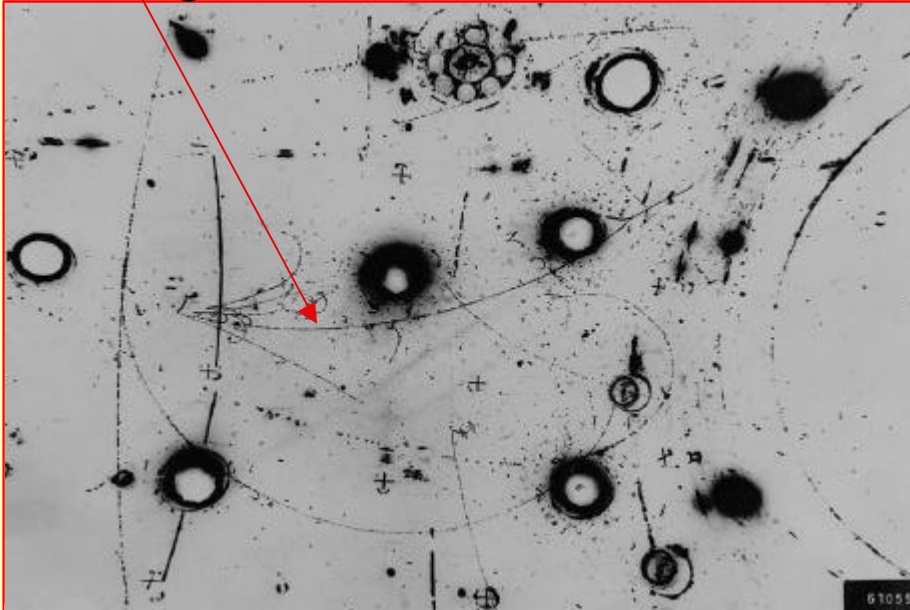


High density  
Non-destructive readout  
Continuously sensitive  
Self-triggering  
 $t_0$  available (scintillation)

# ...an electronic bubble chamber

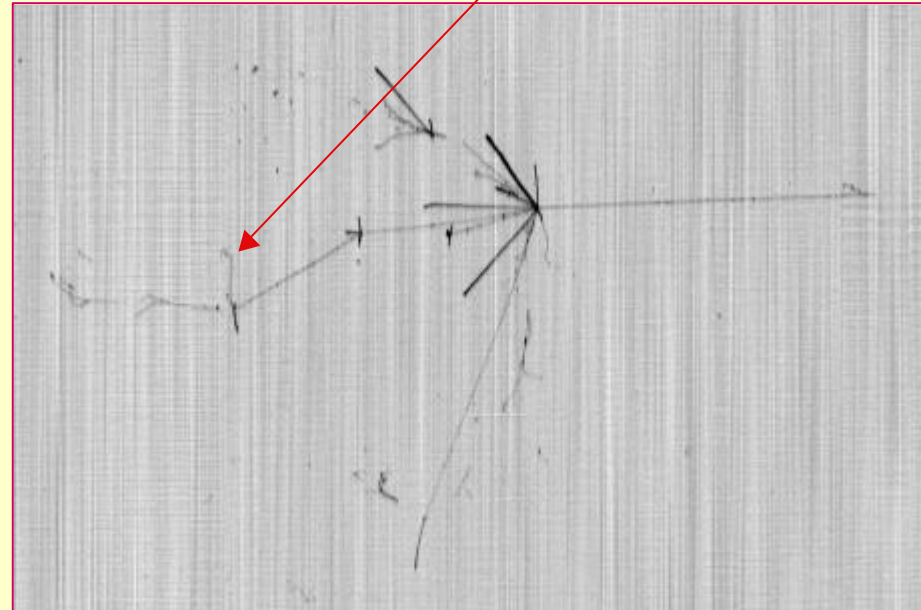
Bubble diameter ~ 3 mm  
(diffraction limited)

## Gargamelle bubble chamber

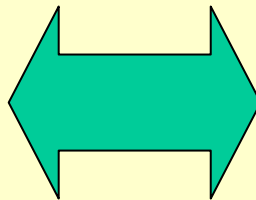


Bubble size ~ 3x3x0.4 mm<sup>3</sup>

## ICARUS electronic chamber



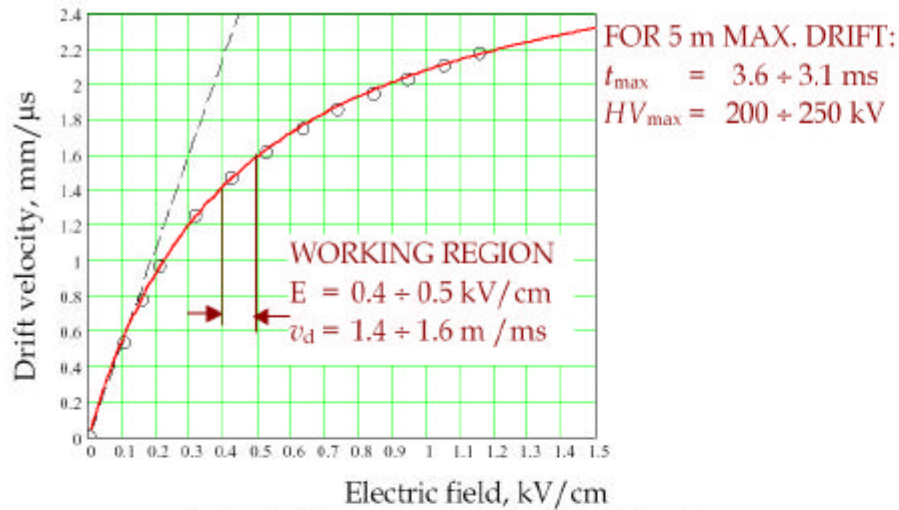
Medium	<i>Heavy freon</i>
Sensitive mass	3.0 ton
Density	1.5 g/cm <sup>3</sup>
Radiation length	11.0 cm
Collision length	49.5 cm
dE/dx	2.3 MeV/cm



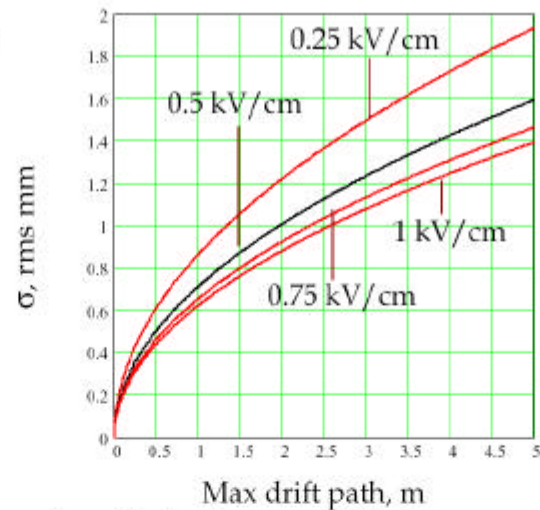
Medium	<i>Liquid Argon</i>
Sensitive mass	Many ktons
Density	1.4 g/cm <sup>3</sup>
Radiation length	14.0 cm
Collision length	54.8 cm
dE/dx	2.1 MeV/cm



# Electron drift properties in liquid Argon



Drift velocity versus electric field in liquid argon

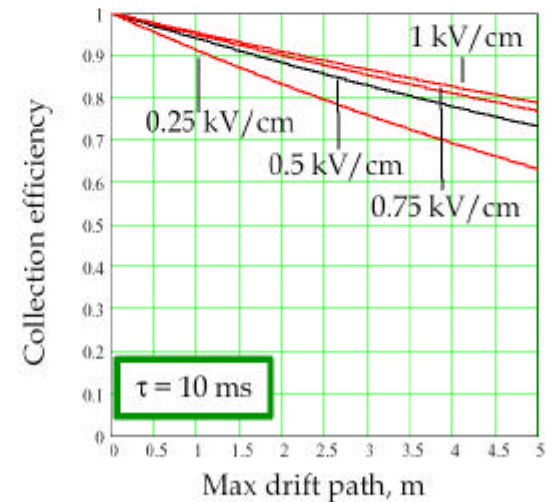
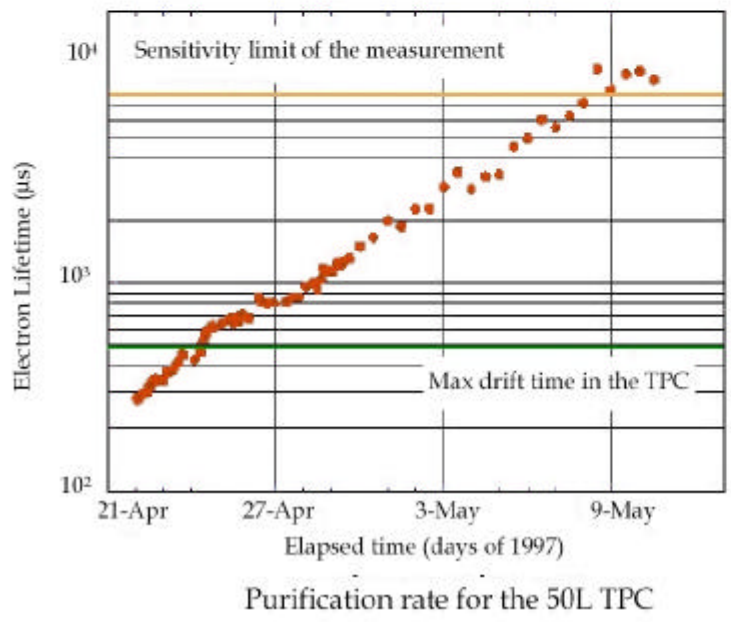


Longitudinal rms diffusion spread versus drift paths at different electric field intensities

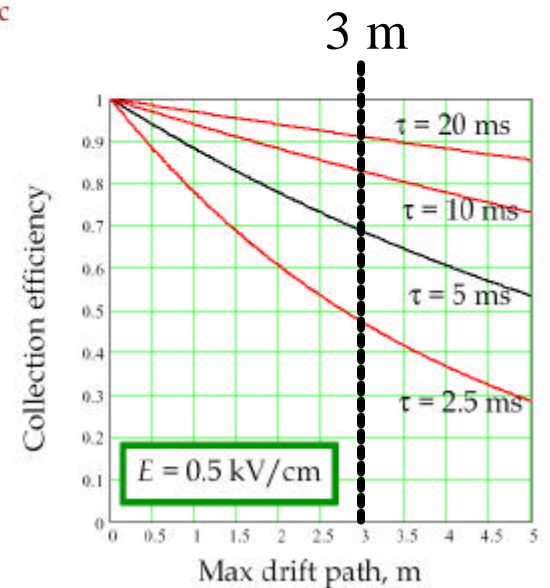
$$\sigma_D = \sqrt{2 \cdot D \cdot \frac{x}{v_d}}$$

$$D = 4.06 \text{ cm}^2/\text{s}$$

$\sigma_D = 0.9 \text{ mm} \cdot \sqrt{T_D [\text{ms}]}$   
 Longitudinal rms diffusion spread at 0.5 kV/cm  
 Average  $\langle \sigma_D \rangle = 1.1 \text{ mm}$   
 Maximum  $\sigma_{Dmax} = 1.6 \text{ mm}$

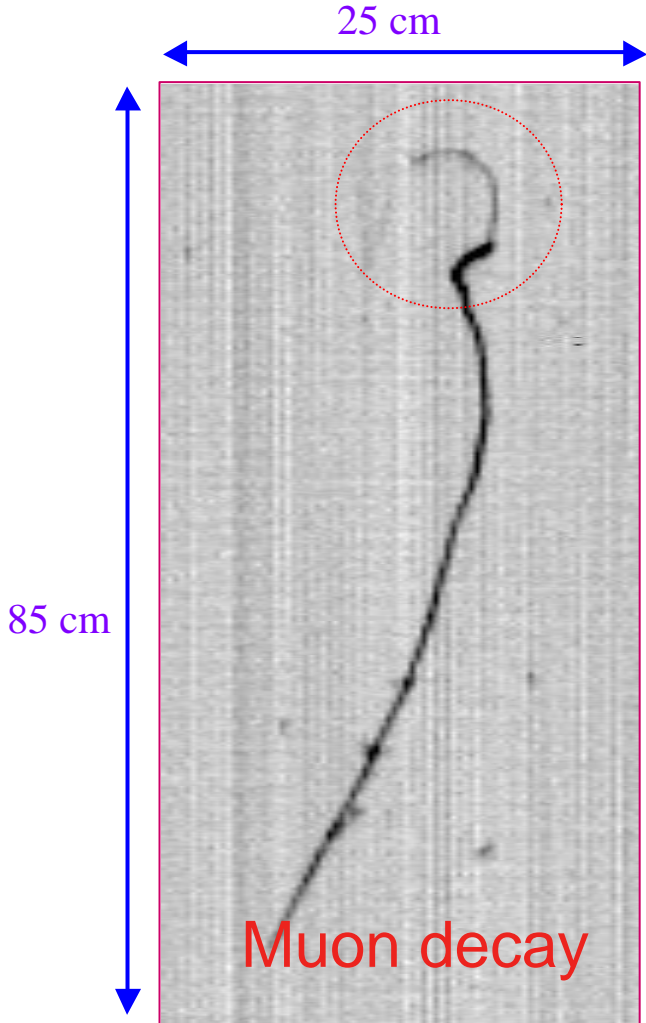


Drifting charge attenuation versus drift paths at different electric field intensities ( $\tau = 10 \text{ ms}$ )



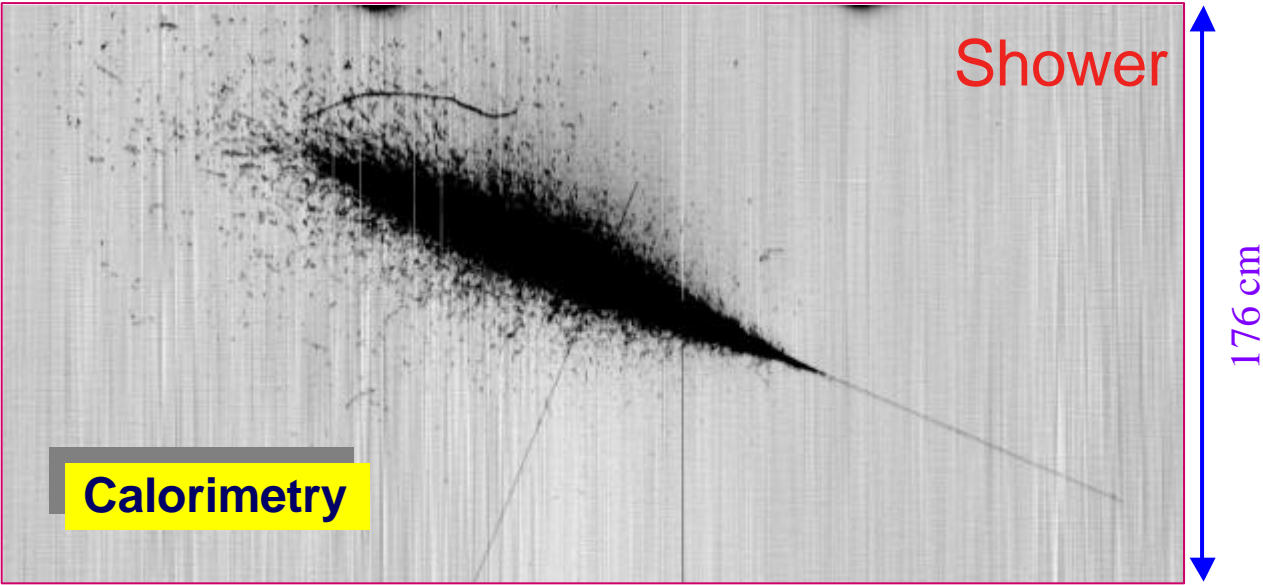
Drifting charge attenuation versus drift path at different electron lifetimes ( $E = 0.5 \text{ kV/cm}$ )

# Cosmic rays events in the ICARUS T300

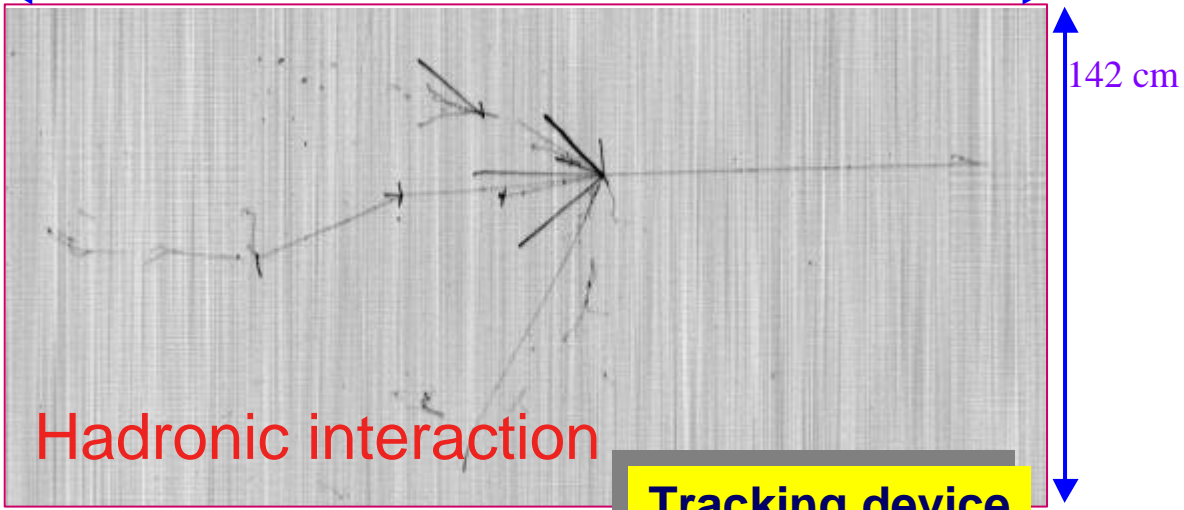


Run 960, Event 4 Collection Left

Measurement of local energy deposition  $dE/dx$

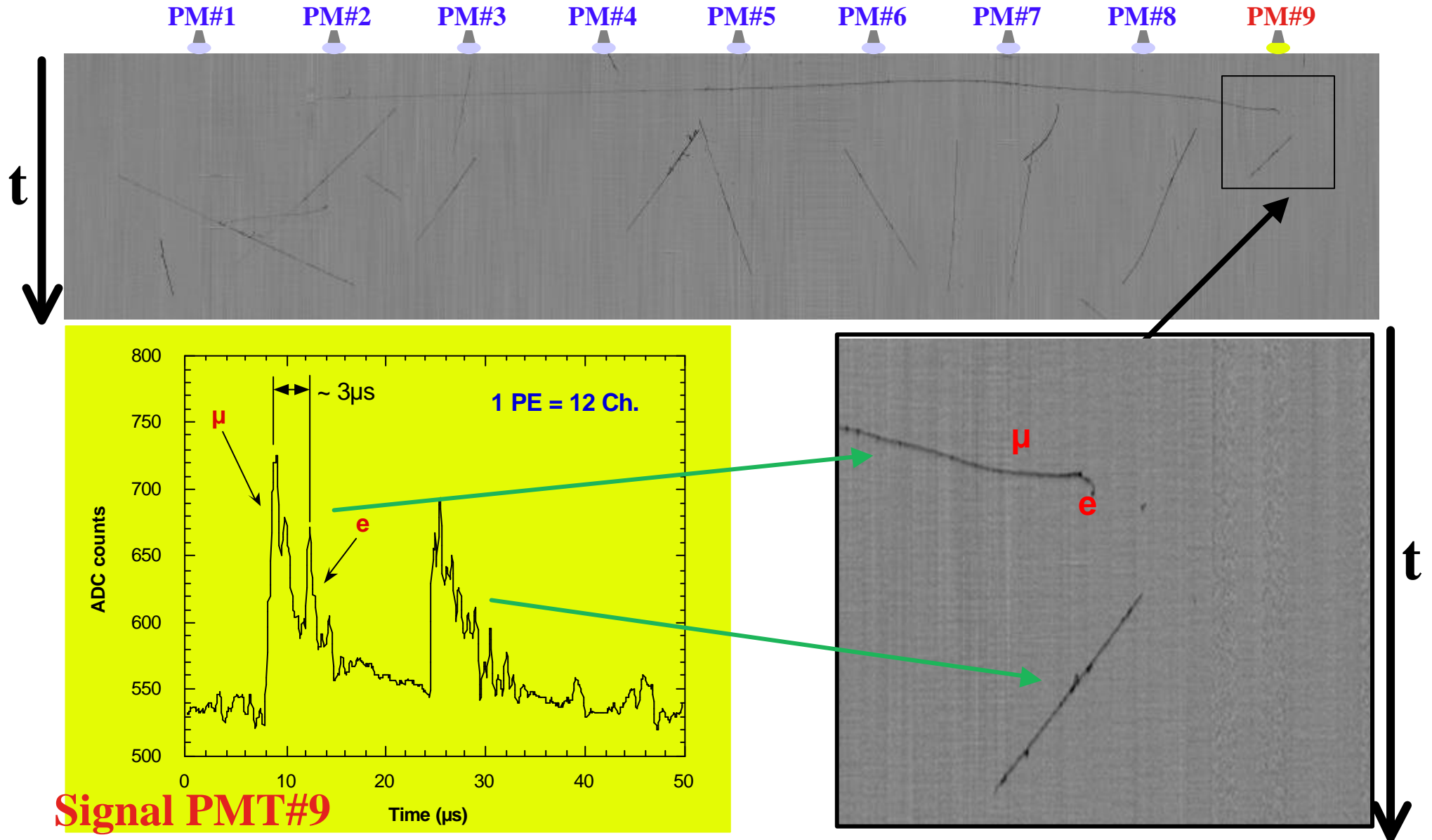


265 cm



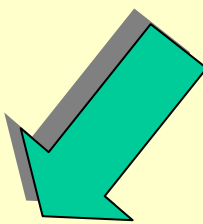
Run 308, Event 160 Collection Left

# VUV scintillation light readout



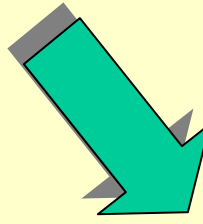
## *Liquid Argon TPC:*

physics calls for applications at two different mass scales



**100 ton**

- Precision studies of  $\nu$  interactions
- Calorimetry
- Near station in LBL facilities



**100 kton**

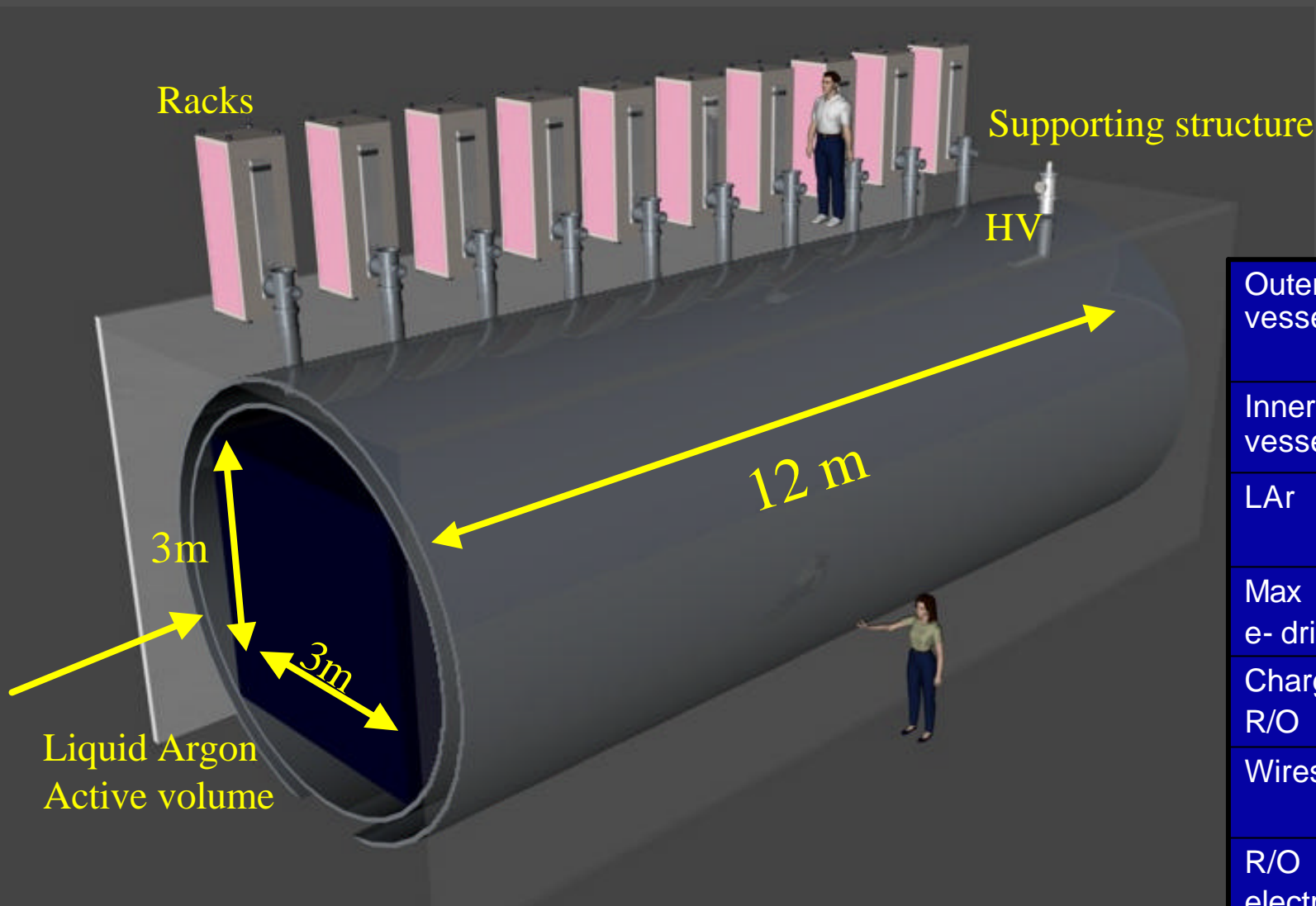
- Ultimate nucleon decay searches
- Astroparticle physics
- CP violation in neutrino mixing



Strong synergy and high degree of interplay

Ideas for a next generation liquid Argon TPC detector for neutrino physics and nucleon decay searches,  
A.Ereditato, A.Rubbia, Memo to the SPSC, April 2004.

**Conceptual design of a ~100 ton LAr TPC for a near station in a LBL facility:  
a possibility to be further explored**



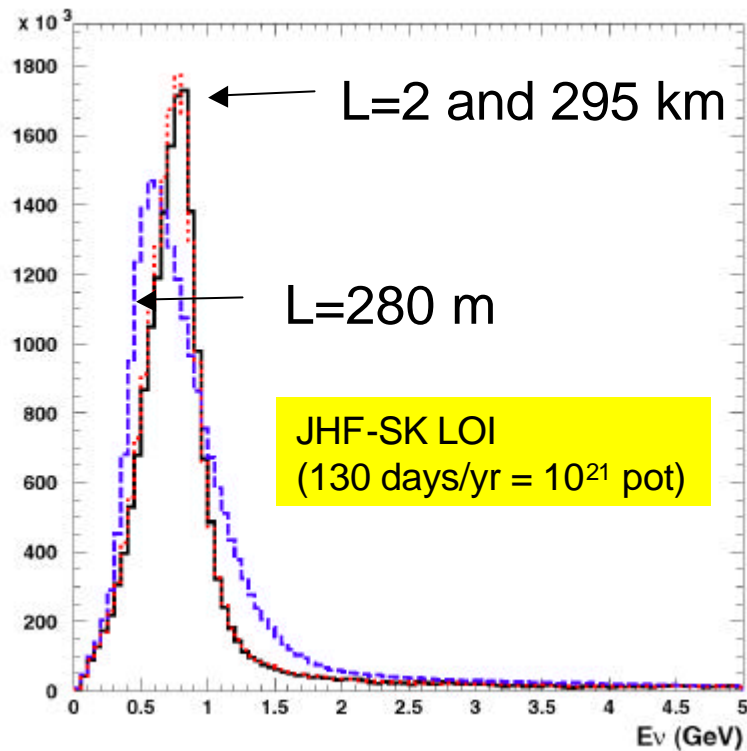
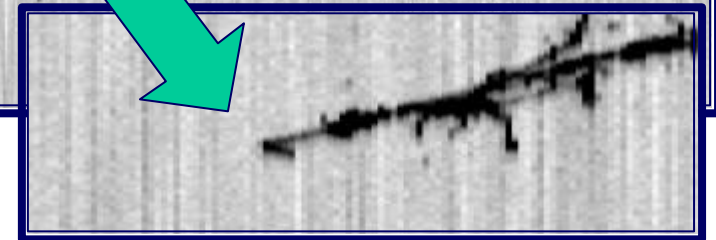
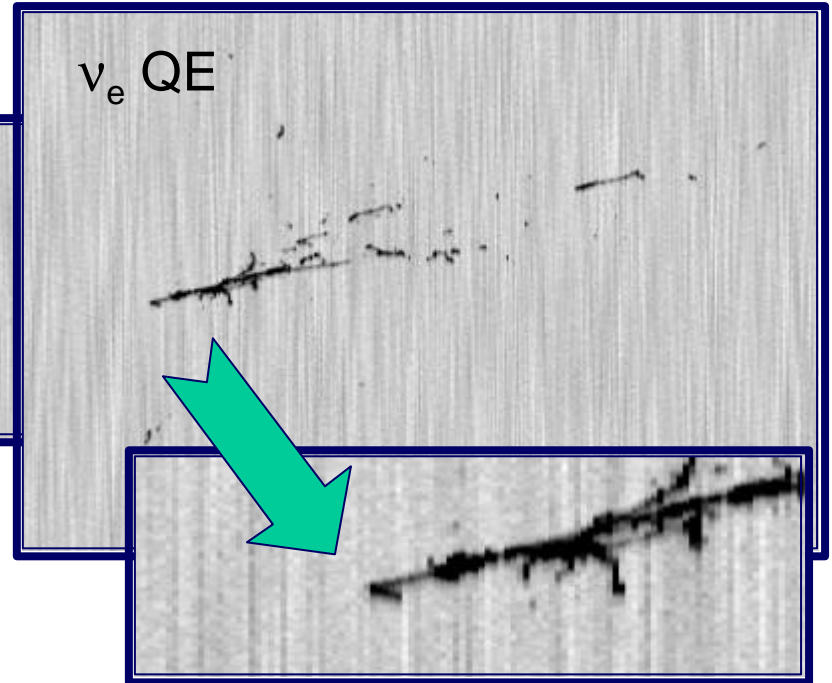
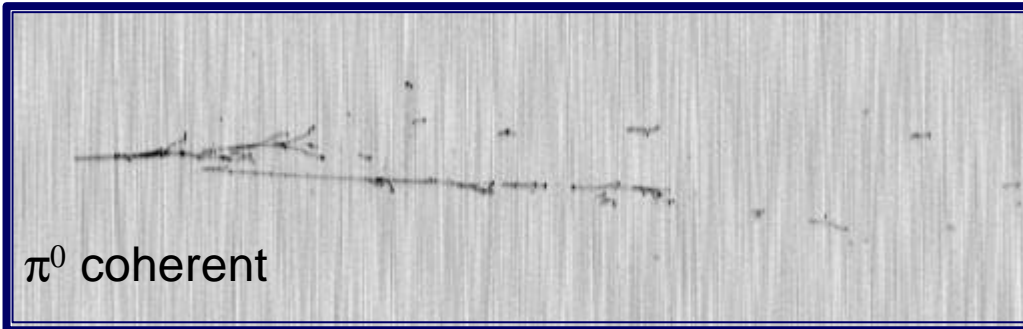
Outer vessel	$\phi \sim 5\text{m}$ , $L \sim 13\text{m}$ , 15mm thick, weight $\sim 22\text{ t}$
Inner vessel	$\phi \sim 4,2\text{ m}$ , $L \sim 12\text{ m}$ , 8 mm thick, $\sim 10\text{ t}$
LAr	<b>Total <math>\sim 240\text{ t}</math></b> <b>Fiducial <math>\sim 100\text{ t}</math></b>
Max e- drift	3 m @ HV=150 kV $E = 500\text{ V/cm}$
Charge R/O	2 views, $\pm 45^\circ$ 2 (3) mm pitch
Wires	$\sim 10000$ (7000) $f = 150\ \mu\text{m}$
R/O electr.	on top of the dewar
Scintill. light	Also for triggering

**Ideas for future liquid Argon detectors**

A.Ereditato, A.Rubbia, to appear in Proc. of NUI NT04, LNGS, March 2004

T2K would provide an ideal & high intensity beam for such a  $\sim 100$  ton detector

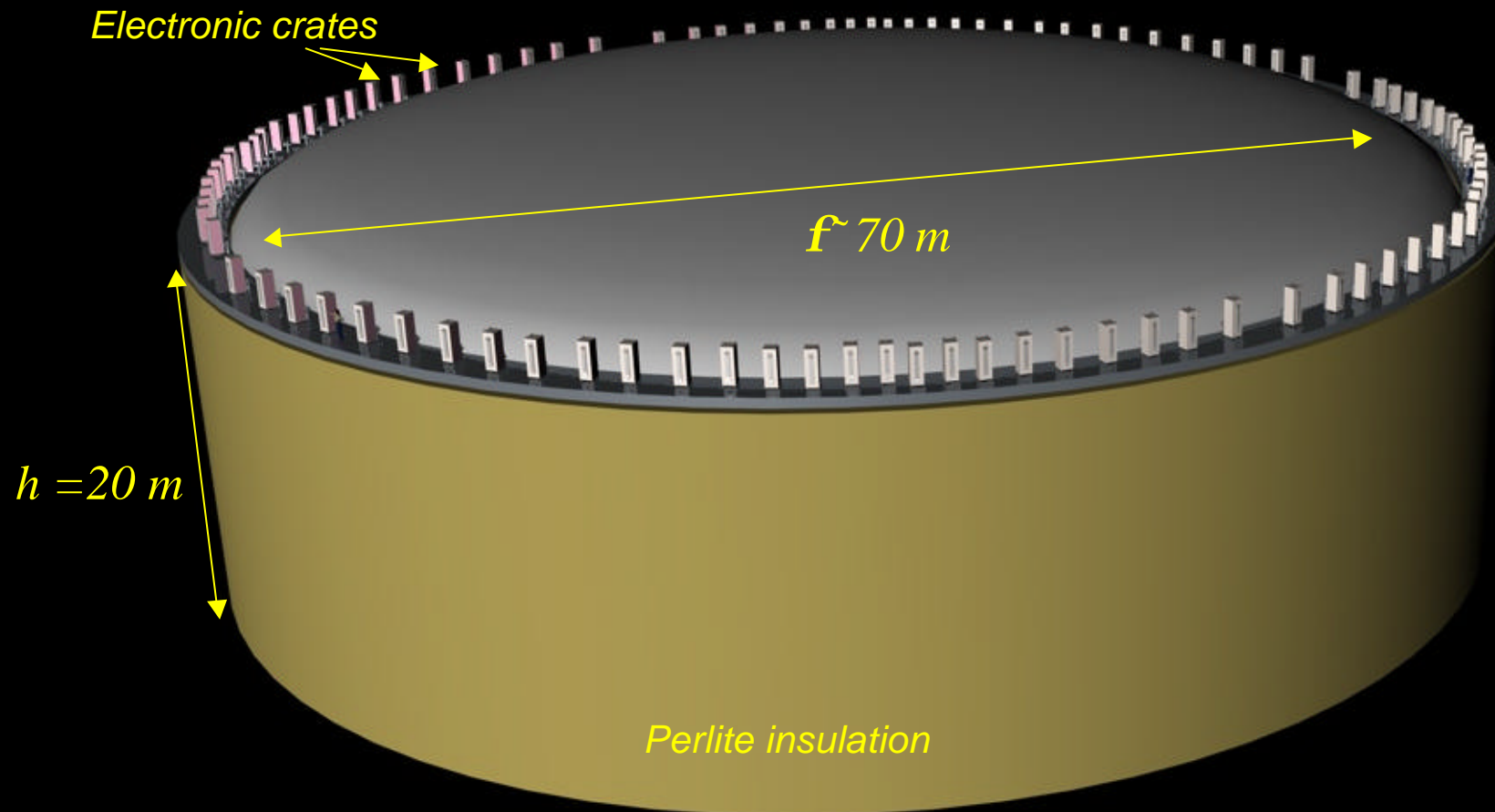
full simulation, digitization, and noise inclusion



For example: 100 ton @ L=2000 m

Beam	$E_{\text{peak}}$ (GeV)	$n_m$	$n_e$
OA2	0.7	300000/yr 0.1/spill	5800/yr 45/day

# 100 kton liquid Argon TPC detector



**Experiments for CP violation: a giant liquid Argon scintillation, Cerenkov and charge imaging experiment.**  
A.Rubbia, Proc. II Int. Workshop on Neutrinos in Venice, 2003, hep-ph/0402110

	<b>Water Cerenkov (UNO)</b>	<b>Liquid Argon TPC</b>
<b>Total mass</b>	650 kton	100 kton
<b>Cost</b>	~ 500 M\$	Under evaluation
<b><math>p \otimes e p^0</math> in 10 years</b>	$10^{35}$ years $e = 43\%$ , ~ 30 BG events	$3 \times 10^{34}$ years $e = 45\%$ , 1 BG event
<b><math>p \otimes n K</math> in 10 years</b>	$2 \times 10^{34}$ years $e = 8.6\%$ , ~ 57 BG events	$8 \times 10^{34}$ years $e = 97\%$ , 1 BG event
<b><math>p \otimes mp K</math> in 10 years</b>	No	$8 \times 10^{34}$ years $e = 98\%$ , 1 BG event
<b>SN cool off @ 10 kpc</b>	194000 (mostly $\bar{n}_e p \otimes e+n$ )	38500 (all flavors) (64000 if NH-L mixing)
<b>SN in Andromeda</b>	40 events	7 (12 if NH-L mixing)
<b>SN burst @ 10 kpc</b>	~330 n-e elastic scattering	380 $n_e$ CC (flavor sensitive)
<b>SN relic</b>	Yes	Yes
<b>Atmospheric neutrinos</b>	60000 events/year	10000 events/year
<b>Solar neutrinos</b>	$E_e > 7$ MeV (central module)	324000 events/year ( $E_e > 5$ MeV)

### Review of massive underground detectors

A.Rubbia, Proc. XI Int. Conf. on Calorimetry in H.E.P., CALOR04, Perugia, March 2004

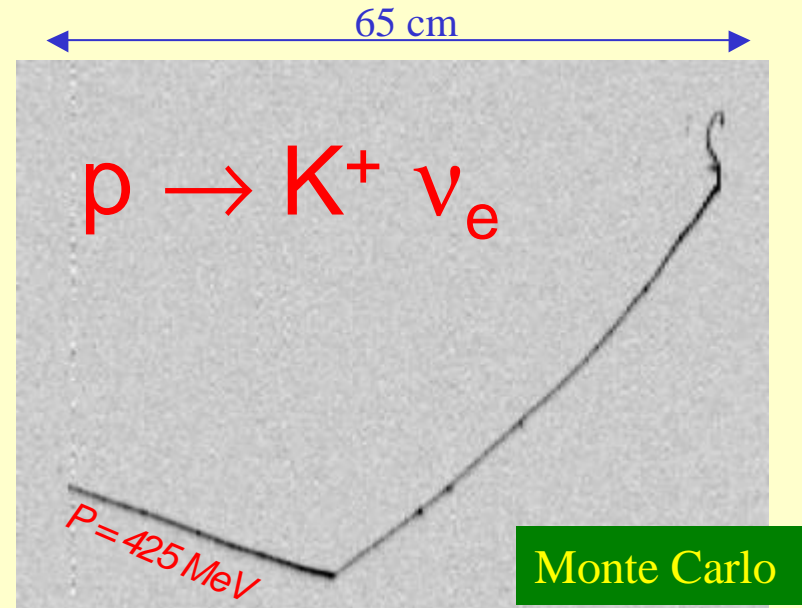
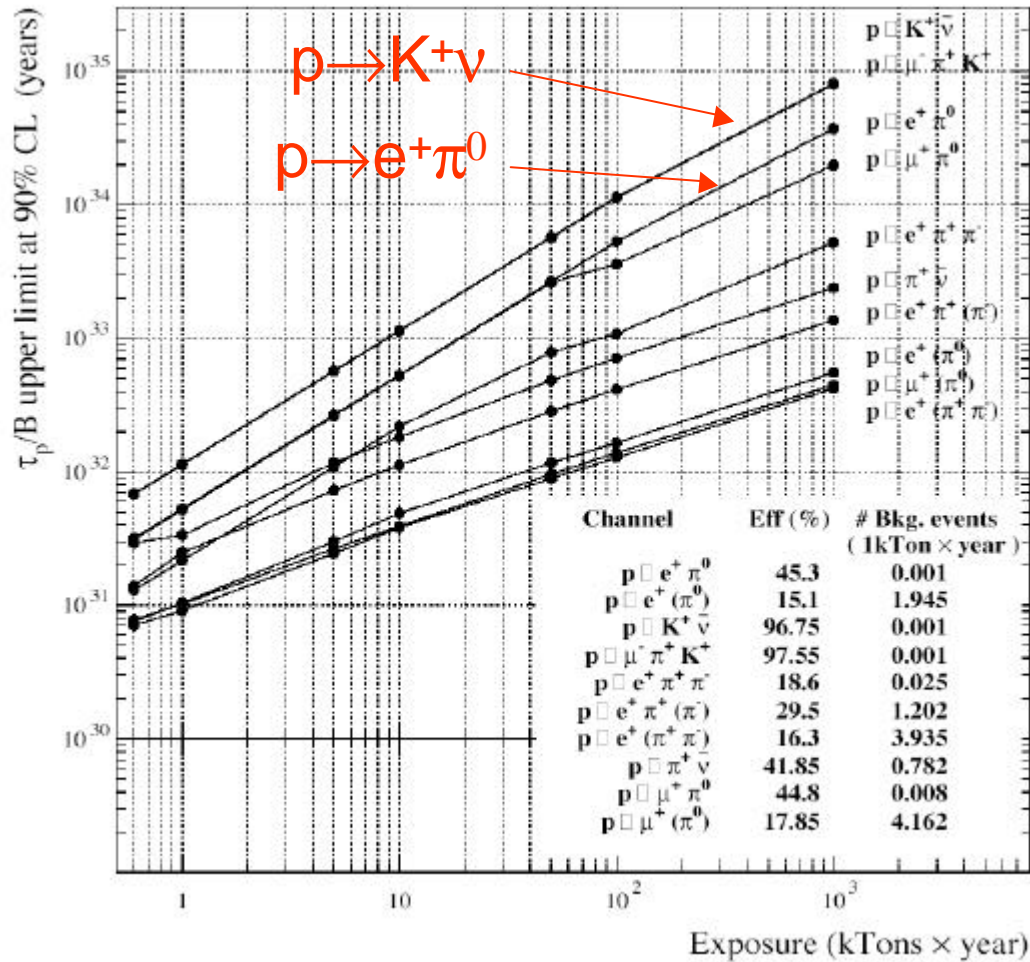
### Operation of a 100 kton LAr TPC in a future neutrino facility:

Super-Beam: 460  $\nu_\mu$  CC per  $10^{21}$  2.2 GeV protons @  $L = 130$  km

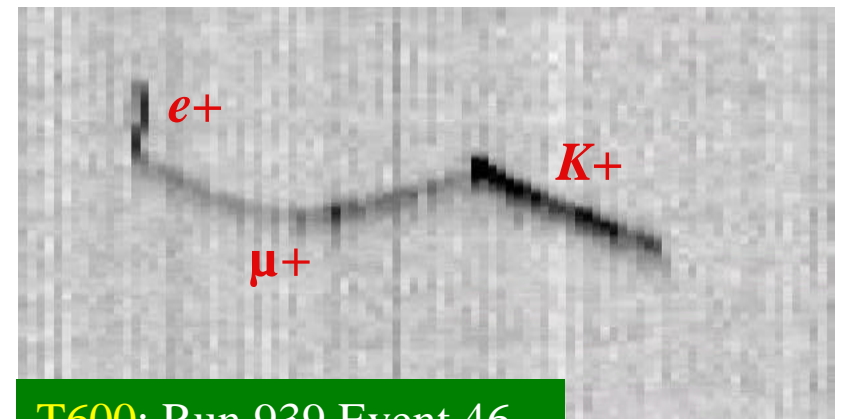
Beta-beam: 15000  $\nu_e$  CC per  $10^{19}$   $^{18}\text{Ne}$  decays with  $\gamma=75$



# Proton decay: sensitivity vs exposure



“Single” event detection capability

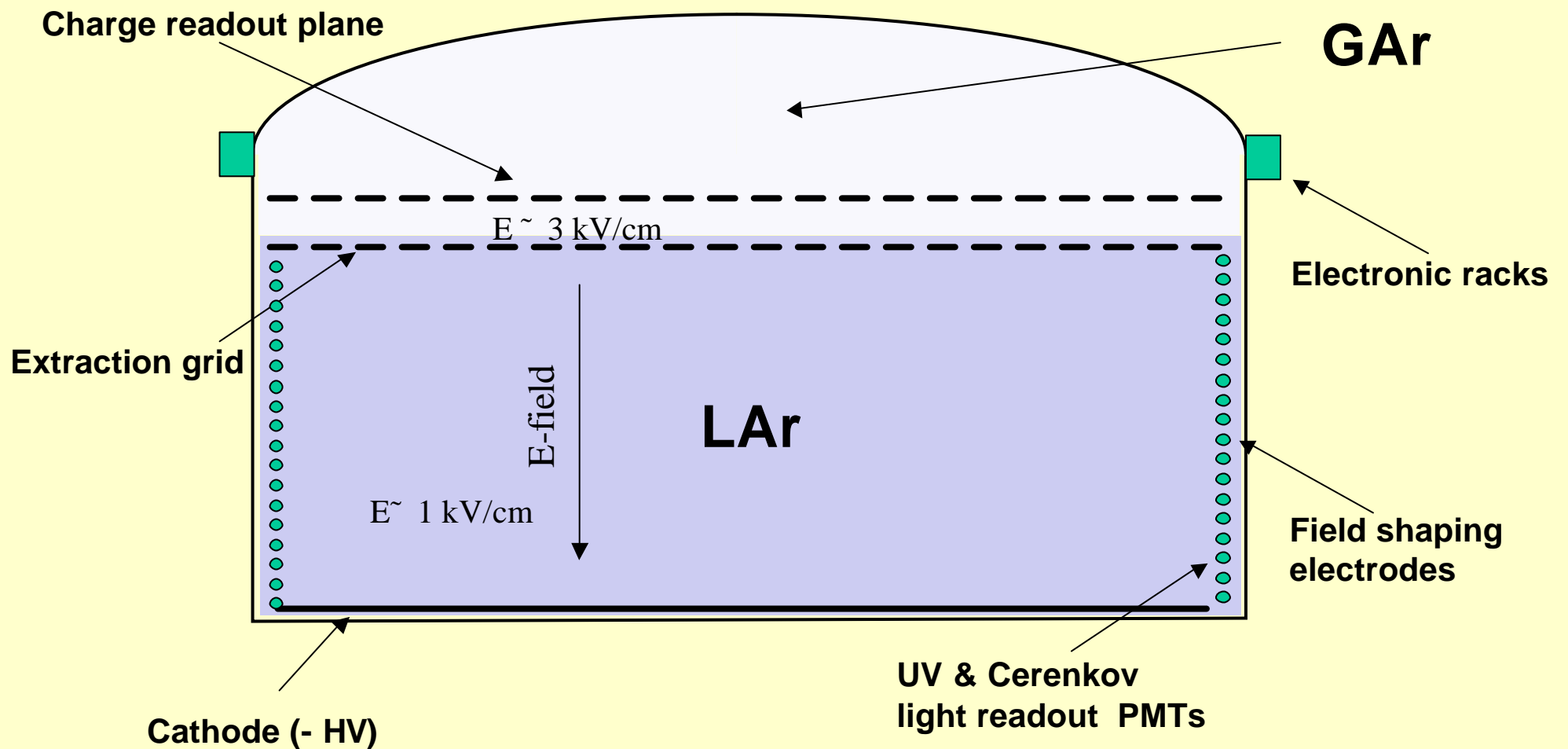


$6 \times 10^{34}$  nucleons  $\Rightarrow$

$t_p / Br > \sim 10^{34}$  years  $\times T(\text{yr}) \sim e @ 90 \text{ CL}$

# A tentative detector layout...

Single detector: charge imaging, scintillation, Cerenkov light



## *...and a tentative parameter list*

<b>Dewar</b>	$\phi \sim 70$ m, height $\sim 20$ m, perlite insulated, heat input $\sim 5$ W/m <sup>2</sup>
<b>Argon storage</b>	Boiling Argon, low pressure (<100 mbar overpressure)
<b>Argon total volume</b>	73000 m <sup>3</sup> , ratio area/volume $\sim 15\%$
<b>Argon total mass</b>	102000 tons
<b>Hydrostatic pressure at bottom</b>	3 atmospheres
<b>Inner detector dimensions</b>	Disc $f \sim 70$ m located in gas phase above liquid phase
<b>Charge readout electronics</b>	100000 channels, 100 racks on top of the dewar
<b>Scintillation light readout</b>	Yes (also for triggering), 1000 immersed 8" PMTs with WLS
<b>Visible light readout</b>	Yes (Cerenkov light), 27000 immersed 8" PMTs of 20% coverage, single g counting capability

# Charge extraction, amplification, readout

## Detector is running in **bi-phase mode**

- Long drift ( $\sim 20$  m)  $\Rightarrow$  charge attenuation to be compensated by charge amplification near anodes located in gas phase (18000 e<sup>-</sup> / 3 mm for a MIP in LAr)
- Amplification operates in proportional mode
- After maximum drift of 20 m @ 1 kV/cm  $\Rightarrow$  diffusion  $\sim$  readout pitch  $\sim$  3 mm

<b>Electron drift in liquid</b>	<b>20 m maximum drift, HV = 2 MV for E = 1 kV/cm, <math>v_d \sim 2</math> mm/<math>\mu</math>s, max drift time <math>\sim 10</math> ms</b>
<b>Charge readout view</b>	<b>2 perpendicular views, 3 mm pitch, 100000 readout channels</b>
<b>Maximum charge diffusion</b>	<b><math>s \sim 2.8</math> mm (<math>v^2 D t_{\max}</math> for D = 4 cm<sup>2</sup>/s)</b>
<b>Maximum charge attenuation</b>	<b><math>e^{-(t_{\max}/\tau)} \sim 1/150</math> for <math>\tau = 2</math> ms electron lifetime</b>
<b>Needed charge amplification</b>	<b>From 100 to 1000</b>
<b>Methods for amplification</b>	<b>Extraction to and amplification in gas phase</b>
<b>Possible solutions</b>	<b>Thin wires (<math>f \sim 30</math> <math>\mu</math>m) + pad readout, GEM, LEM, ...</b>

**LNG = Liquefied Natural Gas**

# **Cryogenic storage tankers for LNG**



support

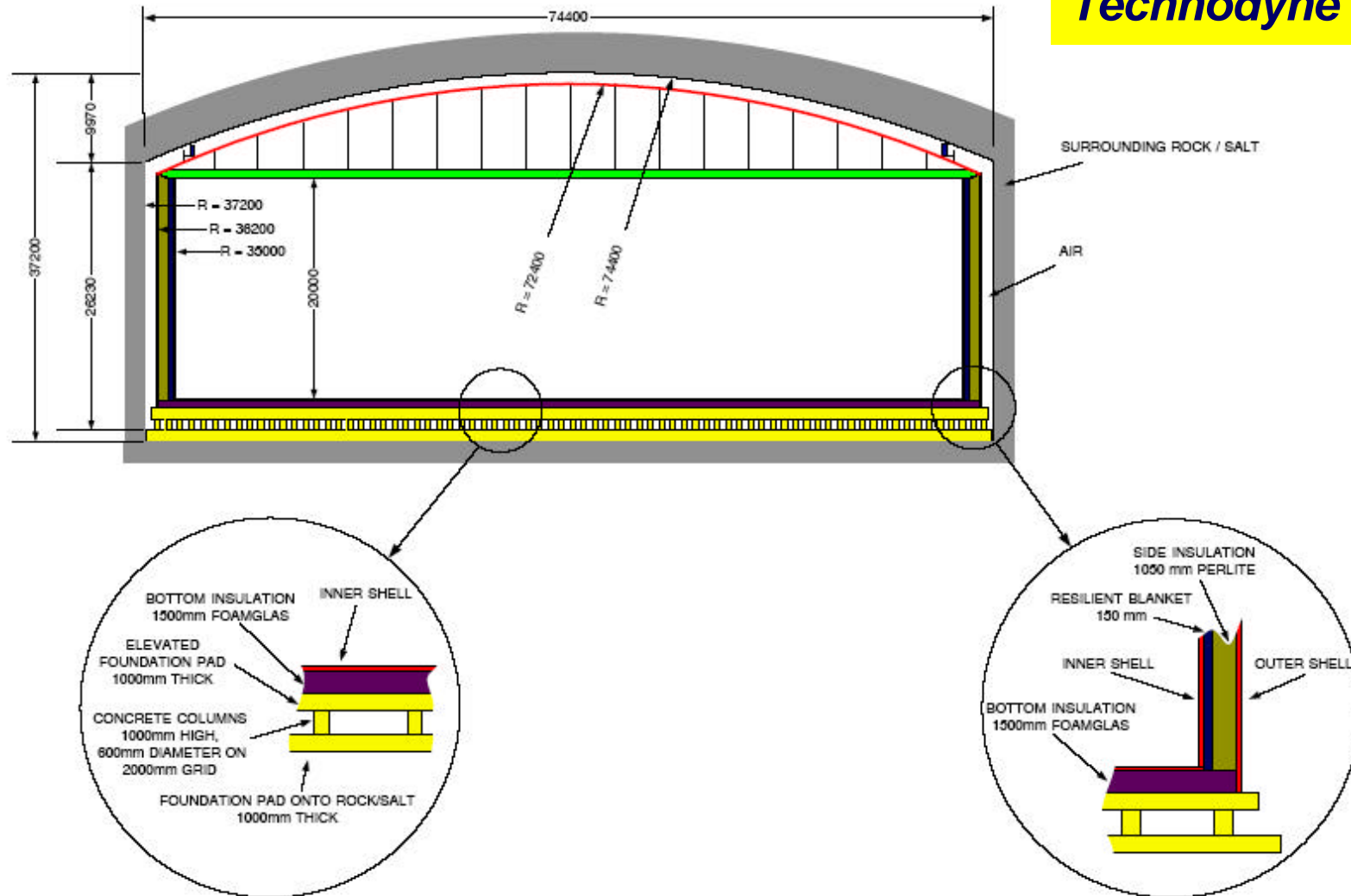
*"I learned a lot from the Shell training course. It was detailed, relevant to our business and moved at the right pace"*  
— An employee, Tigona LNG

About 2000 cryogenic tankers exist in the world, with volume up to ~ 200000 m<sup>3</sup>

Process, design and safety issues already solved by petrochemical industry



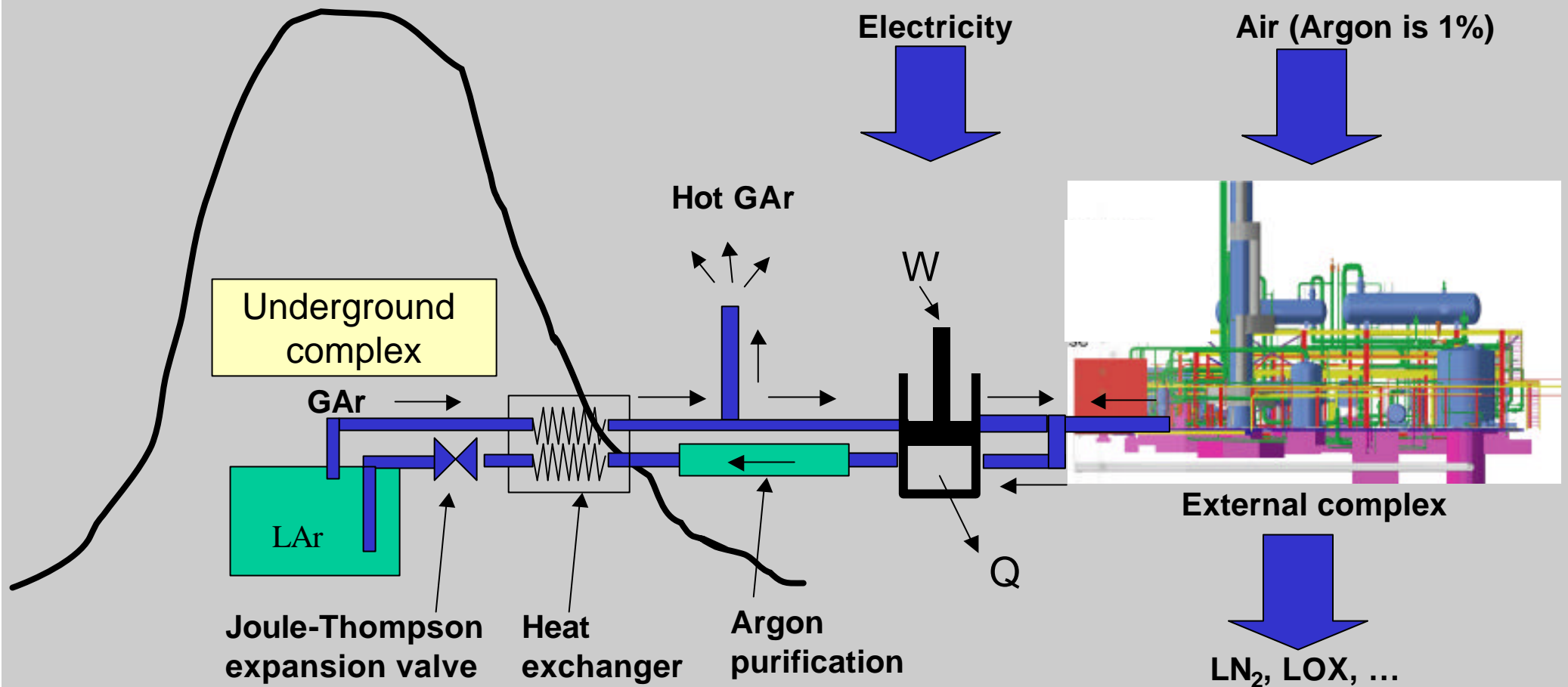
**A feasibility study for a large LAr tanker mandated to Technodyne Ltd (UK)**



**Work in progress:** Underground storage, engineering issues, process system & equipment, civil engineering consulting, safety, cost & time

# Process system & equipment

- Filling speed (100 kton): 150 ton/day  $\rightarrow$  2 years to fill, 10 years to evaporate !!
- Initial LAr filling: decide most convenient approach: transport LAr or in situ cryogenic plant
- Tanker 5 W/m<sup>2</sup> heat input, continuous re-circulation (purity)
- Boiling-off volume at regime: 30 ton/day: refilling



# ***100 kton detector: milestones***

- **Nov 2003: Venice Workshop**
  - Basic concepts: LNG tanker, signal amplification, single detector for charge imaging, scintillation and Cerenkov light readout
  - Design given for proton decay, astrophysics  $\nu$ 's, Super-Beams, Beta-Beams
  - Stressed the need for detailed comparison: 1 Mton water versus 100 kton LAr detector
- **Feb 2004: Feasibility study launched for underground liquid Argon storage**
  - Industry: Technodyne (UK) mandated for the study (expert in LNG design)
  - Design provided as input to the Fréjus underground lab study
  - Salt mine in Poland being investigated as well as other possible sites
- **March 2004: NUINT04 Workshop**
  - Identification of a global strategy: synergy between 'small' and 'large' mass LAr TPC
  - Intent to define a coherent International Network to further develop the conceptual ideas
- **April 2004 : Memo to the SPSC in view of the Villars special session (Sept. 2004)**
- **May 2004 : CERN Workshop on a future Multi MW proton source**
  - Envision a possible 10 kton full scale prototype (10% of the full detector)
  - Physics applications underground (proton decay) or at surface (neutrino beam)



# *Ongoing studies and initial R&D strategy*

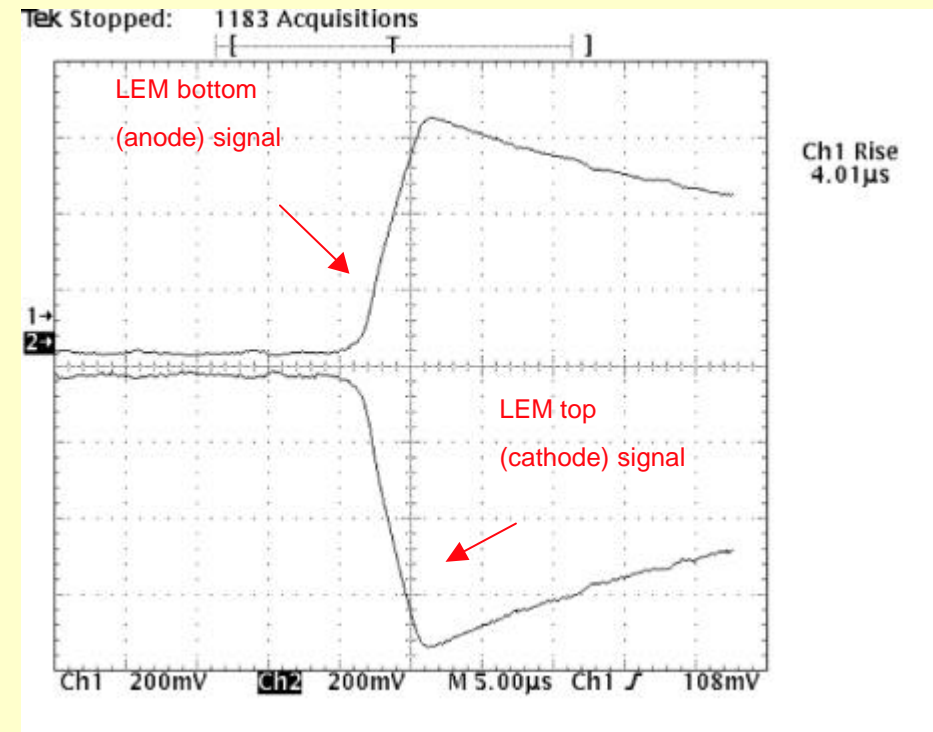
**Engineering studies, dedicated test measurements, detector prototyping, simulations, physics performance studies in progress:**

- 1) Study of suitable charge extraction, amplification and imaging devices**
- 2) Understanding of charge collection under high pressure**
- 3) Realization and test of a 5 m long detector column-like prototype**
- 4) Study of LAr TPC prototypes immersed in a magnetic field**
- 5) Study of logistics, infrastructure and safety issues for underground sites**

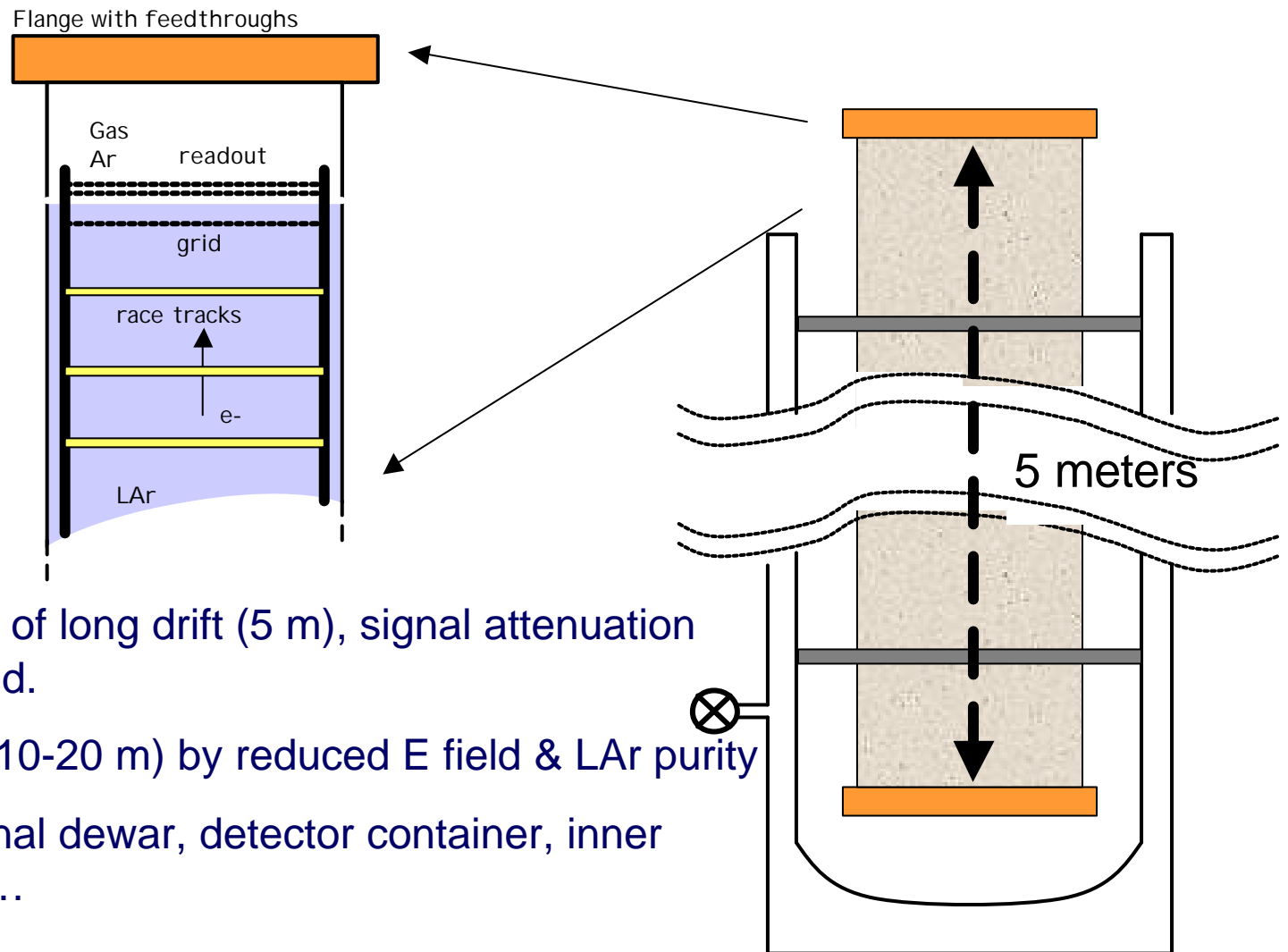
## R&D example 1: amplification with Large Electron Multiplier (LEM)

- A large scale GEM (x10) made with ultra-low radioactivity materials (copper plated on virgin Teflon)
  - In-house fabrication using automatic micro-machining
  - Modest increase in V yields gain similar to GEM
  - Self-supporting, easy to mount in multi-layers
- Resistant to discharges (lower capacitance by segmentation)
  - Cu on PEEK under construction (zero out-gassing)

P. Jeanneret et al.,  
*NIM A 500 (2003) 133-143*



## R&D example 2: long drift, extraction, amplification: test module



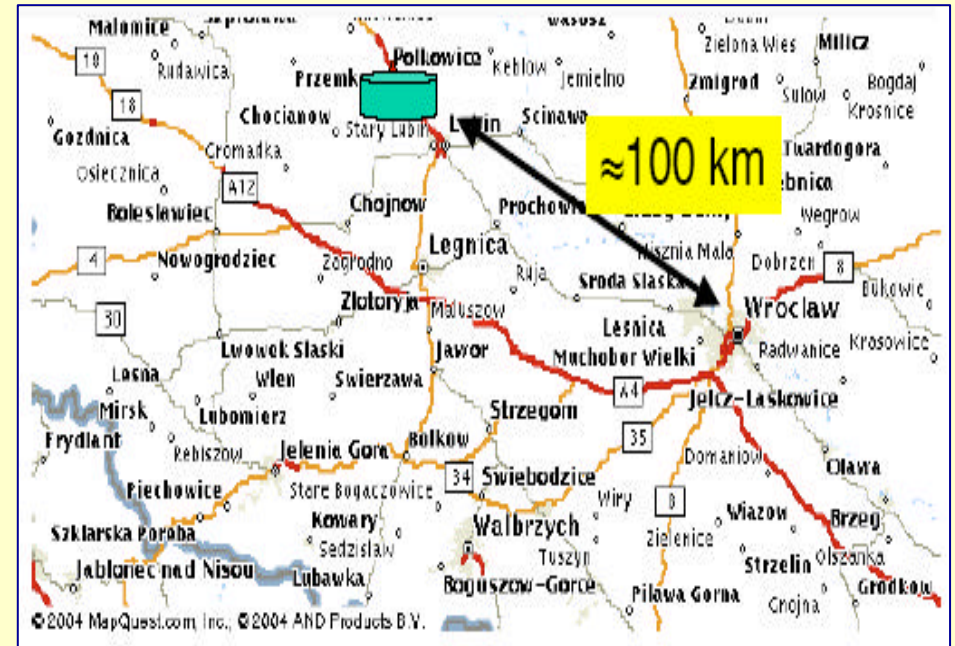
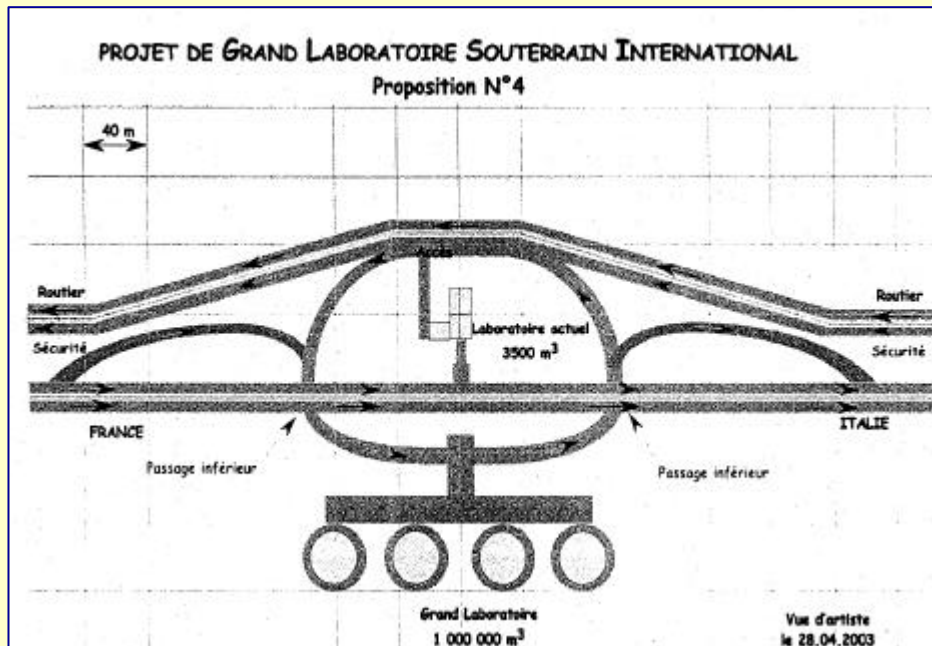
- A full scale measurement of long drift (5 m), signal attenuation and multiplication is planned.
- Simulate 'very long' drift (10-20 m) by reduced E field & LAr purity
- Design in progress: external dewar, detector container, inner detector, readout system, ...

# Possible detectors sites in Europe



# Two different site topologies

1. Hall access via highway tunnel tunnel (Fréjus laboratory project)
2. Deep mine-cavern with vertical access (CUPRUM mines, Polkowice-Sieroszowice)



- cooperation agreement: IN2P3/CNRS/DSM/CEA & INFN
- international laboratory for underground physics
- easy access
- safety issues (highway tunnel)
- caverns have to be excavated

- mines by one of the largest world producers of Cu and Ag
- salt layer at ~1000 m underground (dry)
- large caverns exist for a ~ 80000 m<sup>3</sup> (100 kton LAr) detector
- geophysics under study
- access through vertical shaft

# Argon-Net

- The further developments of the LAr TPC technique, eventually finalized to the proposal and to the realization of actual experiments, could only be accomplished by an international community of colleagues able to identify and conduct the required local R&D work and to effectively contribute, with their own experience and ideas, to the achievement of ambitious global physics goals. In particular, this is true for a large 100 kton LAr TPC detector that would exploit next generation neutrino facilities and perform ultimate non-accelerator neutrino experiments.
- We are convinced that, given the technical and financial challenges of the envisioned projects, the creation of a Network of people and institutions willing to share the responsibility of the future R&D initiatives, of the experiment's design and to propose solutions to the still open questions is mandatory.
- The actions within the Network might include the organization of meetings and workshops where the different ideas could be confronted, the R&D work could be organized and the physics issues as well as possible experiments could be discussed. One can think of coherent actions towards laboratories, institutions and funding agencies to favor the mobility of researchers, to support R&D studies, and to promote the visibility of the activities and the dissemination of the results.

So far colleagues from 21 institutions have already expressed their Interest in joining Argon-Net, to act as 'nodes' of the network

# *Conclusions and Outlook*

- The evidence for neutrino oscillations has opened the way to precision studies of the mixing matrix with accelerator neutrino experiments.
- The generation of running and planned experiments will contribute to narrow-down the errors on the oscillation parameters. The next generation will allow to pin down a non vanishing value of  $\theta_{13}$ .
- The detection of matter and of CP violating effects will likely require another generation of experiments using high intensity ( $> \text{MW}$ ) neutrino facilities with very massive detectors. At present, two options being considered: a 500-1000 kton water Cerenkov detectors (a la SK) and a 100 kton liquid Argon TPC.
- The liquid Argon TPC imaging has reached a high level of maturity thanks to many years of R&D effort conducted by the ICARUS collaboration. The plan is to operate the kton mass scale detector at LNGS with the ICARUS project.
- The technique is suitable for applications at very different different mass scales:
  - ~ 100 kton: proton decay, high statistics astrophysical & accelerator neutrinos
  - ~ 100 ton: systematic study of neutrino interactions, near detectors at LBL facilities
- In particular, a 100 kton, monolithic LAr TPC based on the industrial technology of LNG tankers and on the bi-phase operation is conceivable. R&D studies are in progress and a global strategy is being defined, possibly envisioning the construction of a 10% mass full scale prototype.
- **Look forward to the creation of a dedicated world-wide Network on the LAr TPC technique.**