

# SUPER-ICANOE

*Presented by Antonio Bueno (ETHZ)*

## *ICARUS Collaboration*

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NNN00-Fermilab Nucleon Decay and Neutrino Detector Workshop

August 7-8, 2000

Fermi National Accelerator Laboratory

Batavia, Illinois, USA



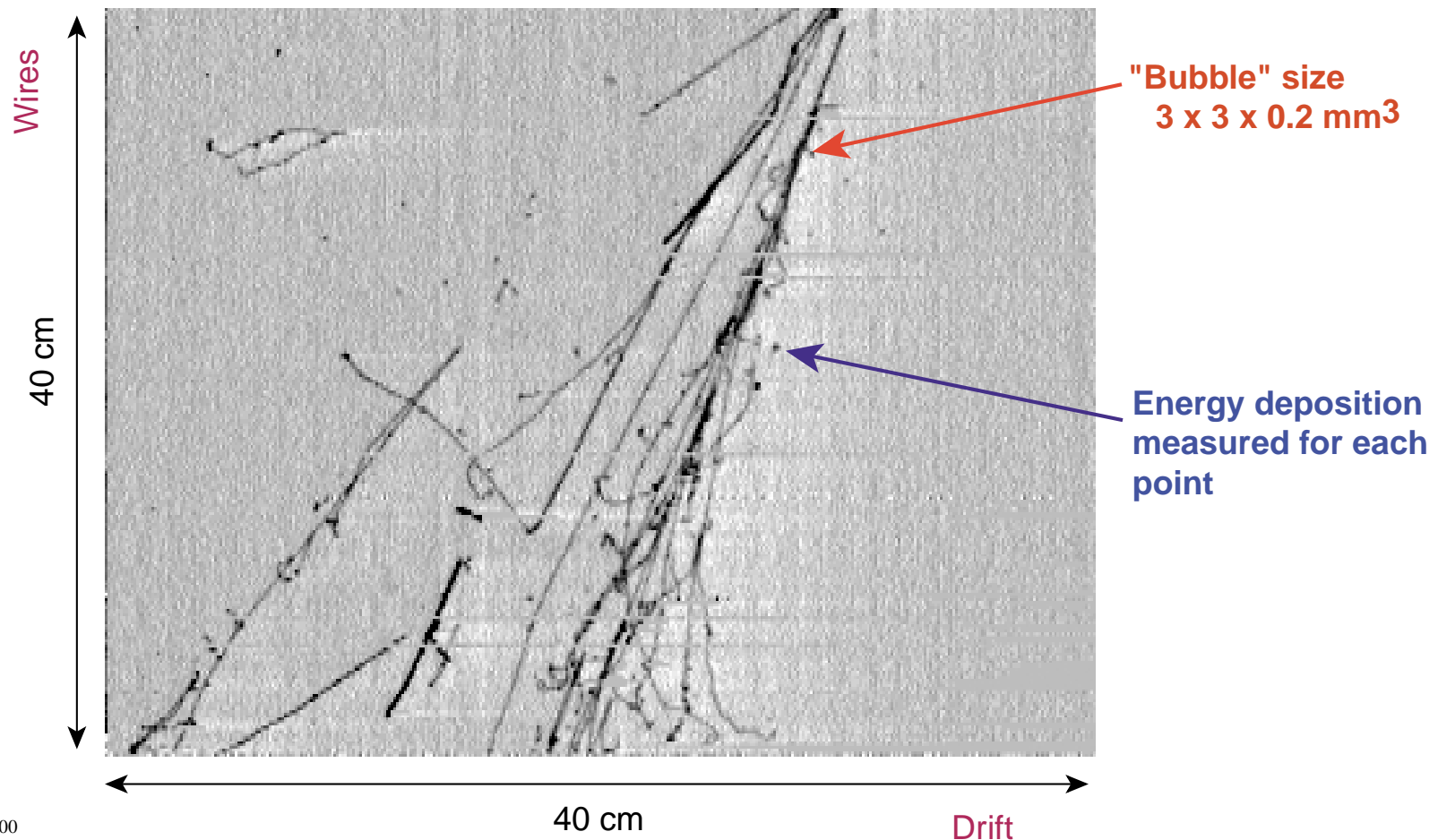
# New challenges

- High-energy physics is now facing the new challenge of thinking of a new generation of experiments beyond the current accelerator & non-accelerator ones in order to:
  - ⇒ **Improve significantly the sensitivity to nucleon decay in the range of  $10^{34}$  years**
  - ⇒ **Allow transcontinental very long-baseline ( $L > 1000\text{km}$ ) oscillations experiments with full event reconstruction**
  - ⇒ **Continue to observe solar, atmospheric neutrinos with higher statistics & better resolution**

**Need to further develop high-granularity detectors with masses in the range of ten's-of-ktons !**

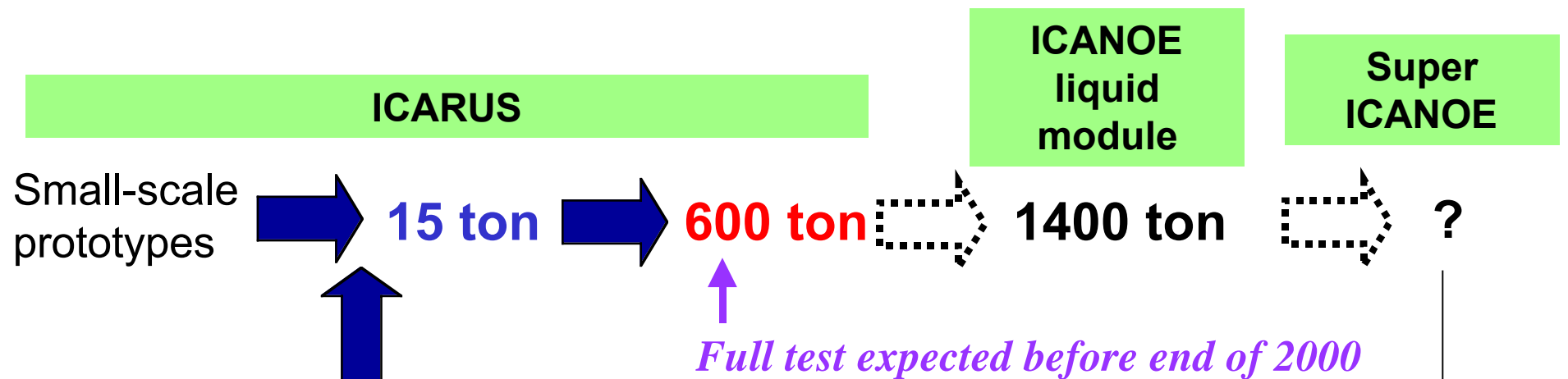
# Liquid argon imaging TPC

- The LAr TPC technique is based on the fact that **ionization electrons can drift over large distances (meters) in a volume of purified liquid Argon under a strong electric field**. If a proper readout system is realized (i.e. a set of fine pitch wire grids) it is possible to realize a massive "electronic bubble chamber", with superb 3-D imaging.



# ICARUS state of the art

- After several years of R&D and prototyping, the ICARUS collaboration is now realizing the first **600 ton module**, which will be installed at Gran Sasso in the year 2001.



## Cooperation with specialized industries:

- ➔ **Air Liquide** for Cryostat and Argon purification
- ➔ BREME Tecnica for internal detector mechanics
- ➔ CAEN for readout electronics

# ICARUS 15 ton ( $10\text{m}^3$ ) prototype

T15 installation @ LNGS (Hall di Montaggio)

- A recent major step of the R&D program has been the construction and operation of a  **$10\text{m}^3$  prototype**

- ① **Test of the cryostat technology**
- ② **Test of the “variable-geometry” wire chamber**
- ③ **Test of the liquid phase purification system**
- ④ **Test of trigger via scintillation light**
- ⑤ **Large scale test of final readout electronics**

*→ First operation of a 15 ton LAr mass as an actual “detector”*



# The ICARUS T600 module

**Under construction**

Number of independent containers = 2

Single container Internal Dimensions: Length = 19.6 m , Width = 3.9 m , Height = 4.2 m

Total (cold) Internal Volume = 534 m<sup>3</sup>

Sensitive LAr mass = 476 ton

Number of wires chambers = 4

Readout planes / chamber = 3 at 0° , ± 60° from horizontal

Maximum drift = 1.5 m

Operating field = 500 V / cm

Maximum drift time 1 ms

Wires pitch = 3 mm

Total number of channels = 58368

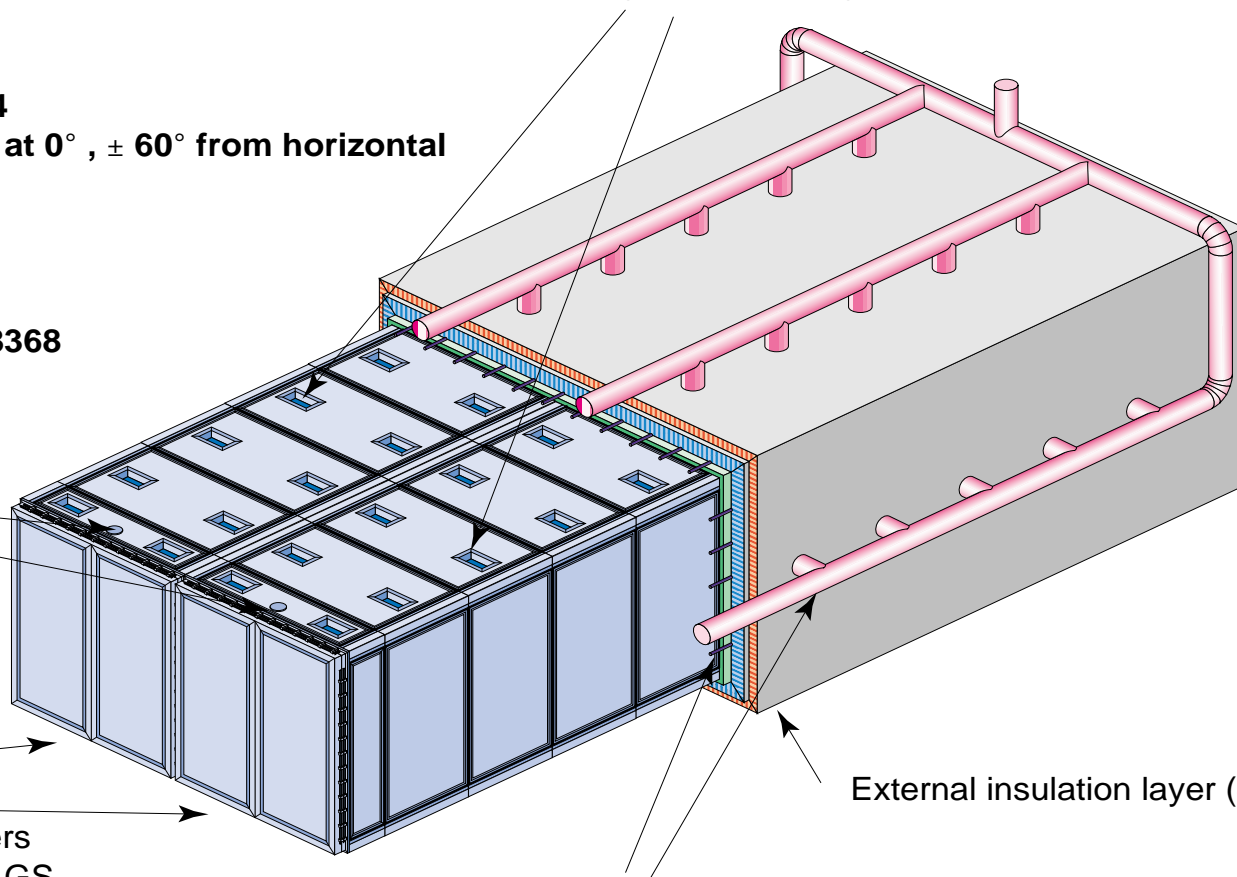
HV feedthroughs

Signal feedthroughs

External insulation layer (400 mm)

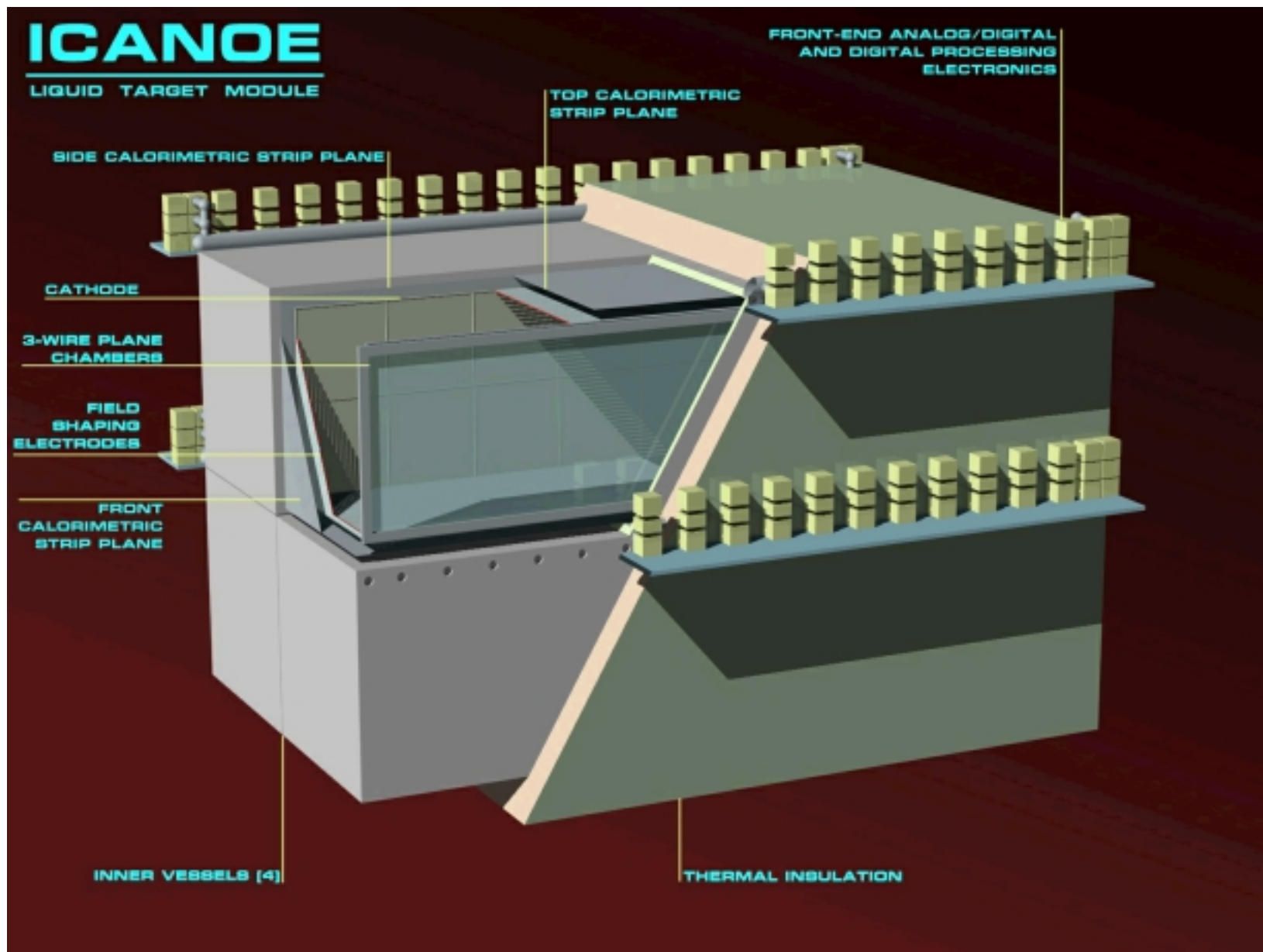
LN2 cooling circuit

2 independent aluminum containers  
each one transportable inside the GS  
Laboratory





# The ICANOE T1400 module



# Design considerations

## What we get for 30 ktons:

- **Number of targets for nucleon stability:**

- $2 \times 10^{34}$  nucleons  $\Rightarrow \tau_p$  ( $10^{33}$  years)  $> 3.6 \times T(\text{yr}) \times \epsilon$  @ 90 C.L.

- **Neutrino factory:**

- **7000**  $\nu_\mu$  CC per  $10^{20}$   $\mu$  @  $L = 7400$  km

- **Atmospheric:**

- **6000** atm CC events / year

- $\approx 30$   $\nu_\tau$  CC /year from oscillations

- **Solar:**

- $\approx 50000$  solar neutrinos / year @  $E > 5$  MeV

**Of course, MASS is not the whole story!**

**➔ We want the factor MASS  $\times$  EFFICIENCY high and BACKGROUNDS low!!!**



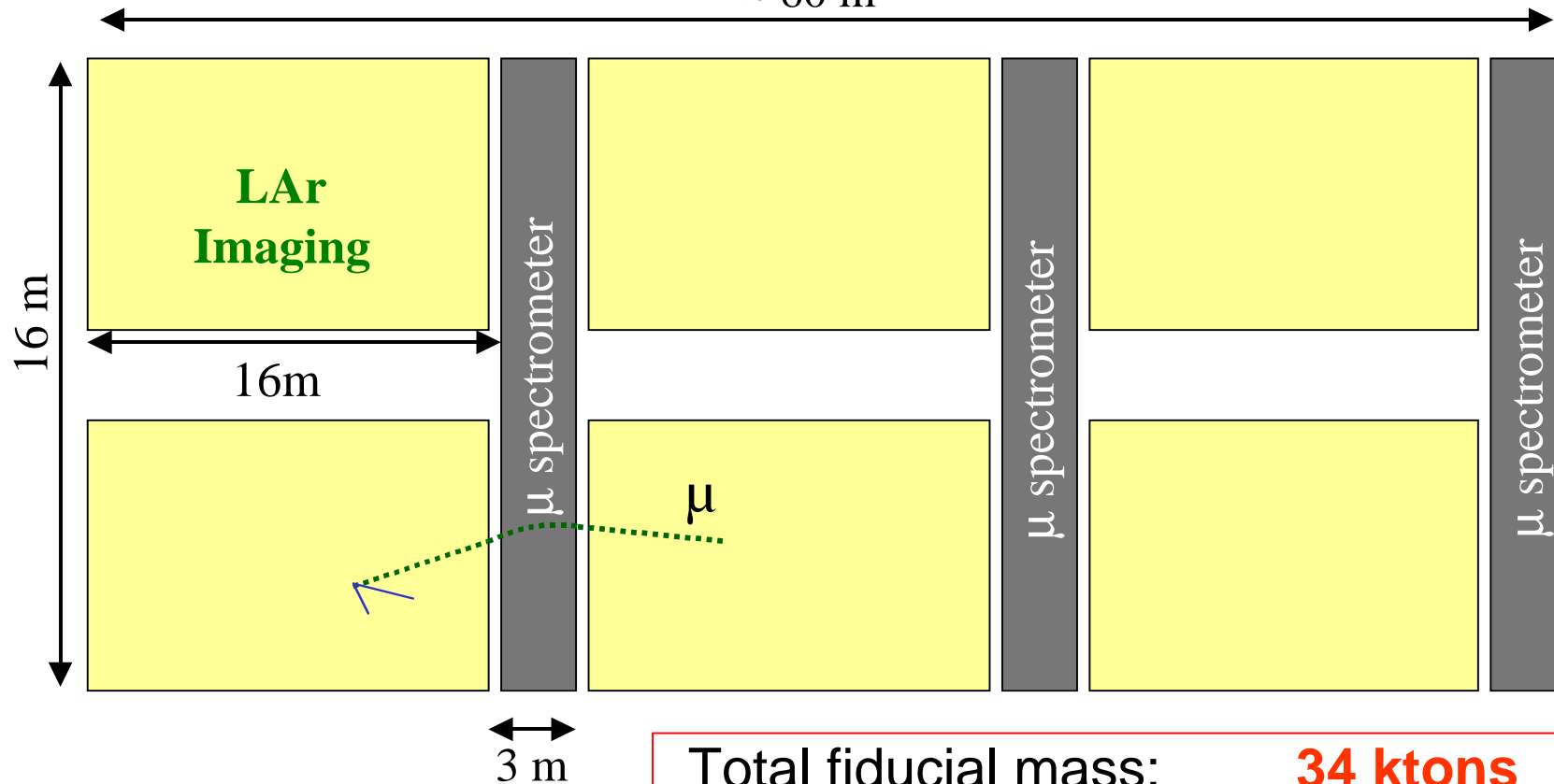
# Design option: Modular

- The ICARUS detector can be built in **very large sizes** and after having been developed in laboratory units, the liquid argon technology is now fully industrialized
  - ⇒ **Modules can be “ordered” to the producer**
- As a case study, the simplest way to extend the mass is to **replicate** a large number of times the current ICANOE supermodule
  - ⇒ **Conservative cost estimate (from ICANOE proposal):**  
**≈ 13 M\$ / supermodule (1400 tons)**
  - ⇒ **Total for 24 supermodules:** **≈310 M\$ ⇒ 34 ktons**
- **Cost reduction** possible by further geometry optimization
  - ⇒ **c.f. SUPERICARUS proposal CERN/SPSC 98-33**
- Other options: fill existing cavern e.g. SuperK ?

# Possible baseline configuration

(side view)

≈ 60 m

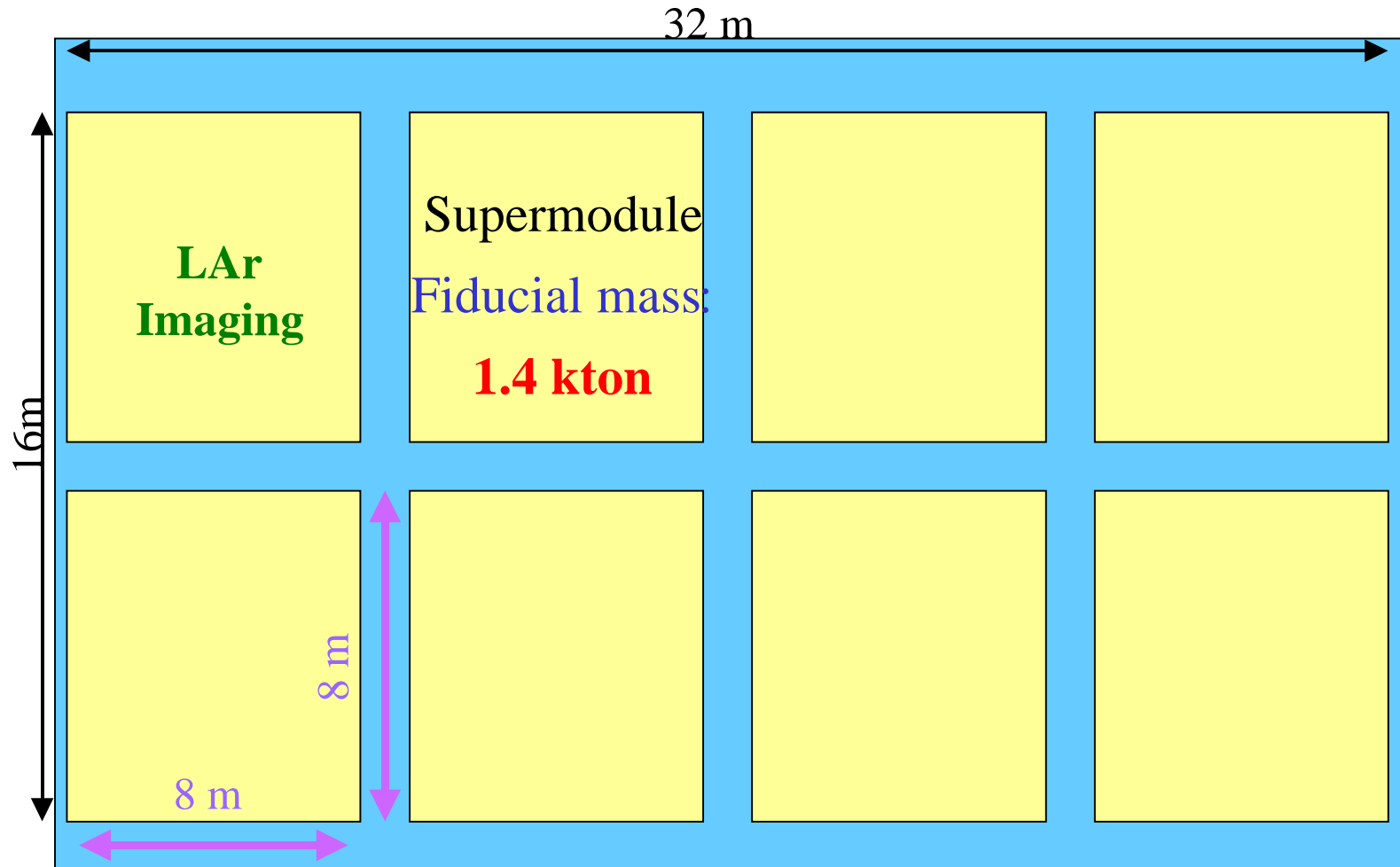


**SuperICANOE**  
(case study for  
this workshop)

Total fiducial mass:	<b>34 ktons</b>
wires/supermodule:	<b>53248</b>
channels/supermodule:	<b>26624</b>
maximum drift length:	<b>4 m</b>

# Possible baseline configuration

(front view)



Three arrays of eight supermodules: **24 supermodules in total**

# Design limitations?

- Detector is a surface detector, so gain in volume/surface ratio
  - ⇒ **Like Water Cerenkov**
- Practical limitations for size of monolithic supermodule (1000 m<sup>3</sup> for ICANOE)
  - ⇒ **Drift length**
    - Can one drift longer than currently assumed 4 meters? ⇒ R&D
  - ⇒ **Readout-wires length is not a limiting factor**
    - Variable geometry chambers (spring system tested in T15)
    - Diameter 150 μm
    - Eventually, limited by wire capacitance (noise)
  - ⇒ **Quantity of Argon**
    - ≈30 kt is equivalent to the Italian production in one year
    - However, Argon not “used”, just “stored” in the experiment
  - ⇒ **Safety issues**
    - In critical area like LNGS lab, a monolithic volume of 1000 m<sup>3</sup> seems to be a limit ⇒ this constraint most likely relaxed in other sites

# Physics potential (I)

## 👍 Proton decay

✳️ Large variety of decay modes accessible

⇒ *study branching ratios free of systematics*

✳️ Background free searches for even for 30 years running!!!

⇒ *linear gain in sensitivity with exposure*

✳️ In case of negative results:

⇒  $\tau_p > 0$  ( $10^{34-35}$  years) in 10 years of data taking

## 👍 Atmospheric neutrinos

✳️ Observation free of experimental biases!

— *Detection down to production thresholds*

— *Complete event final state reconstruction*

— *Measurement of all neutrino flavors in all modes (CC & NC)*

✳️ Excellent resolution on L/E reconstruction

✳️ Direct  $\tau$  appearance search

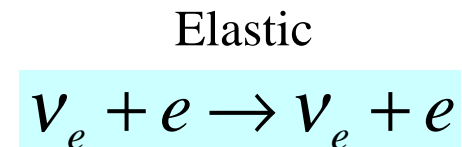
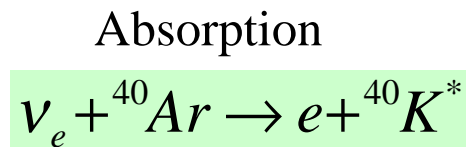
# Physics potential (II)

## 👉 Neutrinos from accelerators ( $\nu$ factory)

- \* Precise measurement of  $\Delta m^2_{23}$ ,  $\Theta_{23}$ ,  $\Theta_{13}$
- \* Matter effects, sign of  $\Delta m^2_{23}$
- \* First observation of  $\nu_e \rightarrow \nu_\tau$  (unitarity of mixing matrix)
- \* CP violation

## 👉 Solar neutrinos

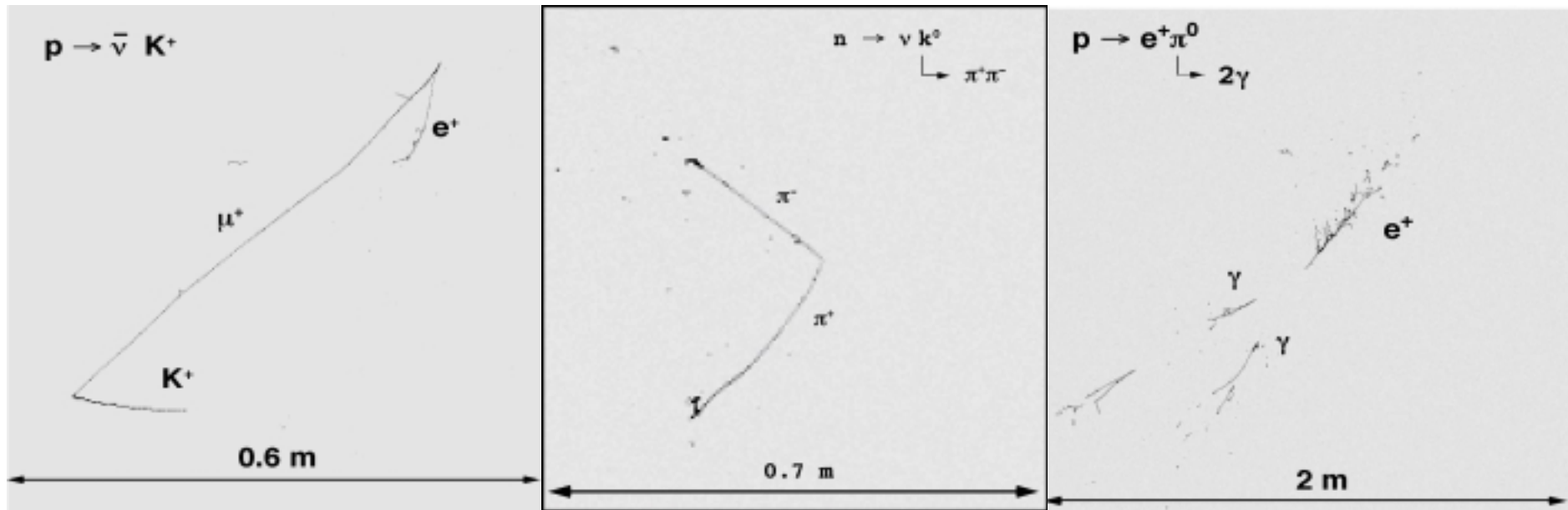
- \* Energy threshold: 5 MeV
- \* Large statistics, high precision measurements
- \* Experimental signal



## 👉 Supernovae



# Nucleon decay search



Thanks to **excellent tracking and particle *id* capabilities**

LAr unique tool for

Extremely efficient background rejection  
High detection efficiency

# $p \rightarrow e^+ \pi^0$ decay mode

**Exposure: 1000 kton x year**

Cuts	$e + \pi^0$ Argon	$e + \pi^0$ Oxygen	$\nu_e$ CC	$\bar{\nu}_e$ CC	$\nu_\mu$ CC	$\bar{\nu}_\mu$ CC	$\nu$ NC	$\bar{\nu}$ NC
Initial	100%	100%	59861	11707	106884	27273	64705	29612
One $\pi^0$	54%	70%	5277	1696	11160	4388	6223	2278
One $e$	54%	70%	5277	1696	7	< 1	< 1	< 1
$T_p < 100$ MeV	53%	68%	2505	1256	< 1	< 1	< 1	< 1
$0.8 < \text{Inv Mass} < 1.05$ GeV	38%	53%	306	204	< 1	< 1	< 1	< 1
Total Momentum < 0.25 GeV	19%	24%	1	< 1	< 1	< 1	< 1	< 1

Overall efficiency in Argon

Overall efficiency in Oxygen

**Full simulation of backgrounds**

# $p \rightarrow K^+ \bar{\nu}$ decay mode

Exposure: 1000 kton x year

Cuts	$K + \bar{\nu}$	$\nu$ NC	$\bar{\nu}$ NC
Initial	100%	64705	29612
No primary $\pi^\pm$	99.4%	55481	26033
No primary $\pi^0$	98.7%	48397	23265
Only one kaon	98.5%	108	22
Total Energy < 0.65 GeV	85%	< 1	< 1

Full simulation of backgrounds

# Limits on proton mean life ( $\tau_p$ )

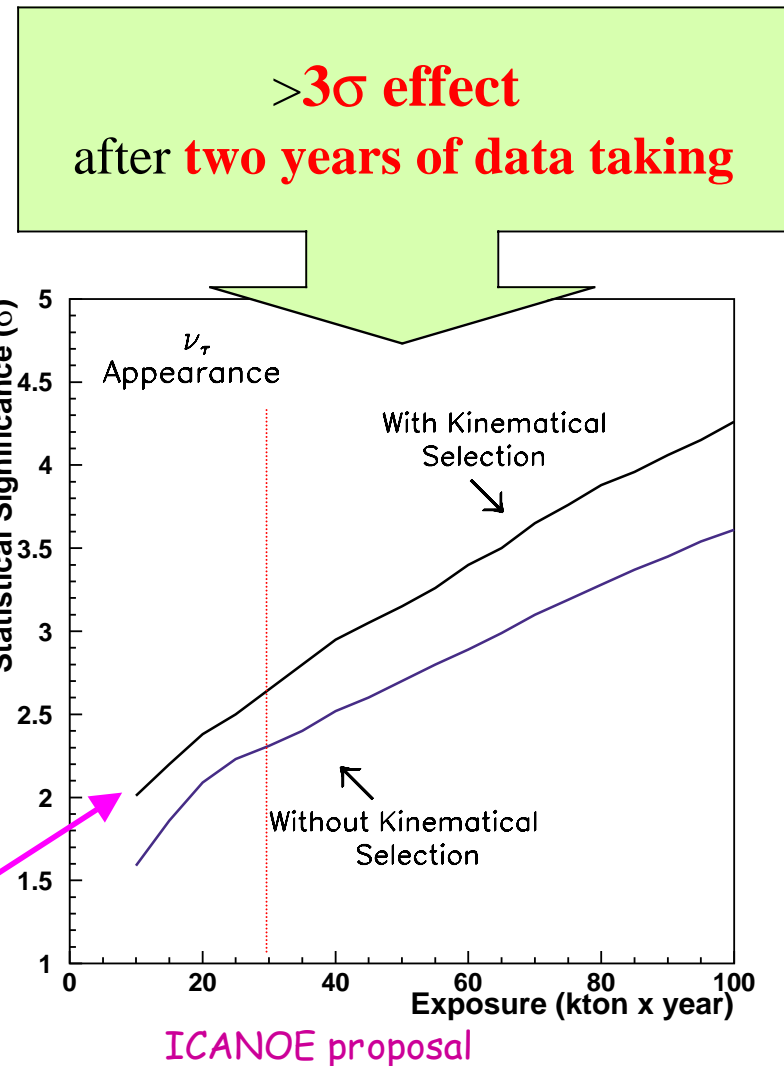
10 years @ SuperICANOE

<i>Exposure: 300 kton × year</i>				
	$p \rightarrow e^+ \pi^0$		$p \rightarrow K^+ \bar{\nu}$	
	Efficiency (%)	$\tau_p$ (years)	Efficiency (%)	$\tau_p$ (years)
No nucl. reinteractions	42	$1.5 \times 10^{34}$	85	$3.1 \times 10^{34}$
Nucl. reinteractions (FLUKA)	19	$6.8 \times 10^{33}$	85	$3.1 \times 10^{34}$
<i>Exposure: 1000 kton × year</i>				
	$p \rightarrow e^+ \pi^0$		$p \rightarrow K^+ \bar{\nu}$	
	Efficiency (%)	$\tau_p$ (years)	Efficiency (%)	$\tau_p$ (years)
No nucl. reinteractions	42	$5.0 \times 10^{34}$	85	$1.0 \times 10^{35}$
Nucl. reinteractions (FLUKA)	19	$2.3 \times 10^{34}$	85	$1.0 \times 10^{35}$

30 years @ SuperICANOE

# Atmospheric direct $\tau$ appearance

- Search for  $\nu_\tau CC$  at high energy ( $\tau$  threshold)
  - $-30 \nu_\tau CC/\text{year}$  expected
- Compare NC(top) to NC(bottom) at high energy
  - Requires good discrimination of NC event direction
- Exploit precise measurement of all final state particles
  - Count events as a function of visible energy
  - Improved discrimination by a study of the event kinematical properties

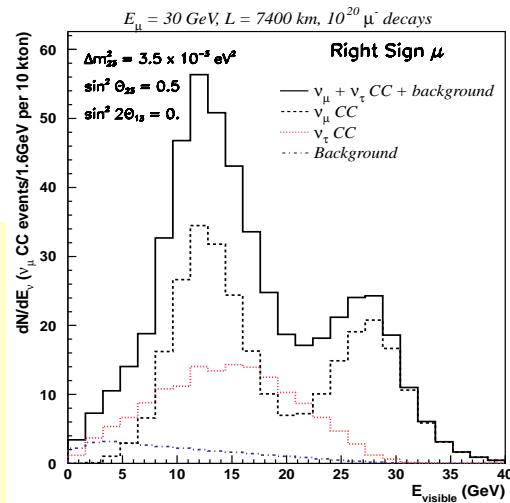
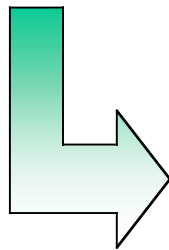


# Event classes at a $\nu$ factory

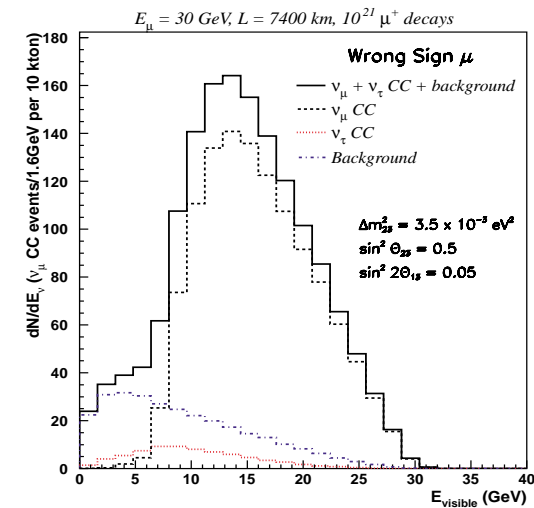
Ideal detector able to measure **12 different oscillation processes**

A. Bueno, M. Campanelli, A. Rubbia, hep-ph/0005007

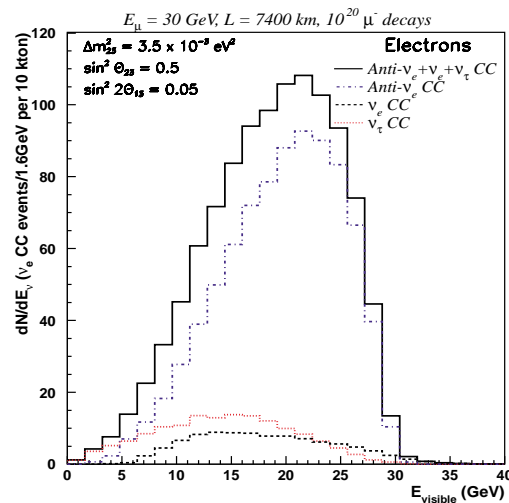
**LAr imaging  
allows precise  
identification of  
all oscillation  
processes**



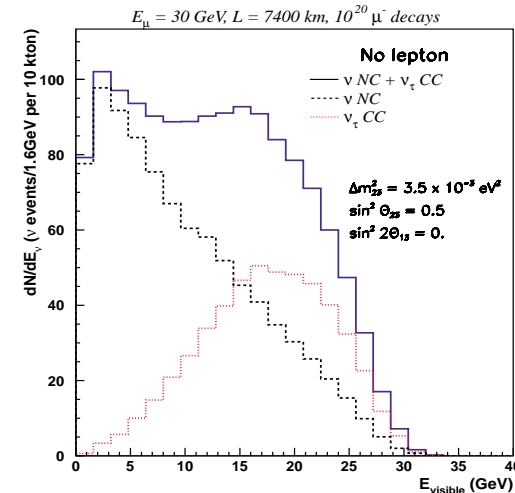
**Right sign muons**



**Wrong sign muons**



**Electrons**

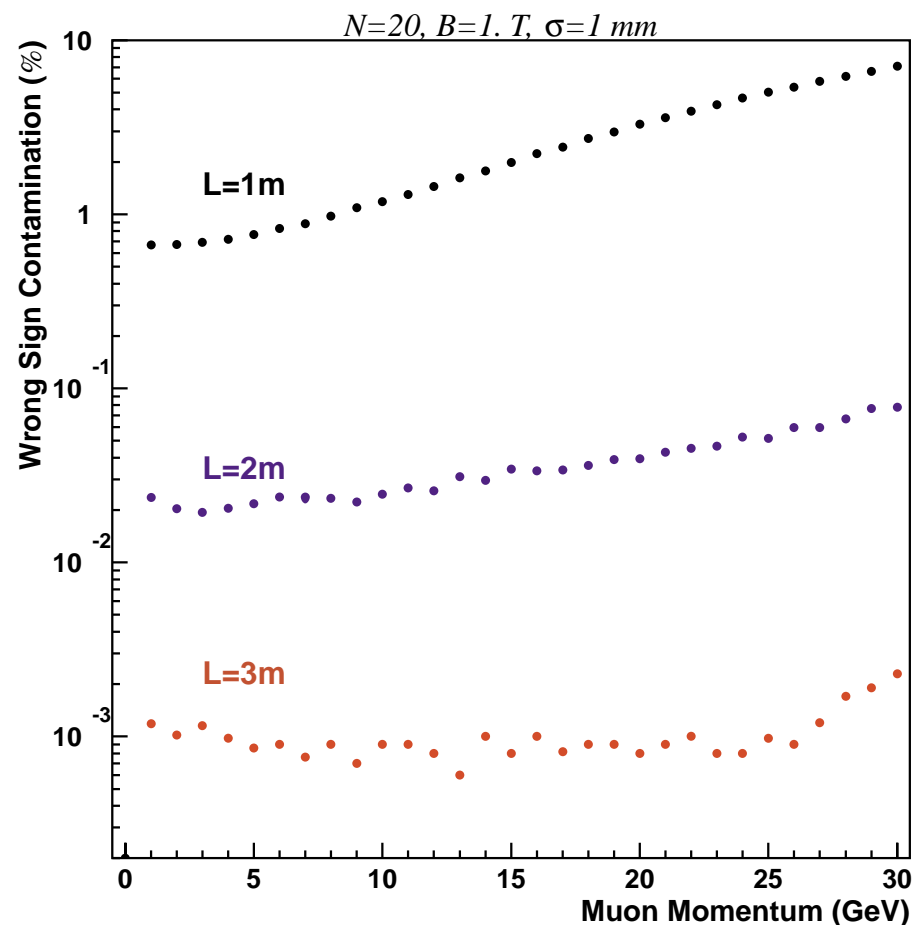


**No lepton**



# Muon identification

- $\mu$  momentum resolution:
  - ⇒ **20% for a 3m long Fe spectrometer with B=1T**
- Wrong sign contamination
  - ⇒ **Charge confusion:  $\sim 10^{-5}$**
- Large detection efficiency for low energy beam
  - ⇒  **$\mu$  detection threshold ( $dE/dx = 240 \text{ MeV/m}$ )**



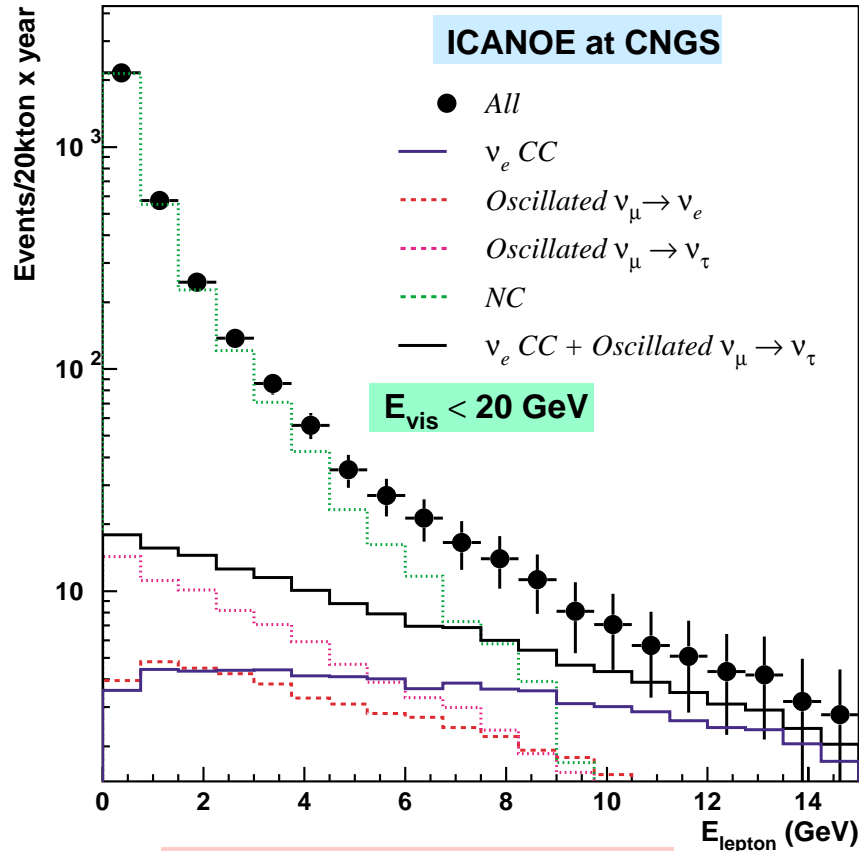
# $\nu_e$ CC versus $\nu$ NC discrimination (I)

NC

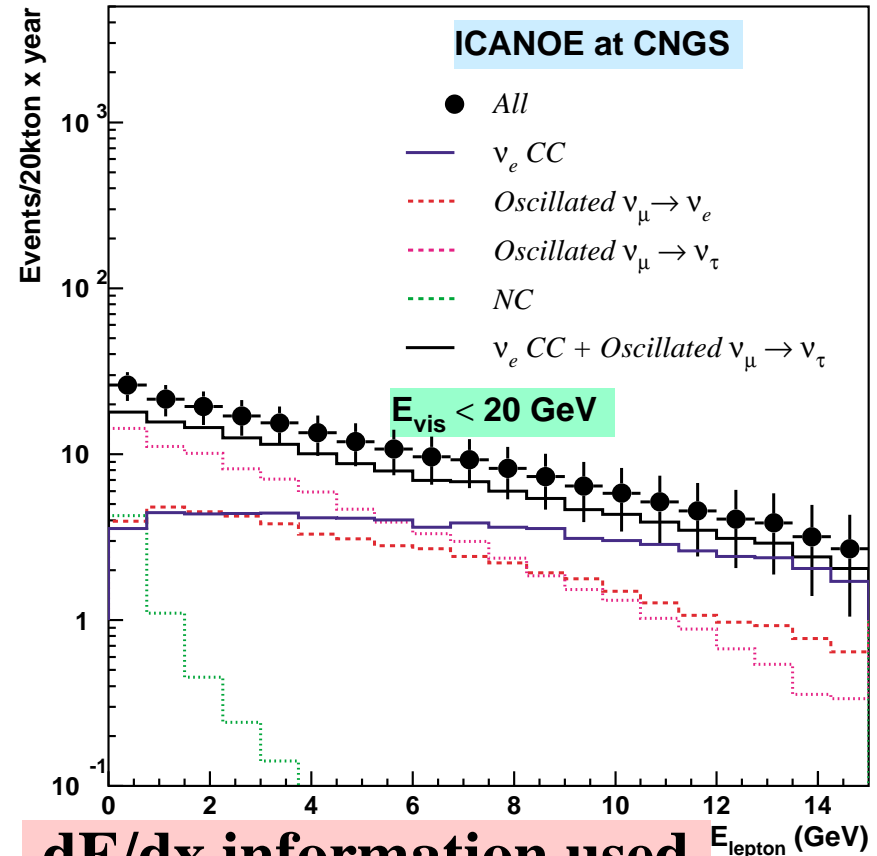
$\gamma$  converting within 3 cm (10 samples) from primary vertex considered as electron candidate

$$L = 732 \text{ Km } \Delta m^2 = 3.5 \times 10^{-3} \text{ eV}^2 \quad \Theta_{23} = 45^\circ \quad \Theta_{13} = 7^\circ$$

$$L = 732 \text{ Km } \Delta m^2 = 3.5 \times 10^{-3} \text{ eV}^2 \quad \Theta_{23} = 45^\circ \quad \Theta_{13} = 7^\circ$$



**NO dE/dx  
information used**



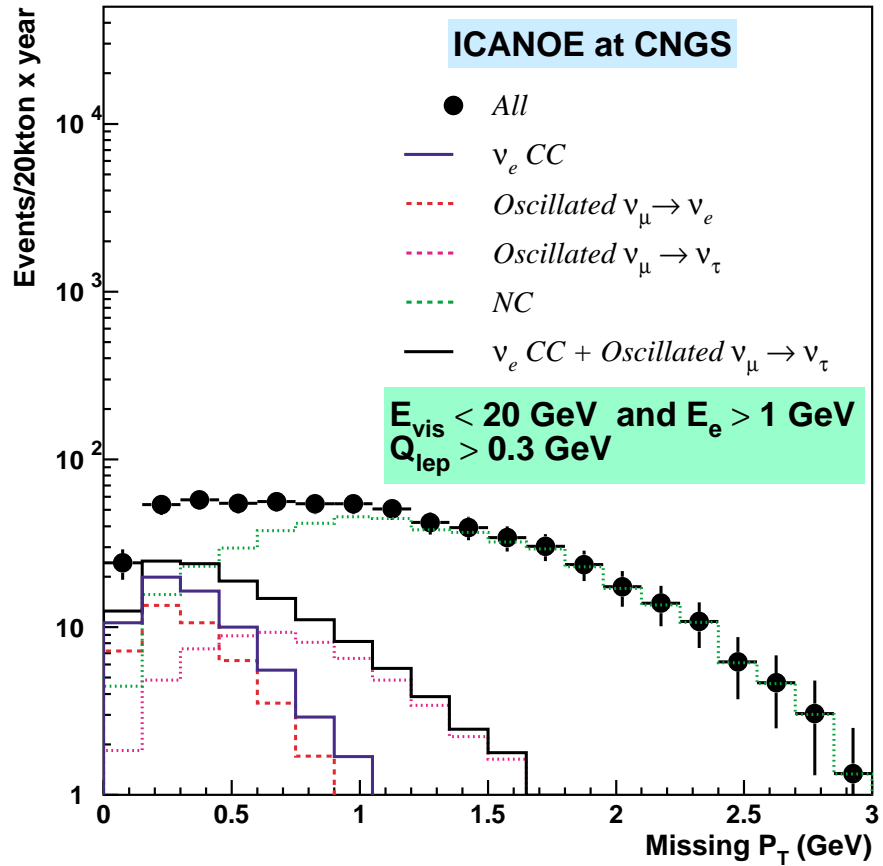
**dE/dx information used**  
**Single vs double m.i.p.**  
**algorithm provides >500**  
**rejection factor**

# $\nu_e$ CC versus $\nu$ NC discrimination (II)

**NC rejection** 

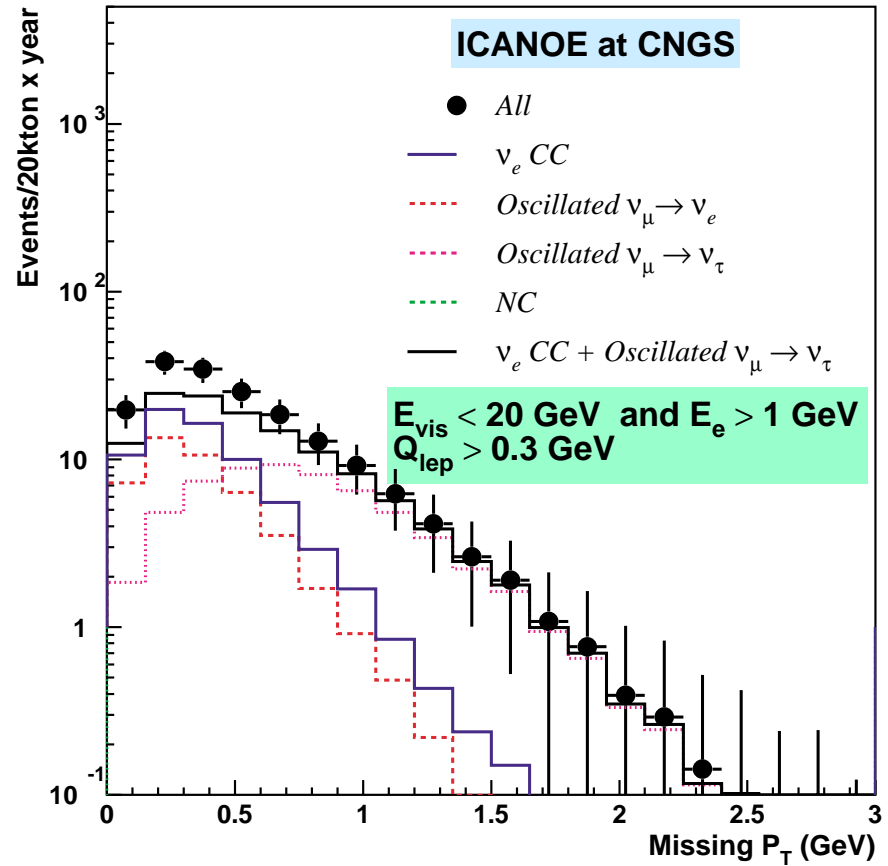
Additional discrimination power provided by event kinematics

$L = 732 \text{ Km } \Delta m^2 = 3.5 \times 10^{-3} \text{ eV}^2 \Theta_{23} = 45^\circ \Theta_{13} = 7^\circ$



**NO dE/dx information used**

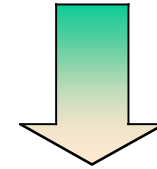
$L = 732 \text{ Km } \Delta m^2 = 3.5 \times 10^{-3} \text{ eV}^2 \Theta_{23} = 45^\circ \Theta_{13} = 7^\circ$



**dE/dx information used**

# Solar neutrinos

Rates for  
**34ktons**  
and  $E_e > 5 \text{ MeV}$

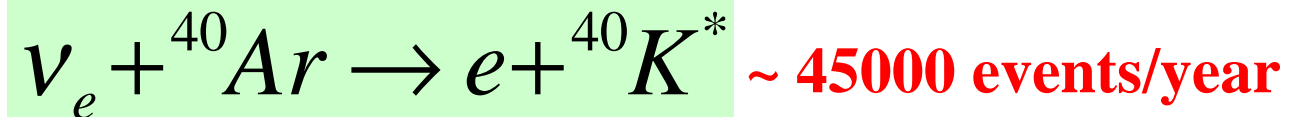


Elastic channel

$$\nu_e + e \rightarrow \nu_e + e \sim 10000 \text{ events/year}$$

Signal

Neutrino absorption



Super allowed process  $\Rightarrow$  larger cross section

Large statistics, high precision measurements

# Conclusions

- Liquid Argon imaging technology allows
  - ⇒ **Very high granularity**
  - ⇒ **Very large mass detectors**
  - ⇒ **Bubble-chamber-like detector**
- It has been *successfully tested* on large prototypes (50l, 3T, 15Ton prototype)
  - ⇒ **Has now entered the fully industrialized era (T600)**
- Timescale:
  - ⇒ **T600 first cool-down planned before end 2000**
  - ⇒ **ICANOE proposal been discussed @ CNGS**
- *The road to SuperICANOE is open...*