

Neutrino factories^(): Detector concepts*

**we do not consider conventional “superbeams” (see hep-ph/0103052)*



*André Rubbia, ETH Zürich
(ICARUS Collaboration)*

Special thanks to A. Bueno & M. Campanelli

***IX International Workshop on
"Neutrino Telescopes"***

6th-10th March, 2001

The oscillation physics program at the NF

$$\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$$

$$\nu_\mu \rightarrow \nu_e$$

appearance

$$\nu_\mu$$

disappearance

$$\nu_\mu \rightarrow \nu_\tau$$

appearance

$$\bar{\nu}_e$$

disappearance

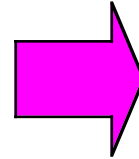
$$\bar{\nu}_e \rightarrow \bar{\nu}_\mu$$

appearance

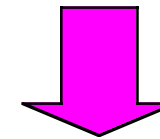
$$\bar{\nu}_e \rightarrow \bar{\nu}_\tau$$

appearance

Plus their charge conjugates with μ^+ beam



Ideal detector should be able to measure **12 different processes as a function of L and E_ν**



$$\begin{cases} \nu_\ell N \rightarrow \ell^- + hadrons \\ \bar{\nu}_\ell N \rightarrow \ell^+ + hadrons \end{cases}$$

$$\begin{cases} \nu_\ell N \rightarrow \nu_\ell + hadrons \\ \bar{\nu}_\ell N \rightarrow \bar{\nu}_\ell + hadrons \end{cases}$$

1. **Particle ID**: charged lepton tags *incoming neutrino flavor*

2. **Charge ID**: sign of lepton charge tags *helicity* of incoming neutrino

3. **Energy resolution**: Reconstructed event energy is $E_\nu = E_\ell + E_{had}$

4. **Various baselines L** could help for detector systematics

The Neutrino Factory



Roughly as many ν_e 's as ν_μ 's

P. Lipari, hep-ph/0102046

Flux
scales as
 E_μ^2/L^2

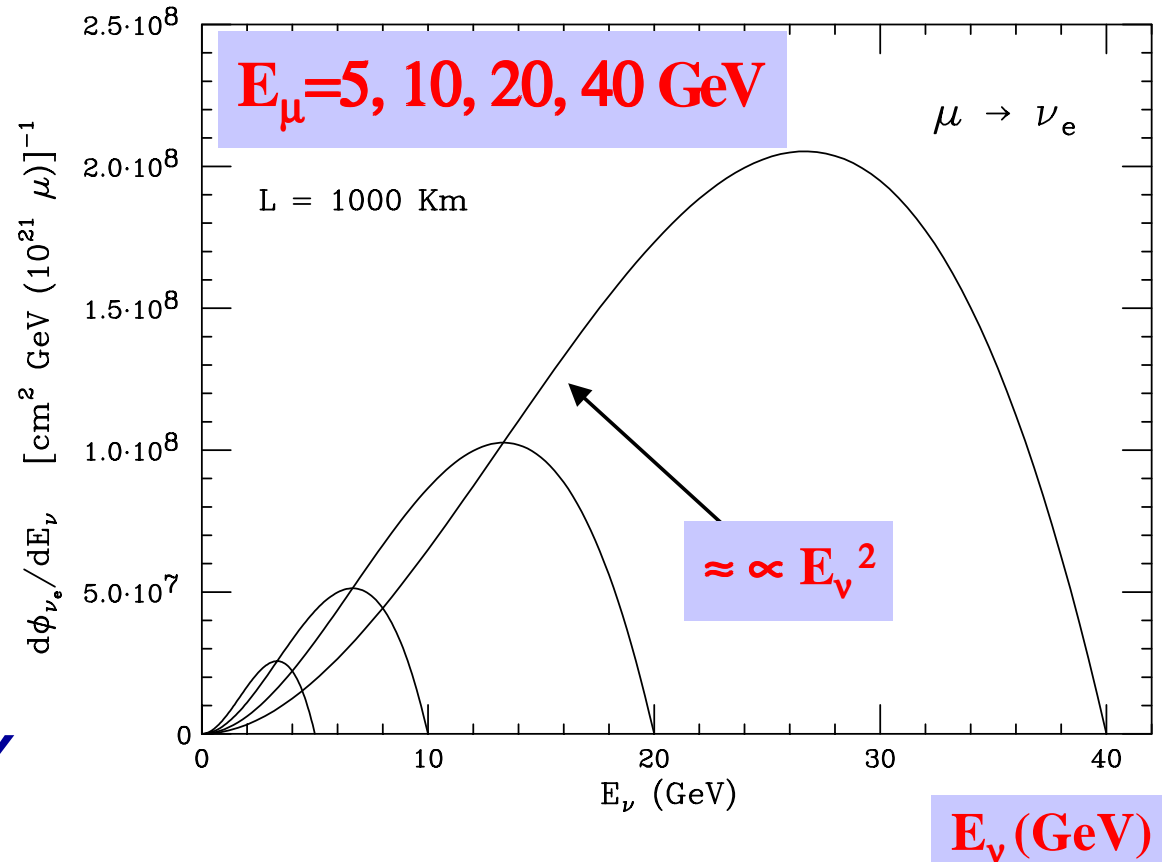
Total event rate
scales as
 $\approx E_\mu^3/L^2$

A very important feature!

The generally adopted
consensus (see later) \Rightarrow

high energy ν -factory

$E_\mu = O(30 \text{ GeV})$



$$\frac{dN}{dx} \propto x^2(1-x) \quad x \equiv E_\nu / E_\mu$$

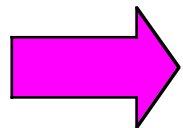
Predicted event rates at a Neutrino Factory

FNAL-FN-692, Apr 2000

10^{20} μ^- decays

No oscillations assumed

Experiment		Baseline (km)	$\langle E_{\nu_\mu} \rangle$ (GeV)	$\langle E_{\bar{\nu}_e} \rangle$ (GeV)	N(ν_μ CC) (per kt-yr)	N($\bar{\nu}_e$ CC) (per kt-yr)
NuMI	Low energy	732	3	—	458	1.3
	Medium energy	732	6	—	1439	0.9
	High energy	732	12	—	3207	0.9
CNGS		732	17	—	2714	1.4
Muon ring	E_μ (GeV)					
	10	732	7.5	6.6	1400	620
	20	732	15	13	12000	5000
	50	732	38	33	1.8×10^5	7.7×10^4
Muon ring	E_μ (GeV)					
	10	2900	7.6	6.5	91	41
	20	2900	15	13	740	330
	50	2900	38	33	11000	4900
Muon ring	E_μ (GeV)					
	10	7300	7.5	6.4	14	6
	20	7300	15	13	110	51
	50	7300	38	33	1900	770



However, in addition to the increased neutrino flux, ambitious oscillation physics program requires detectors in the **10's kton** range to perform experiment with baselines **$L \approx 1000$'s km**

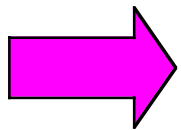
The goal: detect μ^+ , μ^- , e^+ , e^- , τ^+ , τ^- and NC !

★ Particle ID: \Rightarrow *via CC interactions*

- **Muons:** *straight-forward*, look for penetrating particles, but beware π^\pm, K^\pm and charm decays
- **Electrons:** *harder*, look for large & “short” energy deposition, need good granularity for e/π^0 separation
- **Taus:** *hardest*, “kink” or kinematical methods (statistical separation), $\tau \rightarrow \text{hadrons} + \nu$ (Br \approx 60%) look like “NC”

★ Charge ID: \Rightarrow *via magnetic analysis*

- **Muons:** *easy*, muon spectrometer downstream or fully magnetized target
- **Electrons:** *hardest*, need to measure significantly precisely the bending in B-field before start of e.m. shower
- **Taus:** easy for $\tau \rightarrow \mu\nu\nu$ (Br \approx 18%), otherwise *difficult*



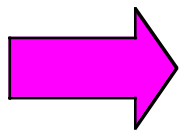
This has to be implemented on multi-kton detectors... various choices & optimizations considered.

1a. Magnetized steel-scintillator sandwich

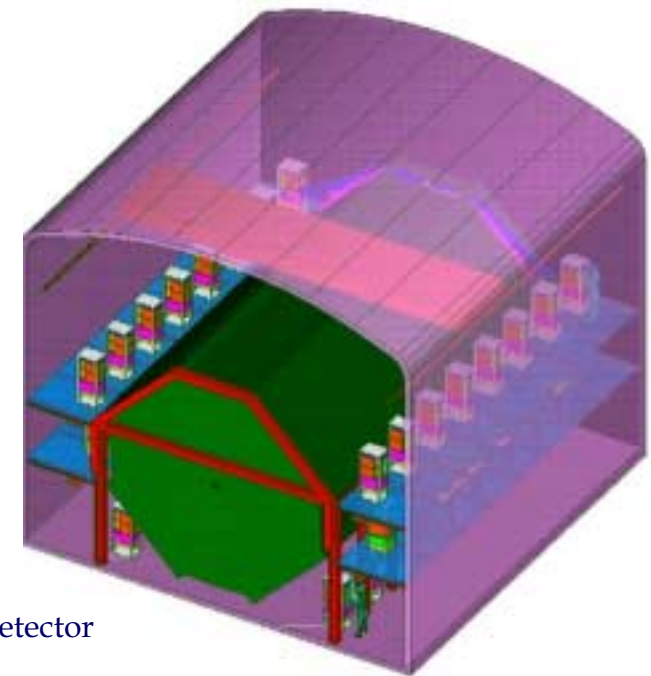
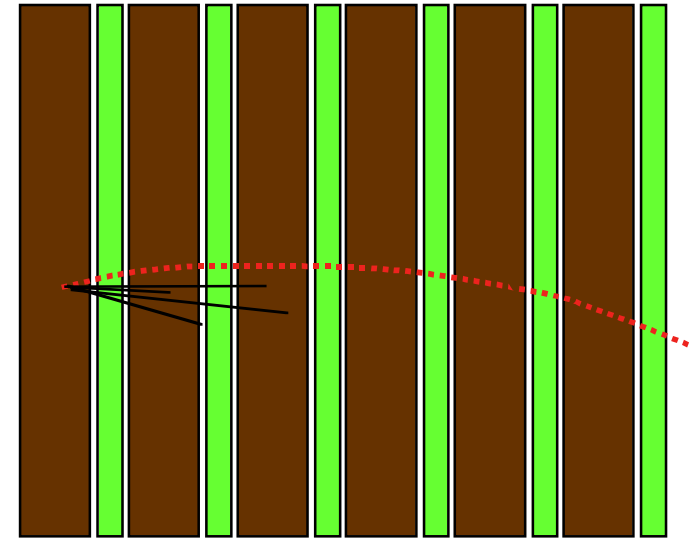
- ★ *High density*
- ★ *Magnetizable* for μ^+/μ^- discrimination
 - Detects: $\mu^+, \mu^-, [e, NC]$
- ★ Good muon and reasonable jet energy resolution $\sigma_{E_h}/E_h \sim 80\%/\sqrt{E_h}$
- ★ *Lots of experience*. e.g. CCFR/NuTeV, CDHS.
MINOS will reach 5.4 kton in 2003.
- ★ Disadvantages:
 - Amount of instrumentation scales with *volume*
 - *Minimum muon energy threshold* (4-6 GeV) in order to separate it from hadrons and muon isolation from jet difficult to tell. Threshold even higher to have excellent wrong-sign- μ @ 10^{-5}
 - *e/h discrimination* rather poor
 - Angular resolution determined by transverse readout segmentation, usually rather modest.

The MINOS far detector

- ★ 8m octagonal tracking calorimeter
- ★ 486 layers of 2.45cm Fe
- ★ 2 sections, each 15m long
- ★ 4.1cm wide solid scintillator strips with WLS fiber readout
- ★ 25,800 m² active detector planes
- ★ Magnet coil $\langle B \rangle = 1.3\text{T}$
- ★ 5.4kton total mass



Can it be scaled to 40 kton ?



Half MINOS detector

MINOS far detector parameters

System	Parameters
MINOS cavern	82.3 m × 13.8 m × 11.6 m (height)
Supermodules	2 supermodules, each 2.7 metric kt, 14.4 m long × 8 m wide
Detector mass	5.14 ktons steel + 261 tons scintillator = 5.4 ktons
Planes/supermodule	486 steel planes and 485 scintillator planes, 2.54 cm pitch
Detector units/plane	192 scintillator strips packaged in 8 modules
Readout	2-ended, with 8 × multiplexing
Channel count	484 planes × 192 strips × 2 ÷ 8 = 23,232 channels
Photodetectors	1,452 16-channel PMTs in 484 MUX boxes
Installation rate	1 plane/1.85 shifts or 24 planes/month (maximum)
Installation time	12 months for first supermodule, 22.5 months for two
Magnetic field	1.3 T at 2 m radius in steel octagon planes
Magnet coils	15 kA-turns, water-cooled copper wire, 58 kW total
Total cavern cooling	292 kW maximum (at the end of the installation period)

Table 13: Summary of some of the major parameters of the far detector and its requirements on the infrastructure systems of the MINOS cavern in the Soudan mine.

Lifting of an instrumented MINOS plane



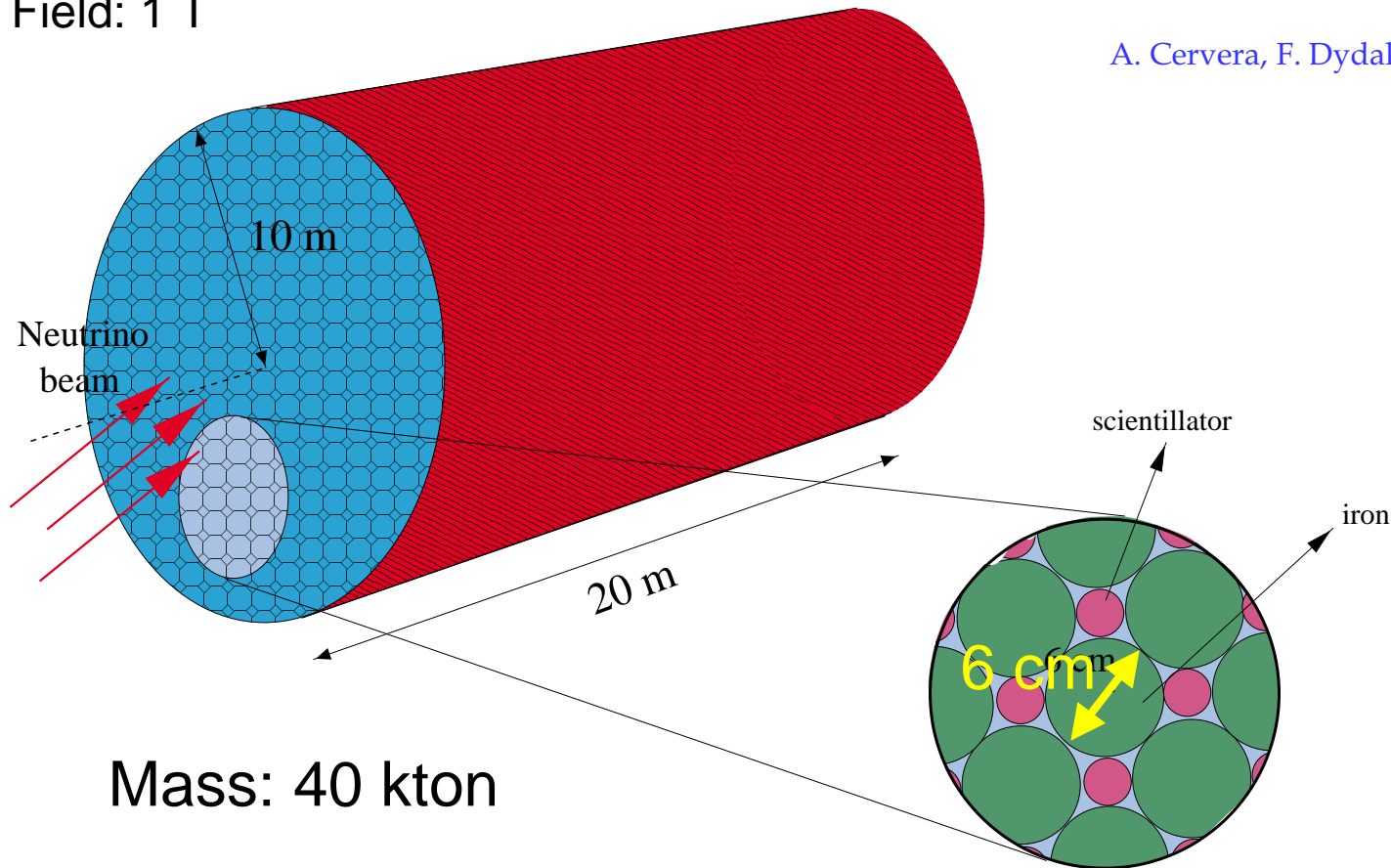
1b. Large Magnetized Detector

Cylindrical symmetry, $R=10\text{m}, L=20\text{m}$

6 cm thick iron rods interspersed with 2cm thick scintillators “longitudinally segmented”

Field: 1 T

A. Cervera, F. Dydak, J.J. Gomez. , NIM.A451 (2000) 123



Mass: 40 kton

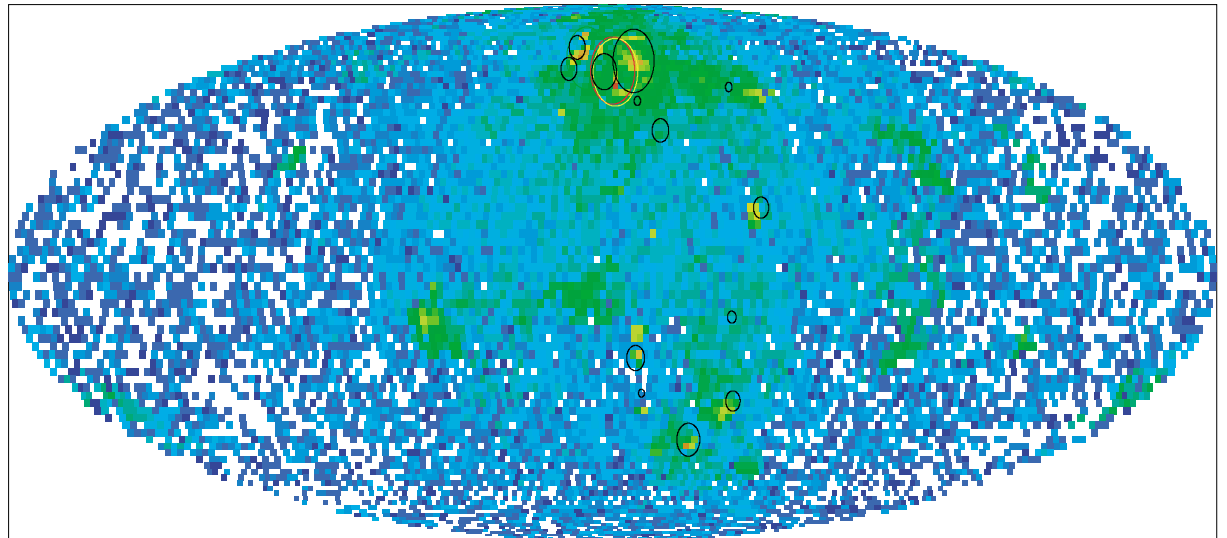
⇒ Muon performance studied with GEANT and assume MINOS-like performance for E_{had} θ_{had} ...

$$\begin{aligned} \sigma_{E_h}/E_h &\sim 76\%/\sqrt{E_h} \oplus 3\% \\ \sigma_{\theta_h}/\theta_h &\sim 17/\sqrt{E_h} \oplus 12/E_h \end{aligned}$$

2. Large Water Cerenkov

- ★ Well proven technology, e.g. *Kam*, *IMB*, *SuperKamiokande*
 - Detects: μ , e , [NC]
- ★ low cost target material, only the surface (not volume) needs to be instrumented, but size eventually limited by water properties
- ★ Next generation 1Mton under consideration C.K. Jung, hep-ex/0005046
- ★ Disadvantages:
 - Reconstructed (pattern) limited to “simple event topologies” (e.g. single-ring)
 - Not compatible with precise reconstruction of high energy neutrinos (e.g. $E_\nu > \approx 5 \text{ GeV}$?)
 - Muon charge ID

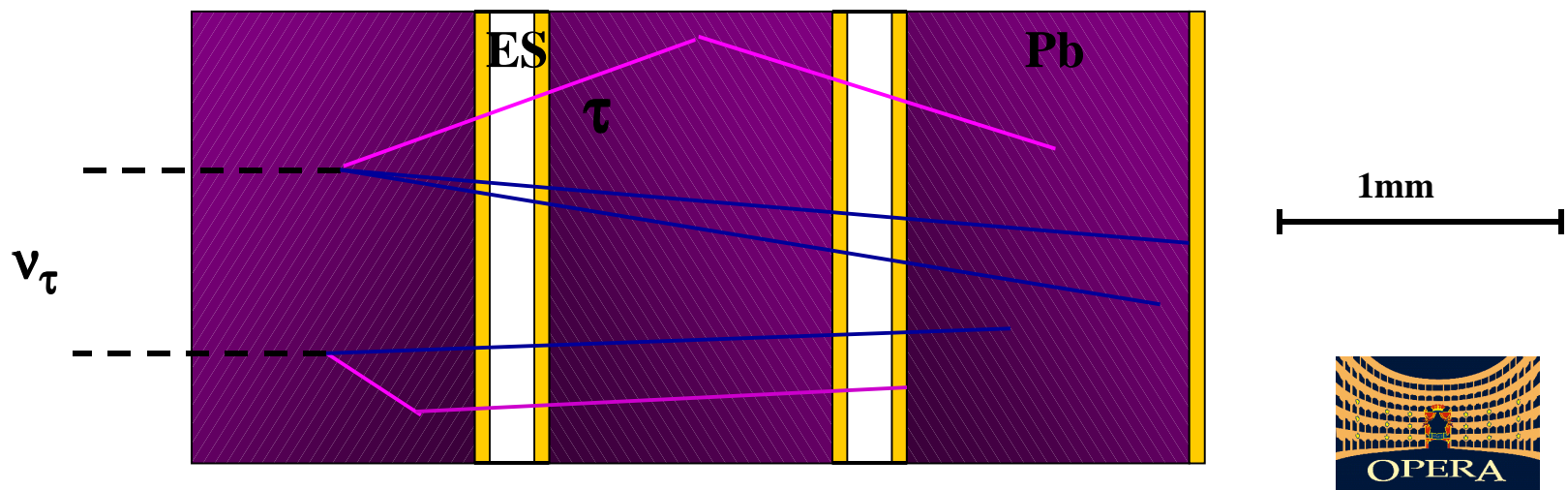
Simulated
neutrino event
from a 50 GeV
storage ring



FNAL-FN-692, Apr 2000

3. Emulsion/target sandwich

- ★ **OPERA**: Technique on the ≈ 1 kton scale to be demonstrated by at LNGS in 2005
 - Detects: μ^+ , μ^- , τ (“kink”), [e,NC]
 - Direct search of $\nu_e \rightarrow \nu_\tau$ oscillations, if charge of the tau can be detected to suppress $\nu_\mu \rightarrow \nu_\tau$ “background”
- ★ Disadvantages:
 - Probably difficult to scale to multi-kton mass
 - Scanning (e.g. $\nu_e \rightarrow \nu_\tau$ to $\sin^2 2\theta_{13} \approx 10^{-4}$?)
 - Possibly severe charm background from ν_e interactions

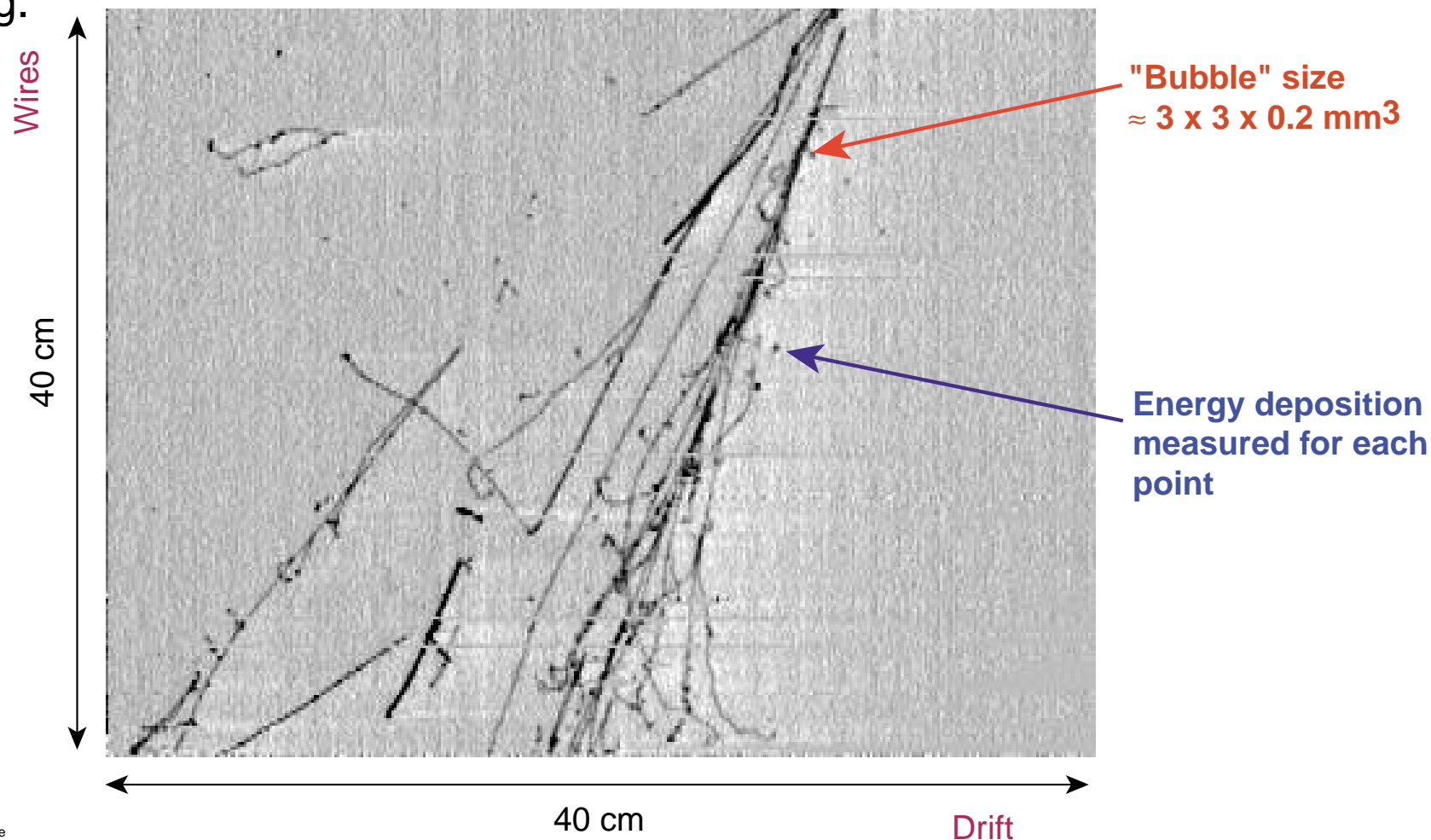


4. Liquid Argon imaging TPC

- ★ **ICARUS**: mature technique, demonstrated up to 15 ton prototype
- ★ Features provided:
 - **Detects: μ^+ , μ^- , e, NC, [τ]**
 - Fully homogeneous, continuous, precise **tracking device** with high resolution dE/dx **measurement** and full sampling **electromagnetic and hadronic calorimetry**
 - Excellent **e identification/measurement** and **e/hadron** separation
 - Very good hadronic energy resolution
- ★ 600 ton prototype construction **very advanced**
 - After the foreseen series of technical tests to be performed in Pavia within the summer 2001, **the T600 module will be ready to be transported into the LNGS tunnel**
- ★ Disadvantages:
 - Muon charge discrimination: target cannot be easily magnetized (but...)
 - Rely on **down-stream muon spectrometer** (low threshold since dE/dx \approx 240 MeV/m)
Idea first implemented in the ICANOE proposal

Liquid Argon technology

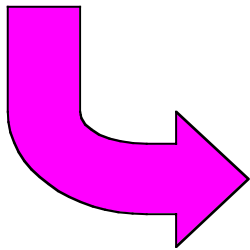
- ★ The LAr TPC technique is based on the fact that ionization electrons can drift over large distances (meters) in a volume of purified liquid Argon under a strong electric field. If a proper readout system is realized (i.e. a set of fine pitch wire grids) it is possible to realize a massive "electronic bubble chamber", with superb 3-D imaging.



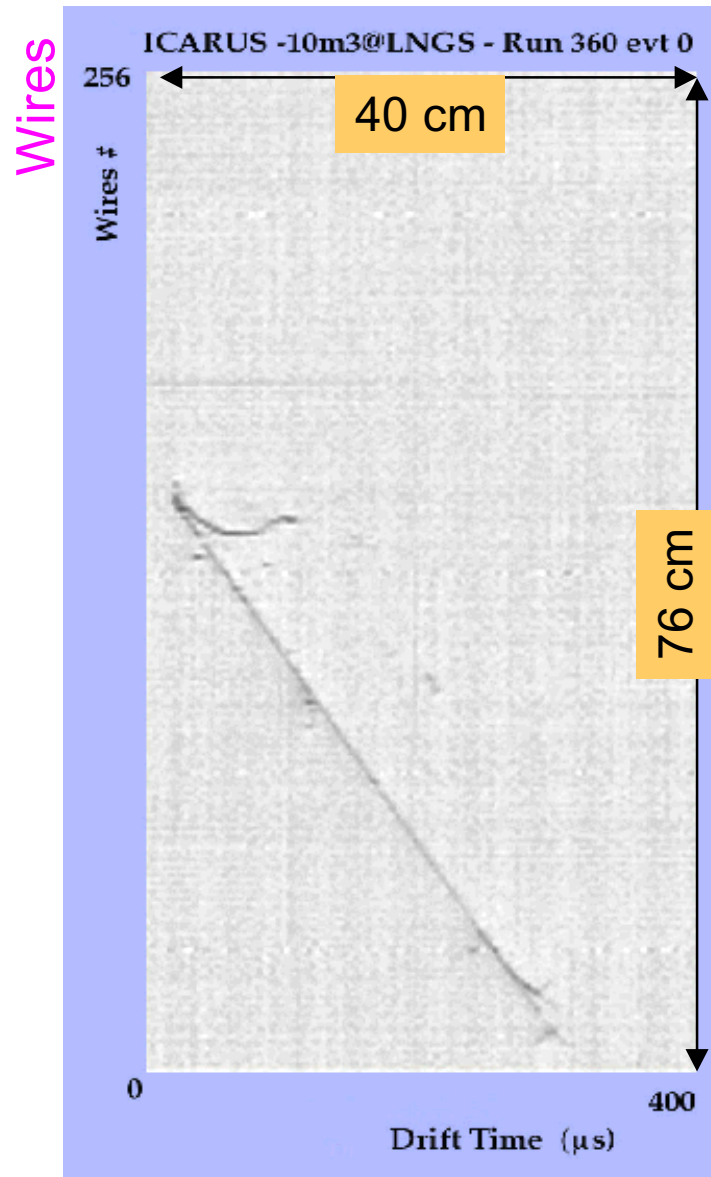
Liquid Argon imaging on large scales

10m³ Module
at LNGS

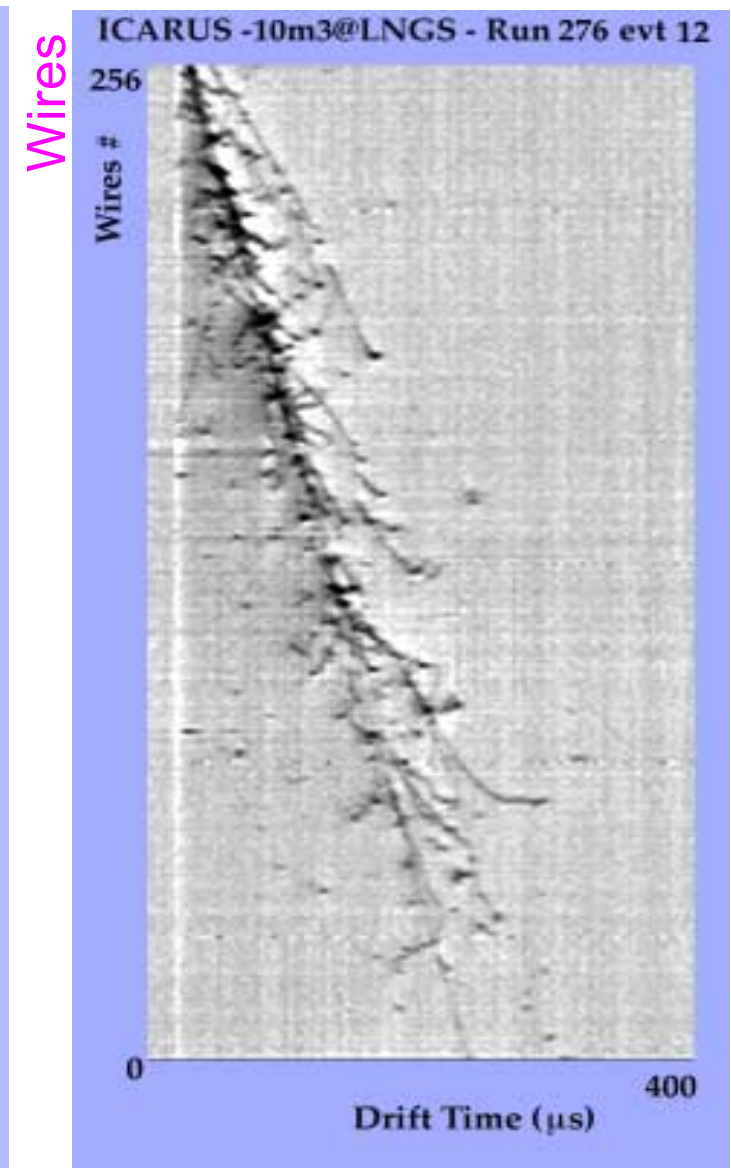
Cosmic Ray tracks
recorded during the
10 m³ operation



*“Big track” in
T600 semimodule
expected soon...*



Drift



Drift

T600 assembly schedule

- ★ Completed *site preparation* in Pavia for the T600 cryostat (**Nov 1999**)
 - “clean room”, “assembly island”, floor, ...
- ★ Delivery of the *1st cryostat* by AirLiquide (**Feb 2000**)
 - Successful vacuum tightness and mechanical stress tests
- ★ Beginning of *assembly of the internal detector mechanics* (**Mar 2000**)
- ★ Completion of assembly and positioning of inner detector frames (**Jul 2000**)
- ★ Installation of *30000 wires + signal cables* (**Jul 2000-Oct 2000**)
- ★ Delivery of the *2nd cryostat* of AirLiquide (**Aug 2000**)
 - Successful vacuum tightness and mechanical stress tests
- ★ Installation of *scintillation light* and all *slow control devices* (**Jul 2000-Dec 2000**)
- ★ *H.V. and field electrodes system* installation (**Oct 2000- Jan 2001**)
- ★ Installation of the *48 electronic racks* on top of dewar (**Dec 2000-ongoing**)
- ★ Installation of *external heat insulation* (for both dewars) and *LAr and LN₂ cryogenic circuits* (**Dec 2000-Jan 2001**)
- ★ Semi-module now ready to be sealed.

The ICARUS T600 module

Under construction

Number of independent containers = 2

Single container Internal Dimensions: Length = 19.6 m , Width = 3.9 m , Height = 4.2 m

Total (cold) Internal Volume = 534 m³

Sensitive LAr mass = 476 ton

Number of wires chambers = 4

Readout planes / chamber = 3 at 0° , ± 60° from horizontal

Maximum drift = 1.5 m

Operating field = 500 V / cm

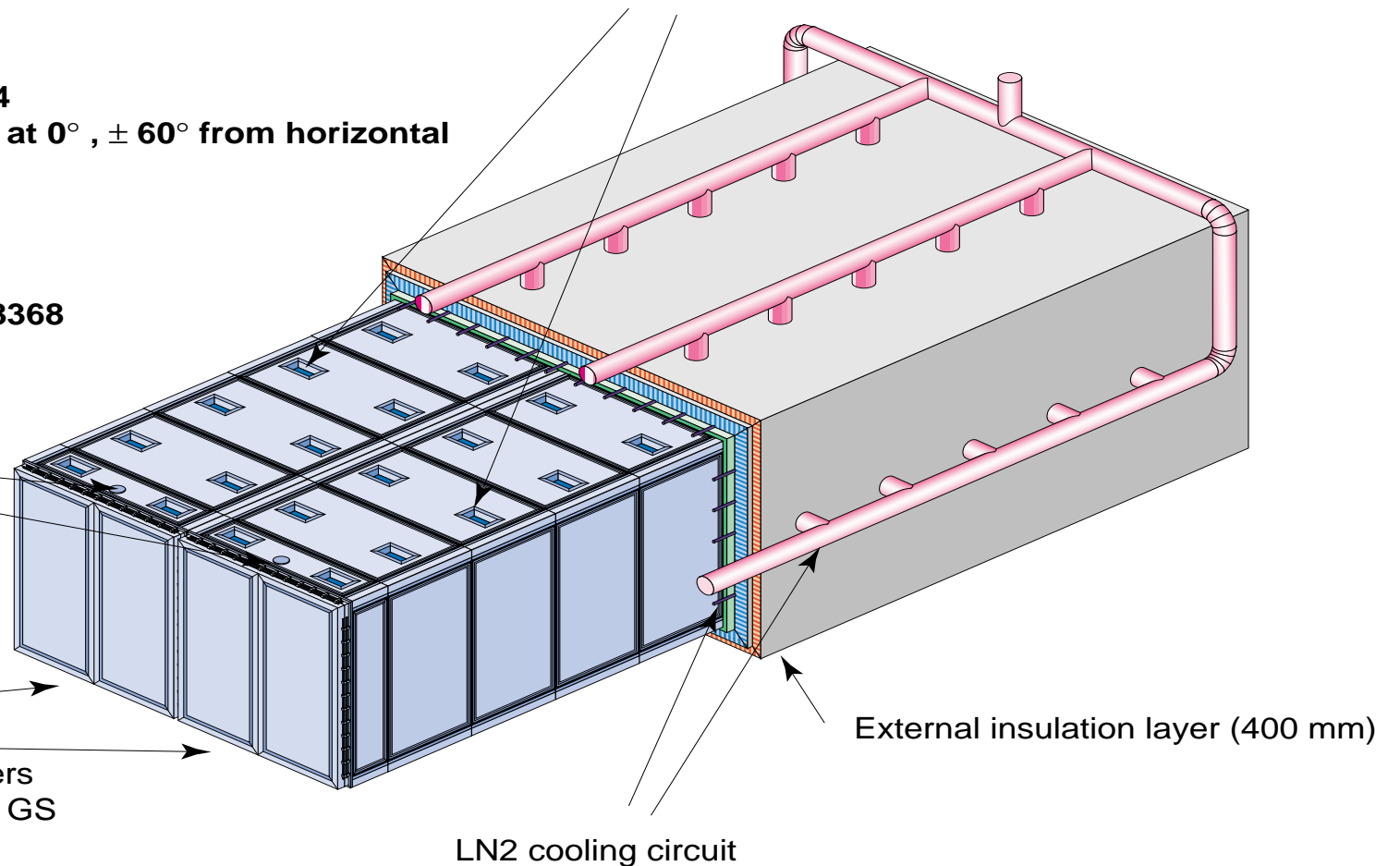
Maximum drift time ≈ 1 ms

Wires pitch = 3 mm

Total number of channels = 58368

HV feedthroughs

Signal feedthroughs

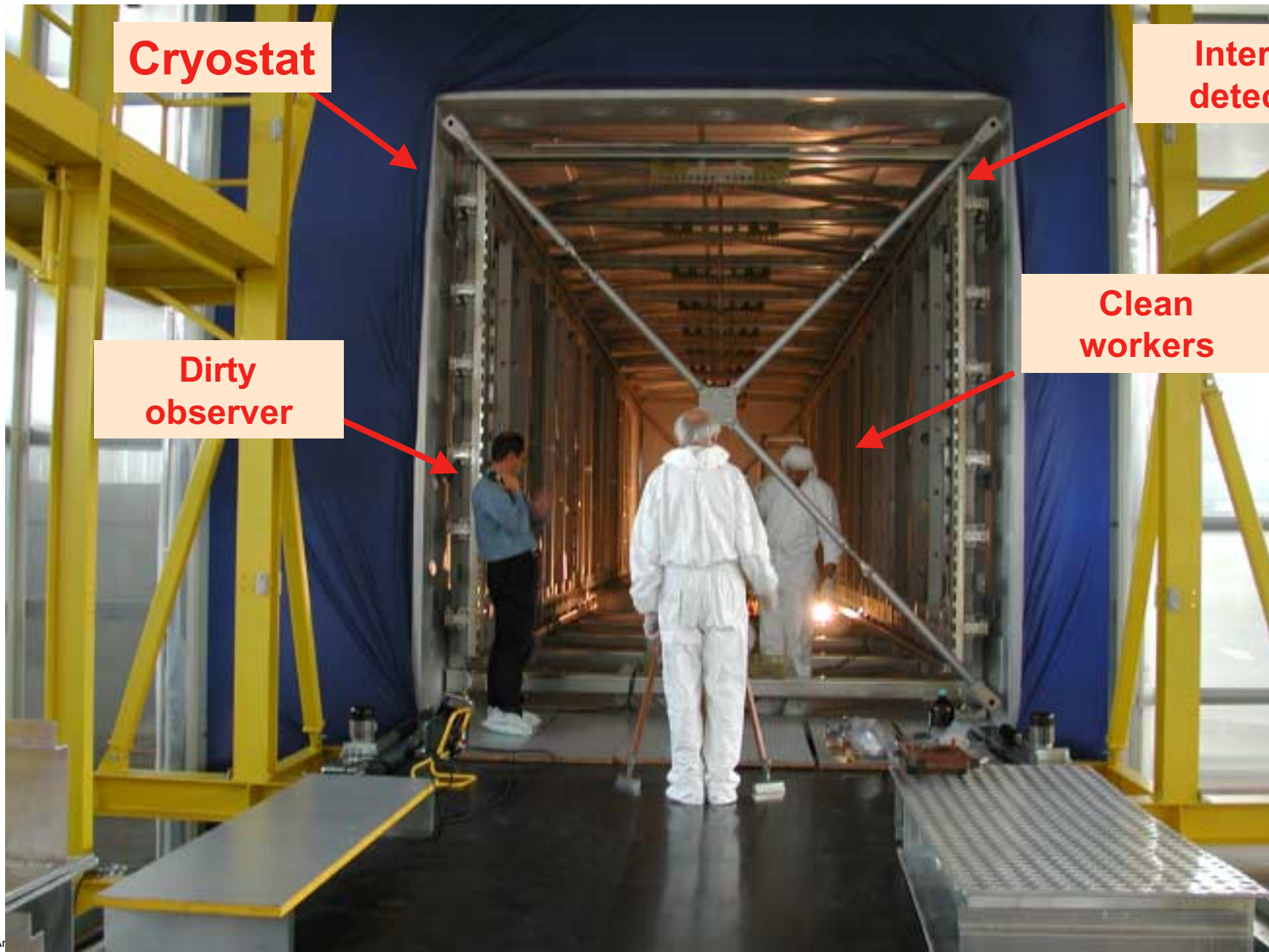


2 independent aluminum containers
each one transportable inside the GS
Laboratory

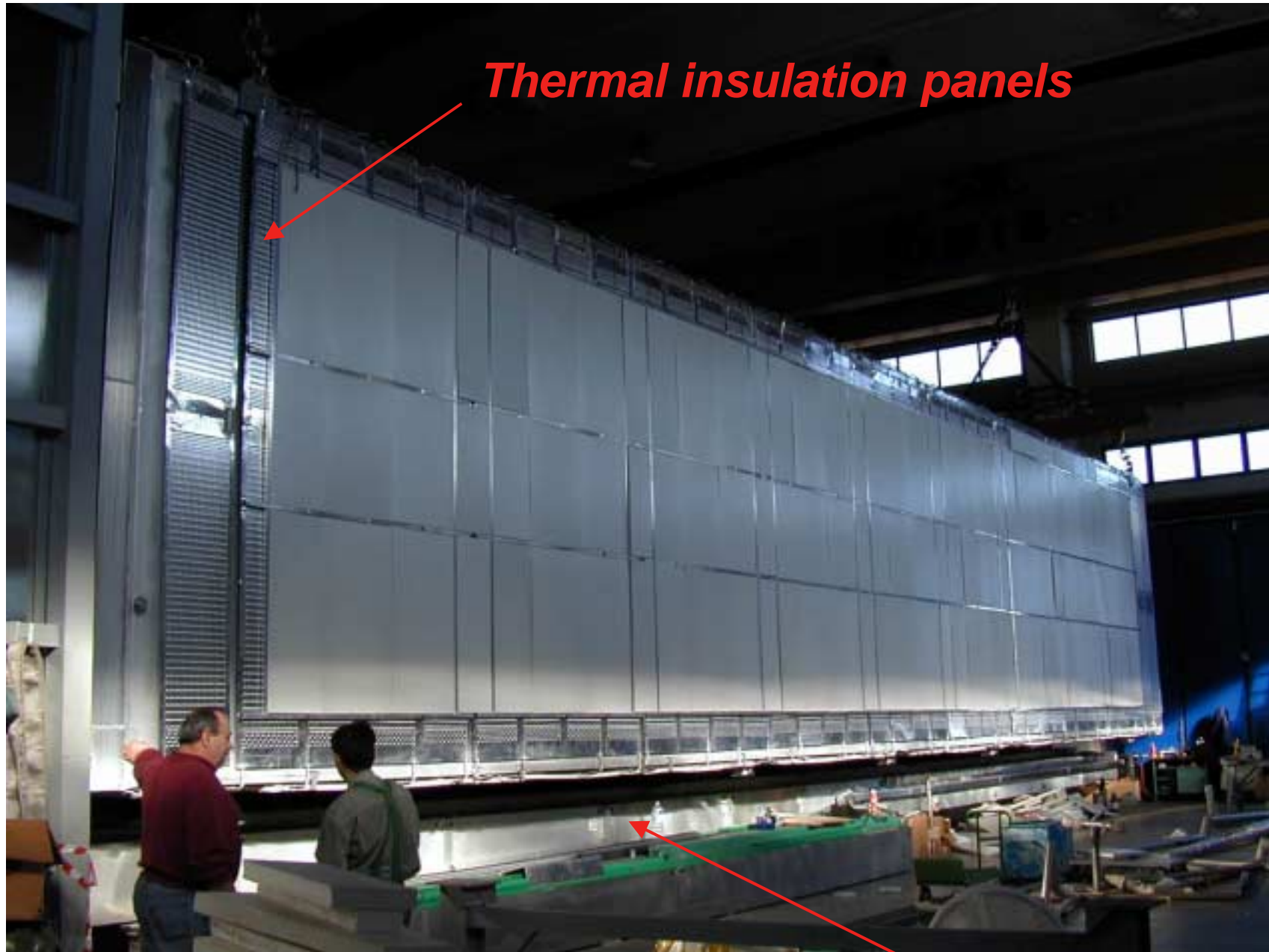
First half-module delivery in Pavia (Feb 29, 2000)



Assembly of the T600 internal detector (Mar-Jul 2000)

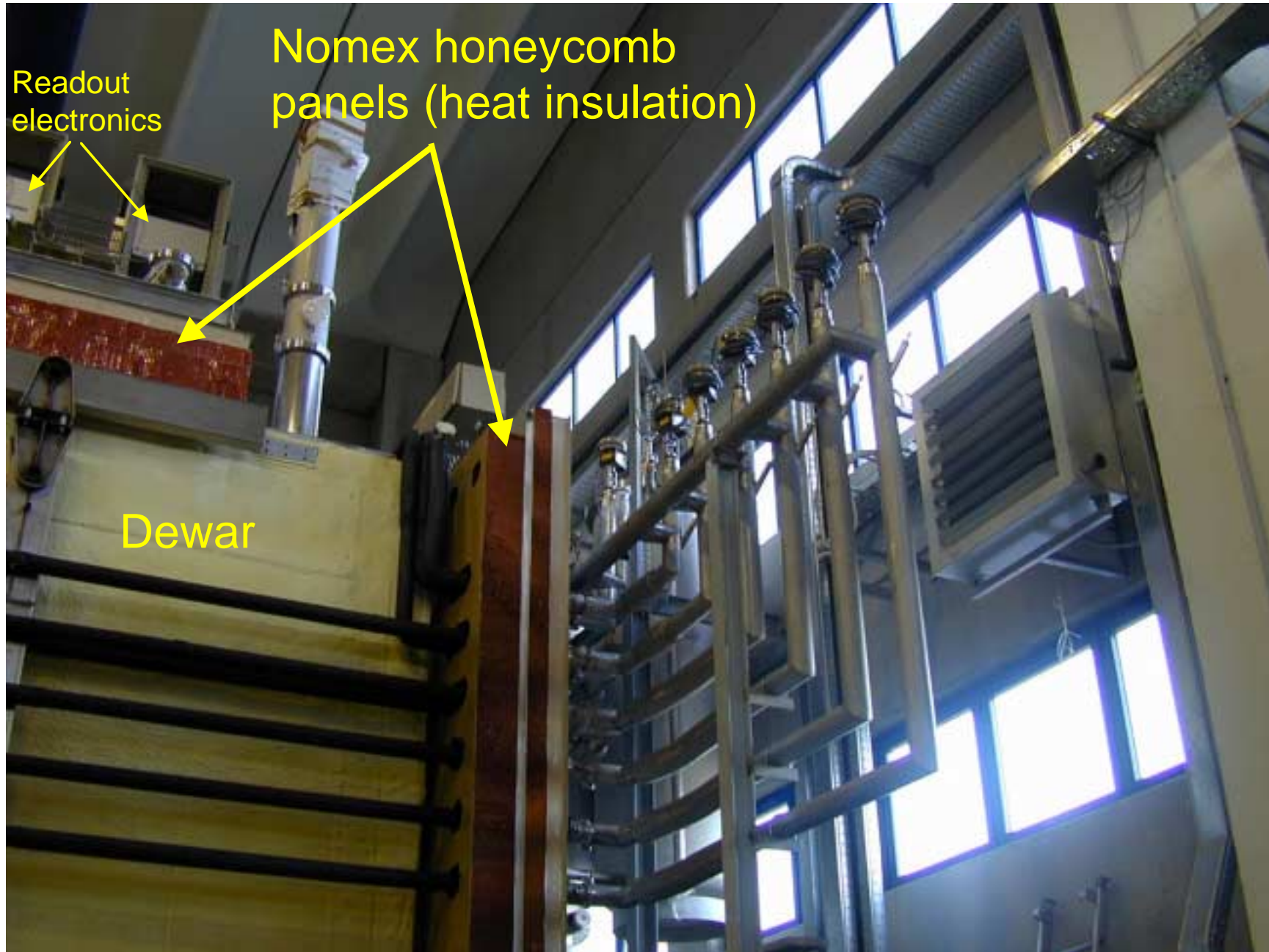


Second half-module (delivered Aug 2000)



Thermal insulation panels

Thermal floor



**Wire installation in T600 internal
detector (Jul-Oct 2000)**

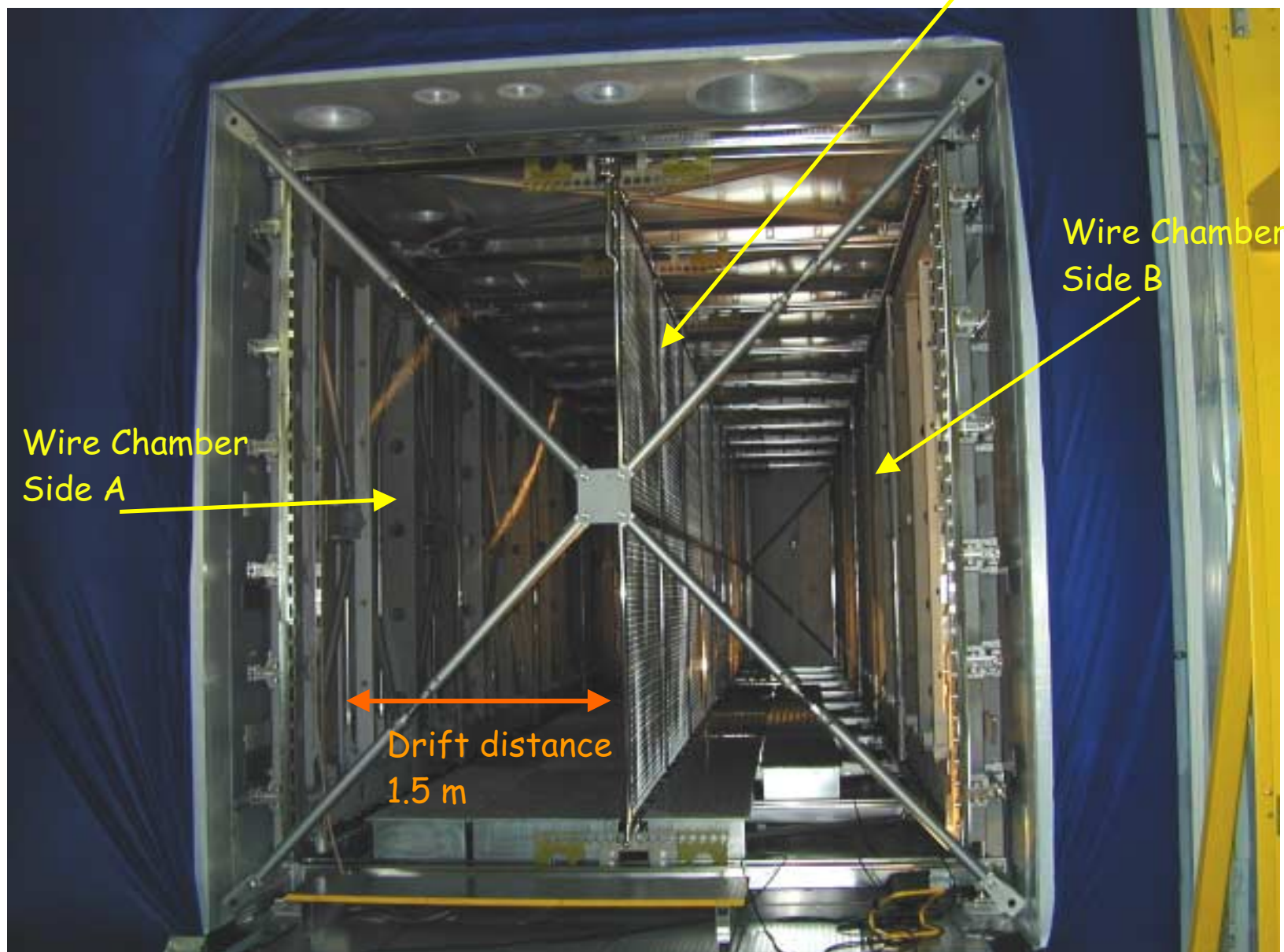


The three wire planes at $0^\circ, \pm 60^\circ$ (wire pitch = 3mm)



and one PMT

T600 - Completed Internal Detector view

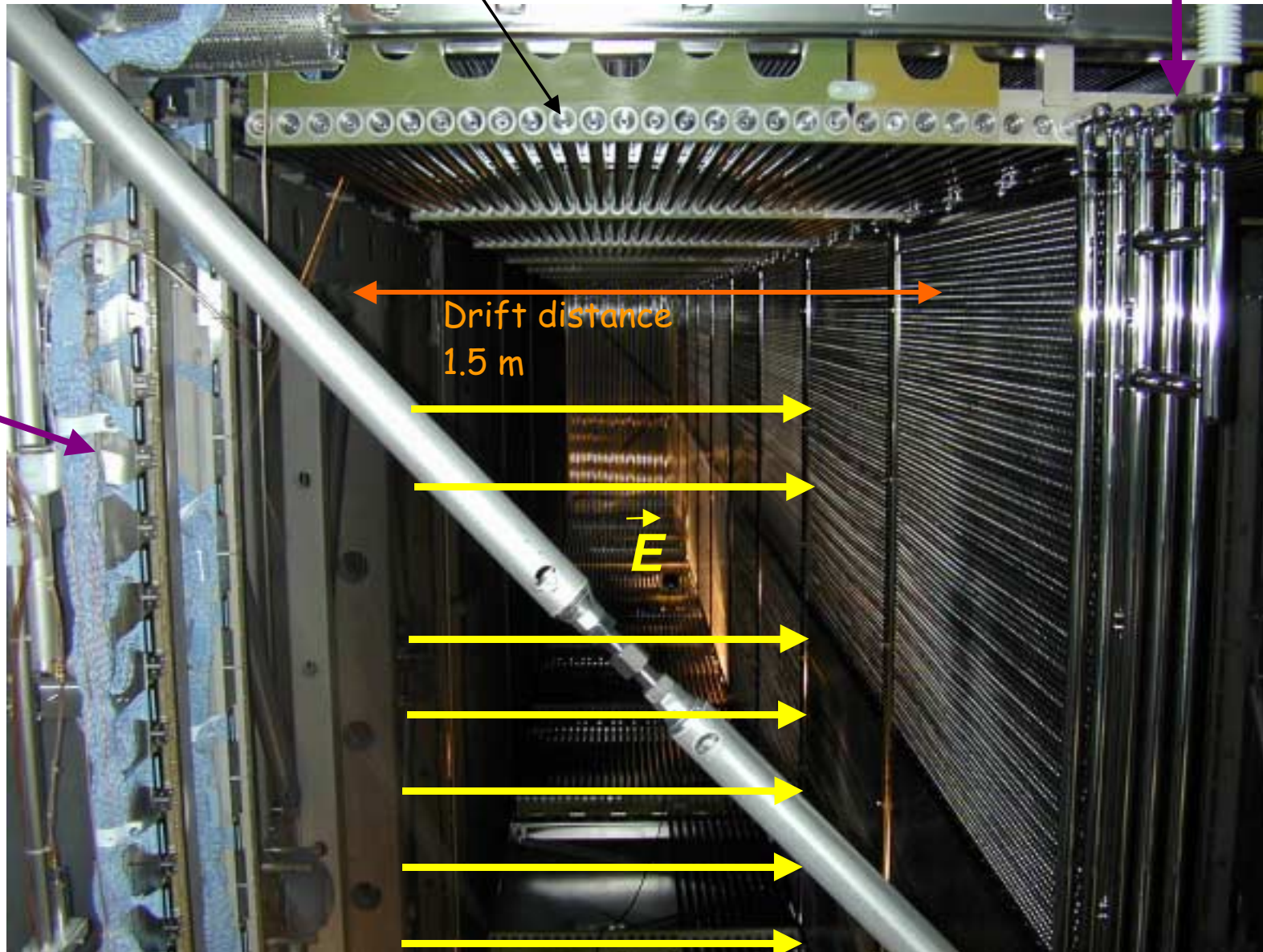


Drift H.V. and field electrodes system

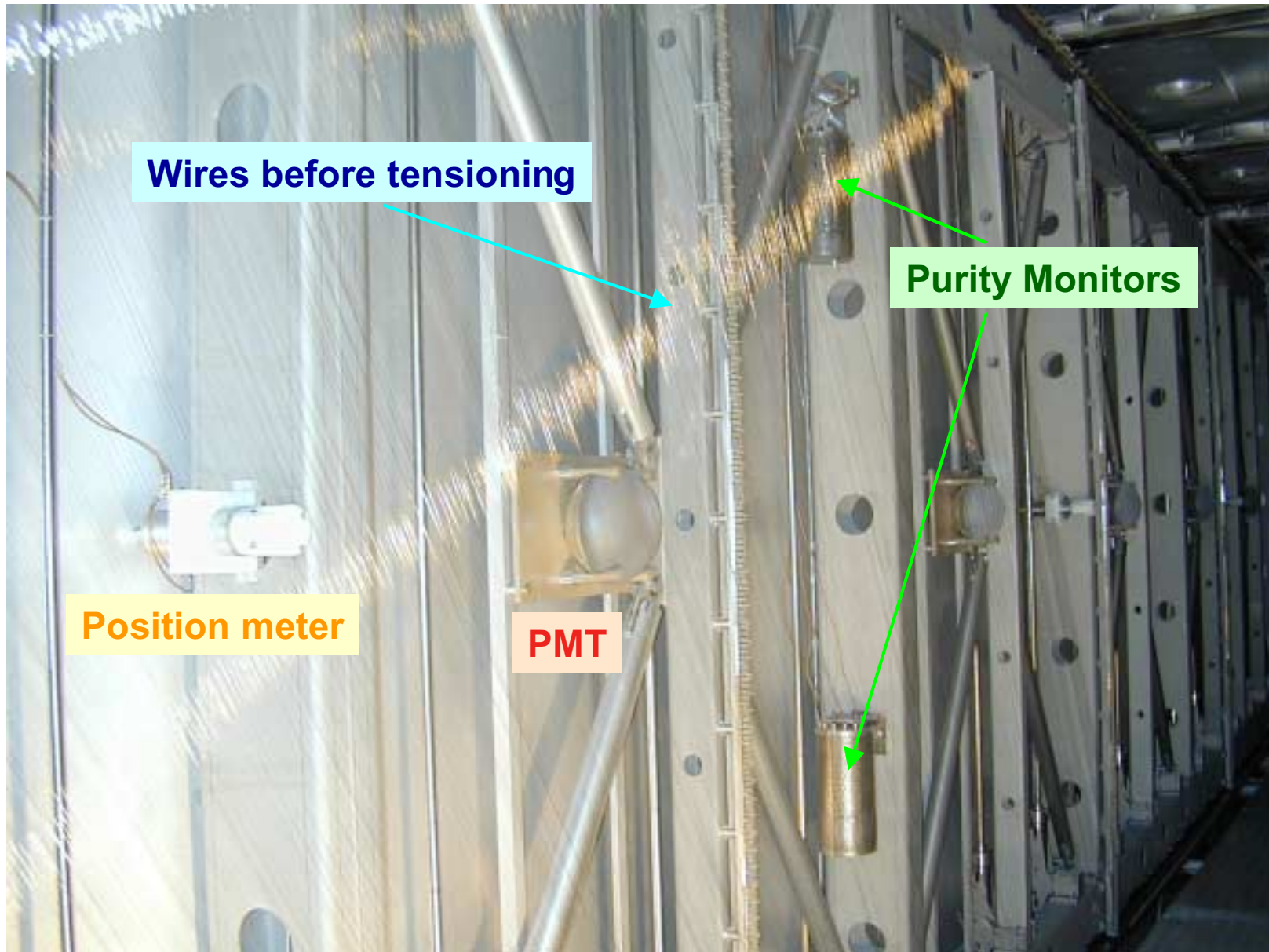
Race-track

-75kV

Horizontal wires readout cables



Slow control sensor (behind wire planes)

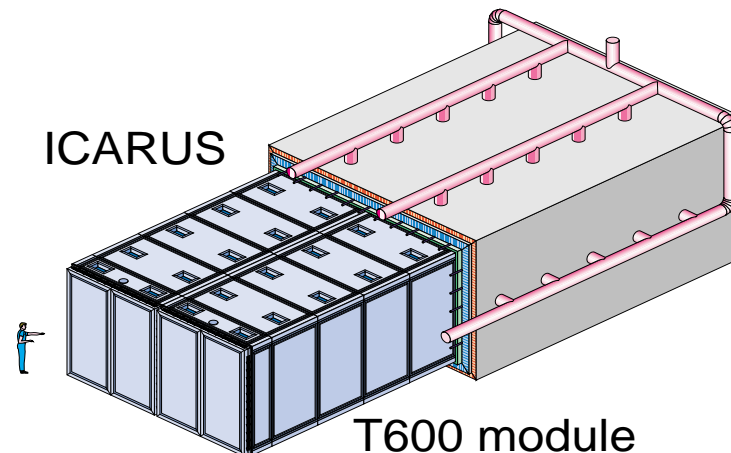


Man-hole (after sealing, the only way to get inside!)

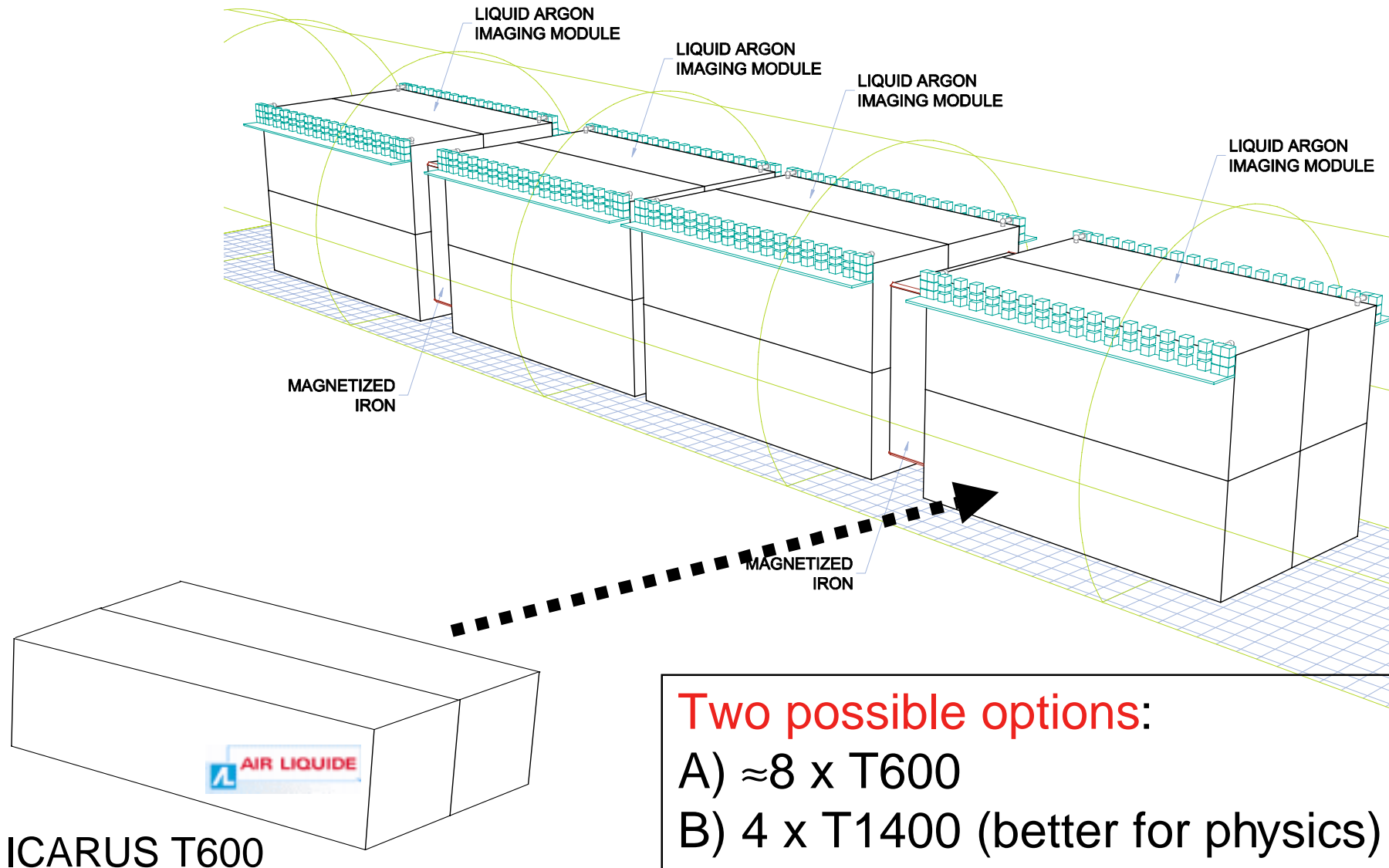


The first ICARUS T600 prototype

- ★ The T600 module is to be considered as a fundamental milestone on the road towards a total sensitive mass in the multi-kton range
 - First piece of the detector to be complemented by further modules of appropriate size and dimension ⇒ *Goal is to reach a multikton mass in LNGS tunnel in a most efficient and rapid way*
- ★ It has a physics program of its own, immediately relevant to neutrino physics, though limited by statistics (see hep-ex/0103008)

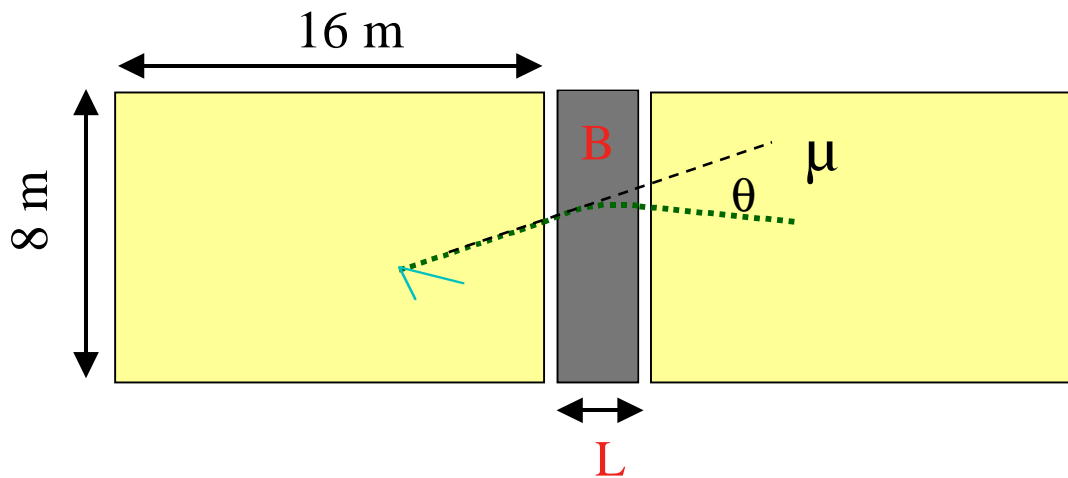


Proposed setup ICARUS 5kt in LNGS Hall B



Muon bending measurement

- ★ We consider a design in which the muon escaping the liquid Argon is bent by a magnetized piece of iron

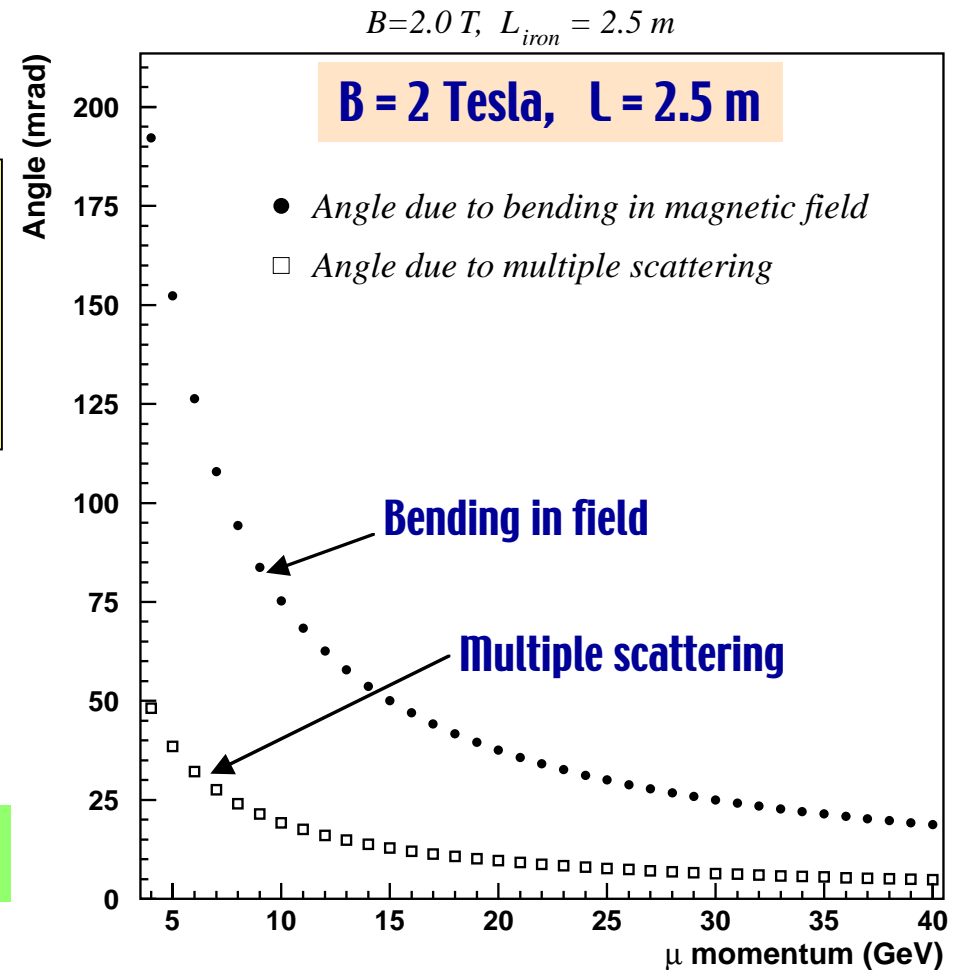


The bending angle θ is measured with the tracks observed in two subsequent liquid argon module

$\Delta p/p \approx 25\%$

Charge confusion: $\sim 10^{-4}$

A simpler solution than in the ICANOE proposal



Muon charge misidentification

★ μ momentum resolution:

→ 25% for a 2.5m long Fe spectrometer with $B=2T$

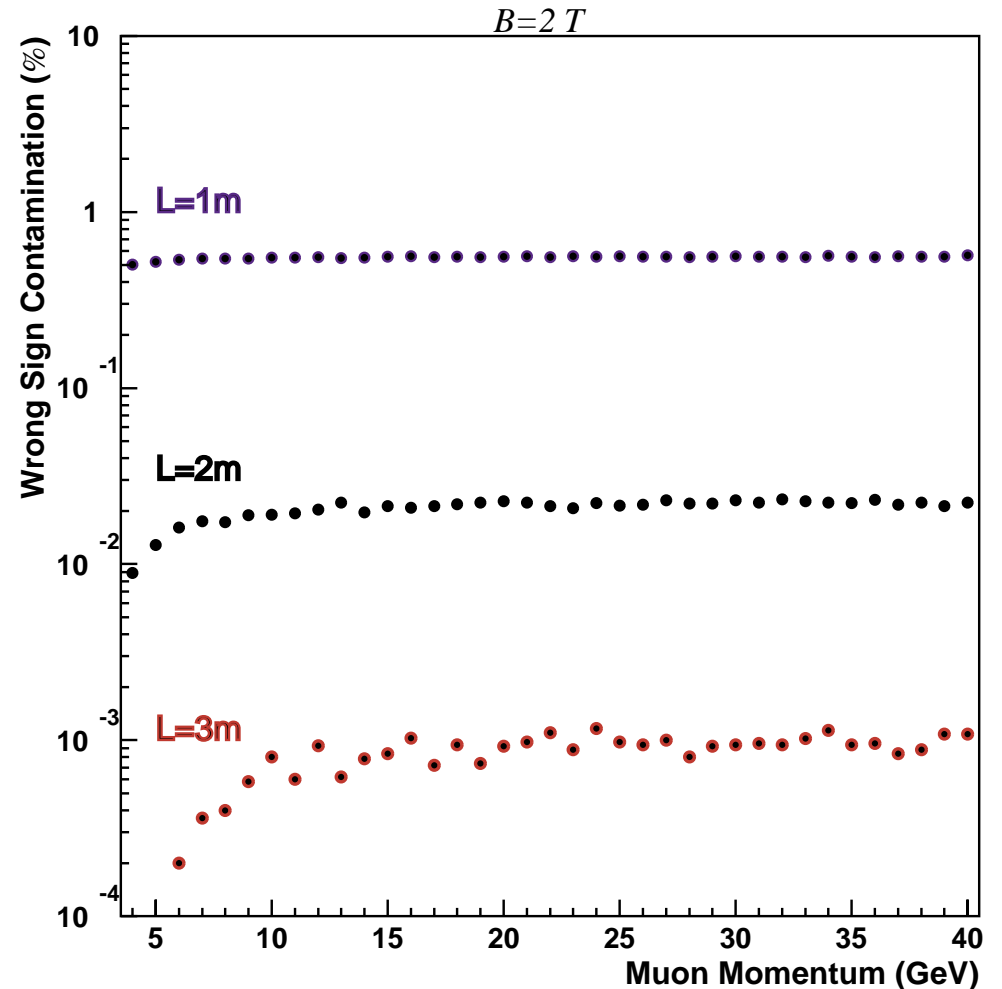
★ **Wrong sign contamination**

→ Charge confusion: $\sim 10^{-4}$

★ **Large detection efficiency for low energy beam**

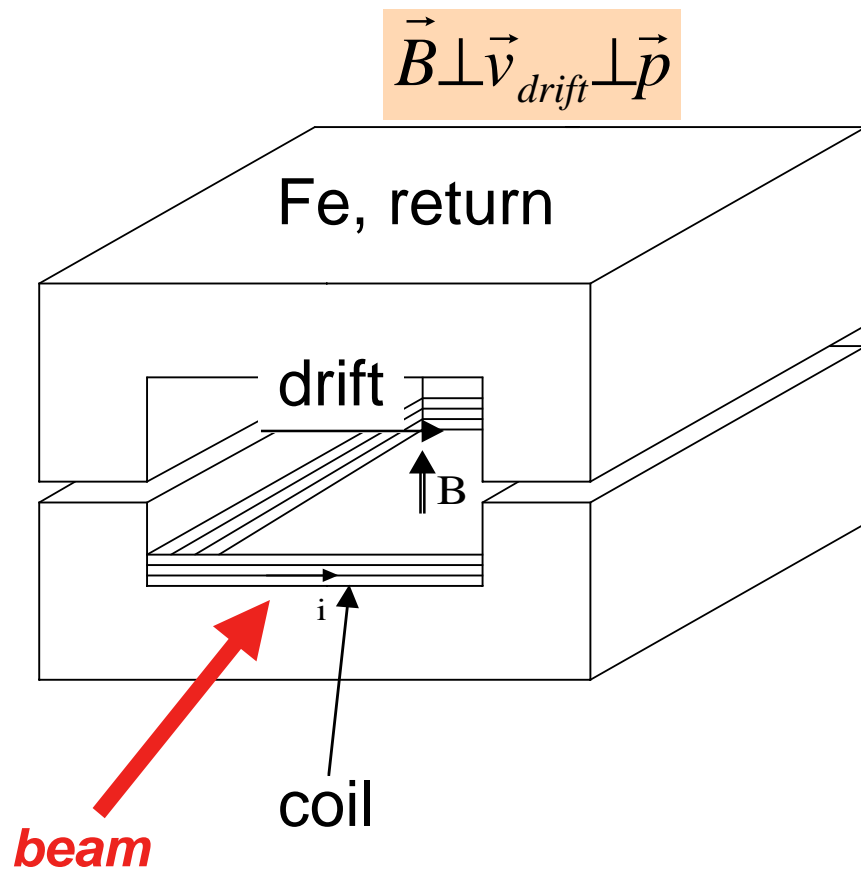
← μ detection threshold ($dE/dx = 240$ MeV/m)

Fraction (%) of misidentified charge



A large magnet ?

An interesting possibility, *to be further understood*, is the creation of the B-field over the large volume encompassing the LAr with the help of a very large solenoid



Joule Power (non-superconducting):

$$P = \rho \frac{2(a+b)hB^2}{md\mu_0^2}$$

d=coil thickness, m=#windings, h=height, a=width, b=length

Parameter	
Argon volume	$8 \times 8 \times 16m^3$
Argon mass	1.4 kton
Magnetic field	1.0 T
Current	2000 A
Conductor length	150 km
Resistance	1 Ω
Dissipated power	4 MW
Iron mass	5 kton

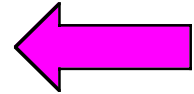
Neutrino Factory wish list



Precise determination of Δm^2_{23} and Θ_{23}



Stringent limit/precise measurement of Θ_{13}



Determination of Δm^2_{23} sign



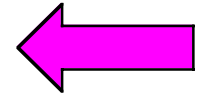
Study matter effects



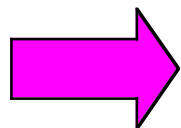
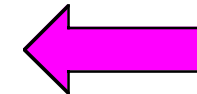
First detection of $\nu_e \rightarrow \nu_\tau$ oscillations



Over-constrain the oscillation parameters



Study CP violation in the leptonic sector



Try to show three concrete examples where detector considerations could be relevant...

Looking at the θ_{13} term

$$P(\nu_e \rightarrow \nu_\mu) = \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2(\Delta m_{32}^2 L/4E)$$

a realistic assumption at the NF \Rightarrow for $\Delta m_{21}^2 (L/4E) \ll 1$

$$\approx \sin^2 2\theta_{13} \sin^2 \theta_{23} (\Delta m_{32}^2)^2 (L/4E)^2$$

Not always a correct assumption at the NF \Rightarrow for $\Delta m_{21}^2 (L/4E) \ll \Delta m_{32}^2 (L/4E) \ll 1$

Similarly,

$$P(\nu_e \rightarrow \nu_\tau) = \sin^2 2\theta_{13} \cos^2 \theta_{23} \sin^2(\Delta m_{32}^2 L/4E)$$

for $\Delta m_{21}^2 (L/4E) \ll 1$

In contrast,

$$P(\nu_\mu \rightarrow \nu_\tau) = \cos^4 \theta_{13} \sin^2 2\theta_{23} \Delta_{32}^2$$

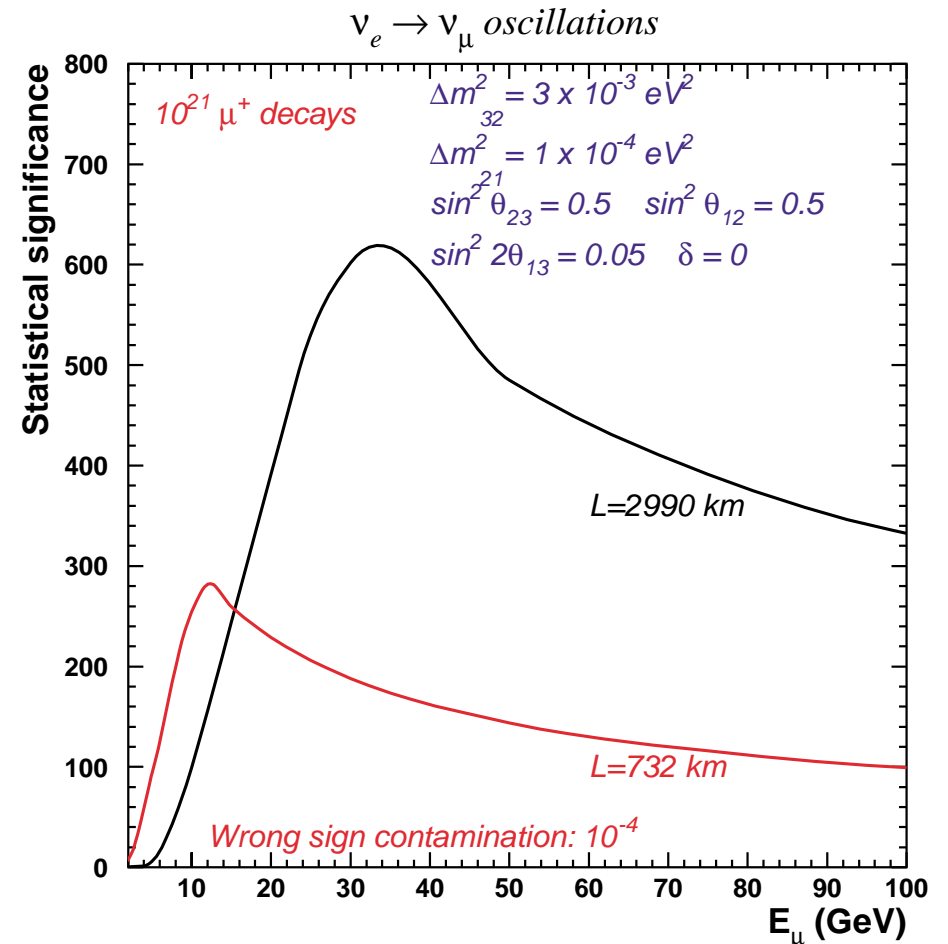
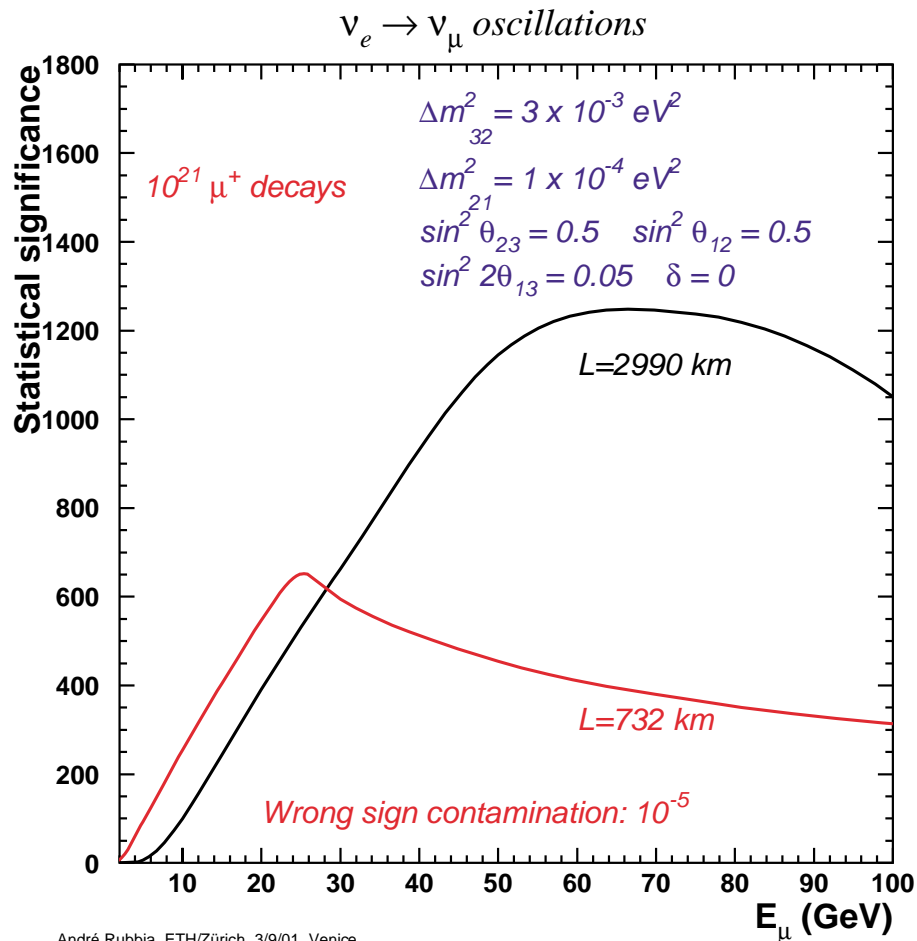
Wrong-sign muon optimization

Optimization of E_μ and L depends on detector background considerations

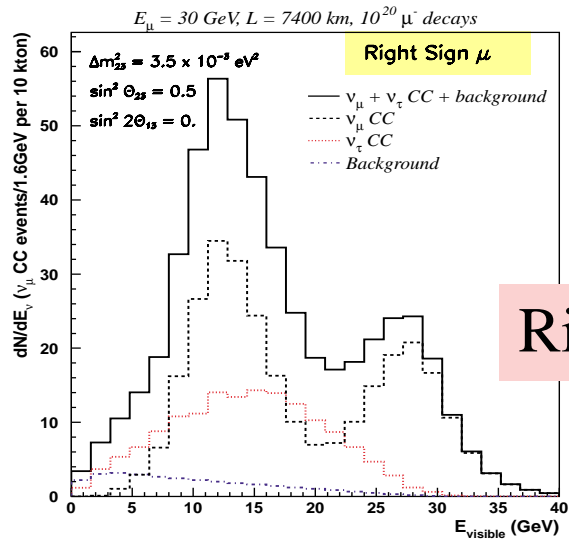
\Rightarrow gain with E_μ , since rate increases like E_μ^3 until background becomes relevant

$$S = \frac{Nws\mu}{B}$$

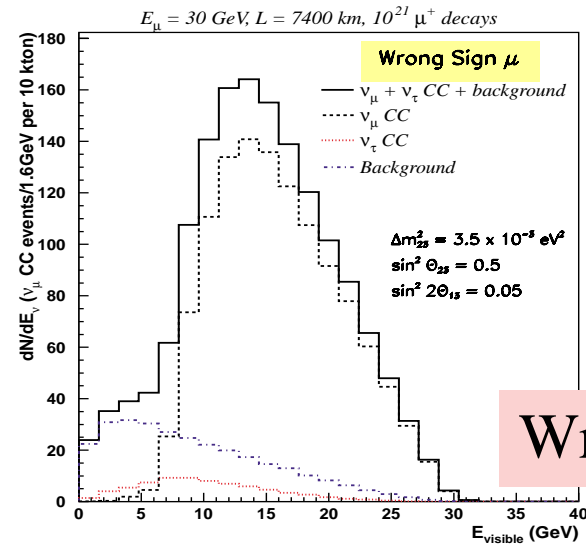
$$B = \begin{cases} \sqrt{Nrs\mu \times \epsilon_{BKG}} & \text{if } Nrs\mu \times \epsilon_{BKG} > 10 \\ \text{Poisson} & \text{otherwise} \end{cases}$$



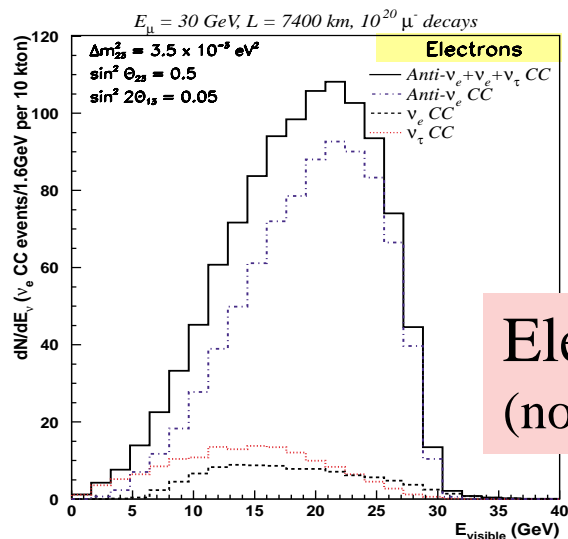
Over-constraining the parameters (I)



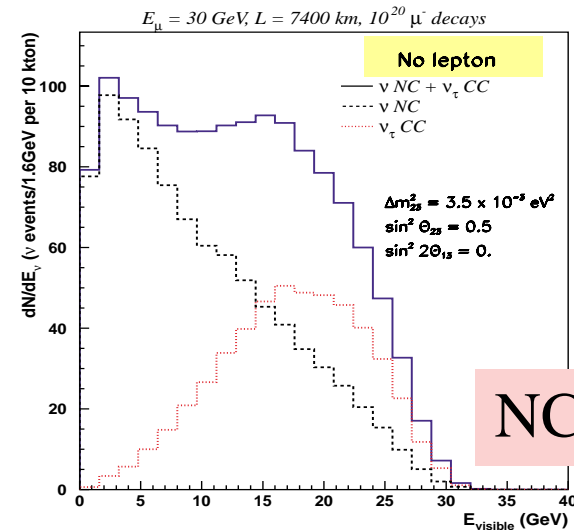
Right sign μ



Wrong sign μ



Electrons
(no charge info)



NC-like

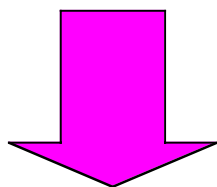
Combining all classes \Rightarrow (over-constrained) sensitivity to all oscillations!

Over-constraining the parameters (II)

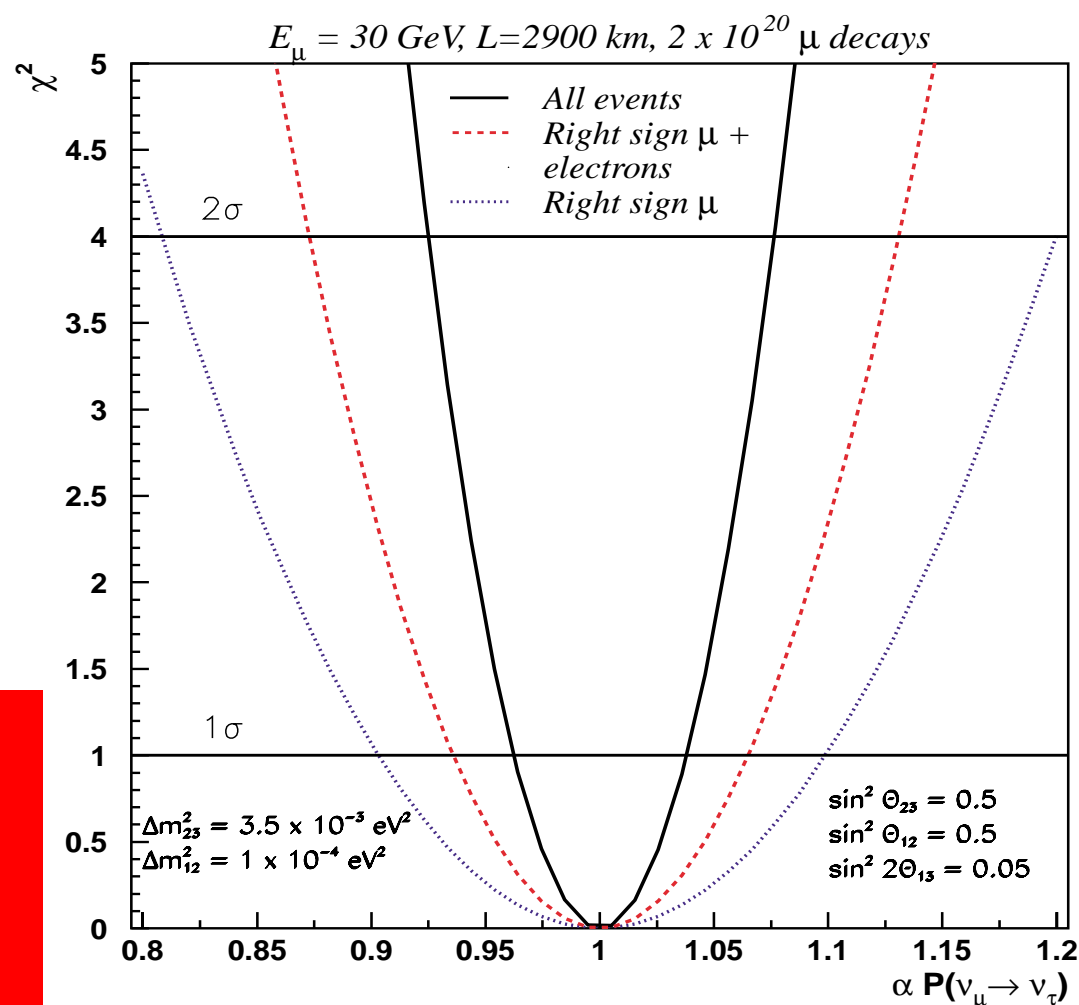
* Check consistency between different observed oscillation processes

* Proof/rule out the existence of sterile neutrinos

* **First observation of $\nu_e \rightarrow \nu_\tau$**



Ability to detect τ appearance is crucial



A. Bueno et.al. , Nucl.Phys.B589 (2000) 577

A way to rescale probabilities...

$$p \equiv P(\nu_e \rightarrow \nu_\mu) \times E_\nu^2 / L^2$$

probability

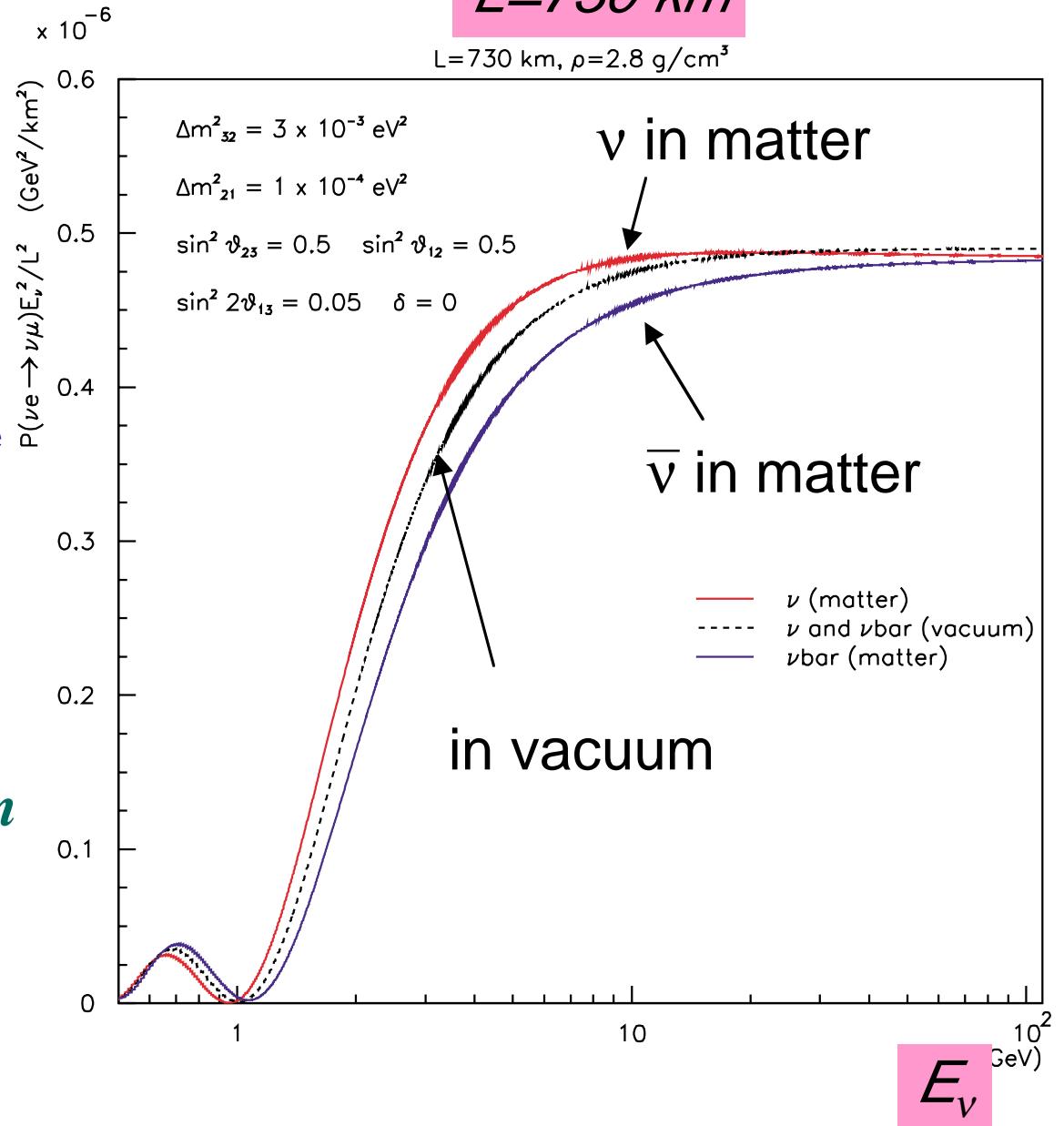
Approximate E_ν -dependence of NF ν -spectrum

Flux attenuation with distance

1. $p \rightarrow \text{const}$ when $E_\nu \rightarrow \infty$
2. It *correctly “weighs”* the probabilities with the E_ν dependence of the NF ν spectrum
3. p can be *directly compared at different baselines*

$L=730 \text{ km}$

$L=730 \text{ km}, \rho=2.8 \text{ g/cm}^3$



Matter effects

$$\sin^2 2\theta_m(D) = \frac{\sin^2 2\theta}{\sin^2 2\theta + \left(\pm \frac{D}{\Delta m^2} - \cos 2\theta \right)^2}$$

+ for neutrinos
- for antineutrinos

$$\lambda_m = L \times \sqrt{\sin^2 2\theta + \left(\pm \frac{D}{\Delta m^2} - \cos 2\theta \right)^2}$$

where

$$D(E_\nu) = 2\sqrt{2}G_F n_e E_\nu \approx 7.56 \times 10^{-5} \text{ eV}^2 \left(\frac{\rho}{\text{gcm}^{-3}} \right) \left(\frac{E}{\text{GeV}} \right)$$

For example, for neutrinos:

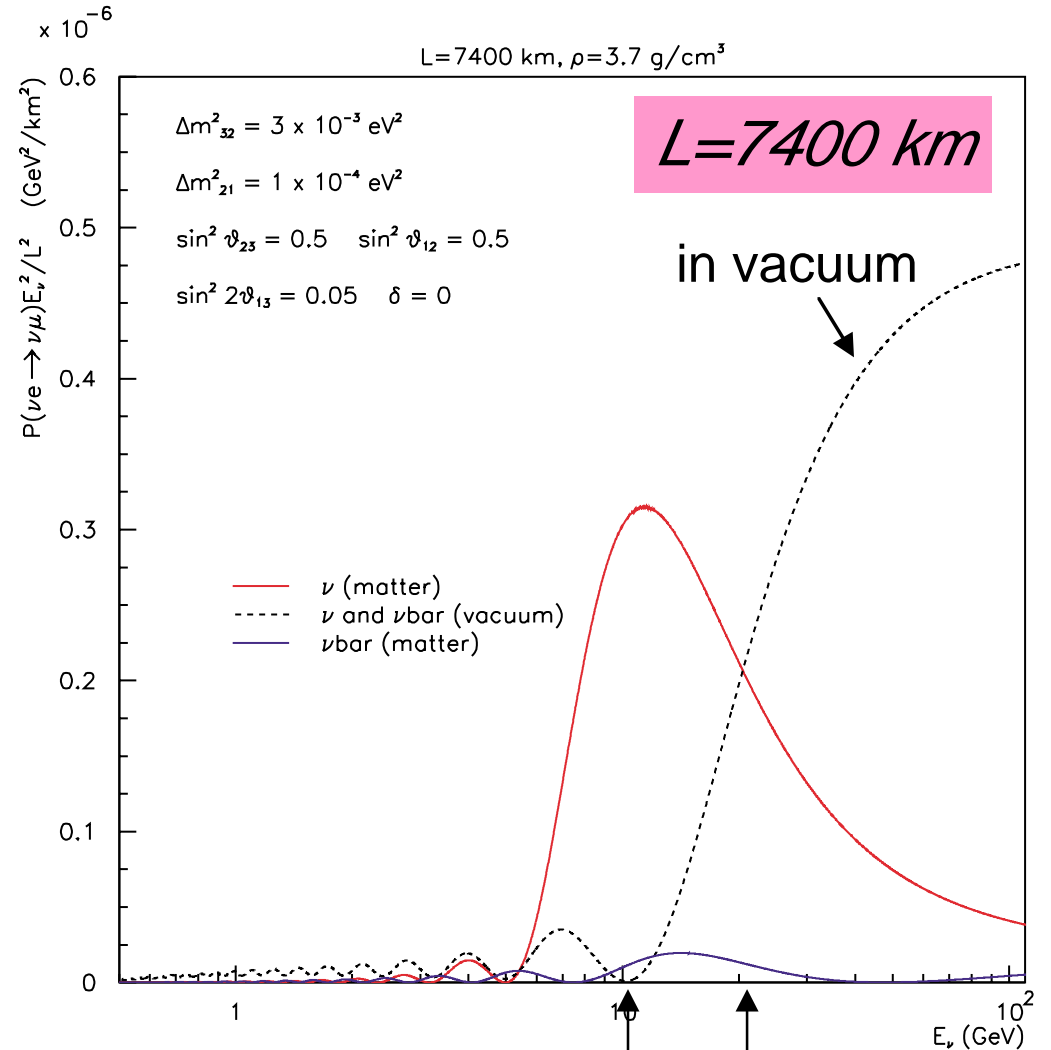
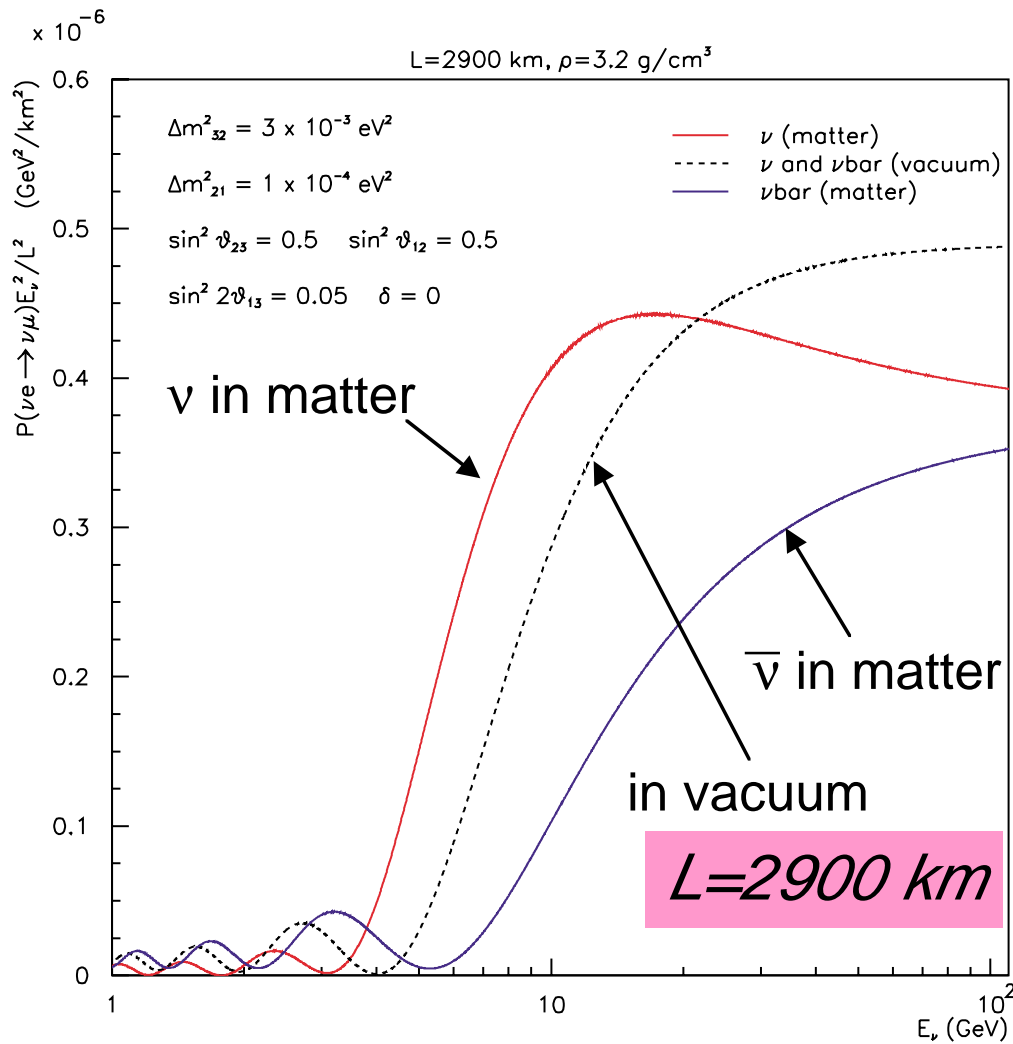
Resonance: $D \approx \Delta m^2 \cos 2\theta$ \longrightarrow $\sin^2 2\theta_m(D) \approx 1$

Suppression: $D > 2\Delta m^2 \cos 2\theta$ \longrightarrow $\sin^2 2\theta_m(D) < \sin^2 2\theta$

*Mixing in matter smaller than
in vacuum*

Effect tends to become “visible” for $L > \approx 1000 \text{ km}$

Behaviour at larger distances...



At large distances, matter effect suppresses oscillations!

$$D \approx \Delta m^2 \cos 2\theta$$

$$D \approx 2\Delta m^2 \cos 2\theta$$

Looking for effects of δ !

One of the main motivations of NF is to try to look for effects induced by the phase δ

Effect “largest” when beat of three sin-functions:

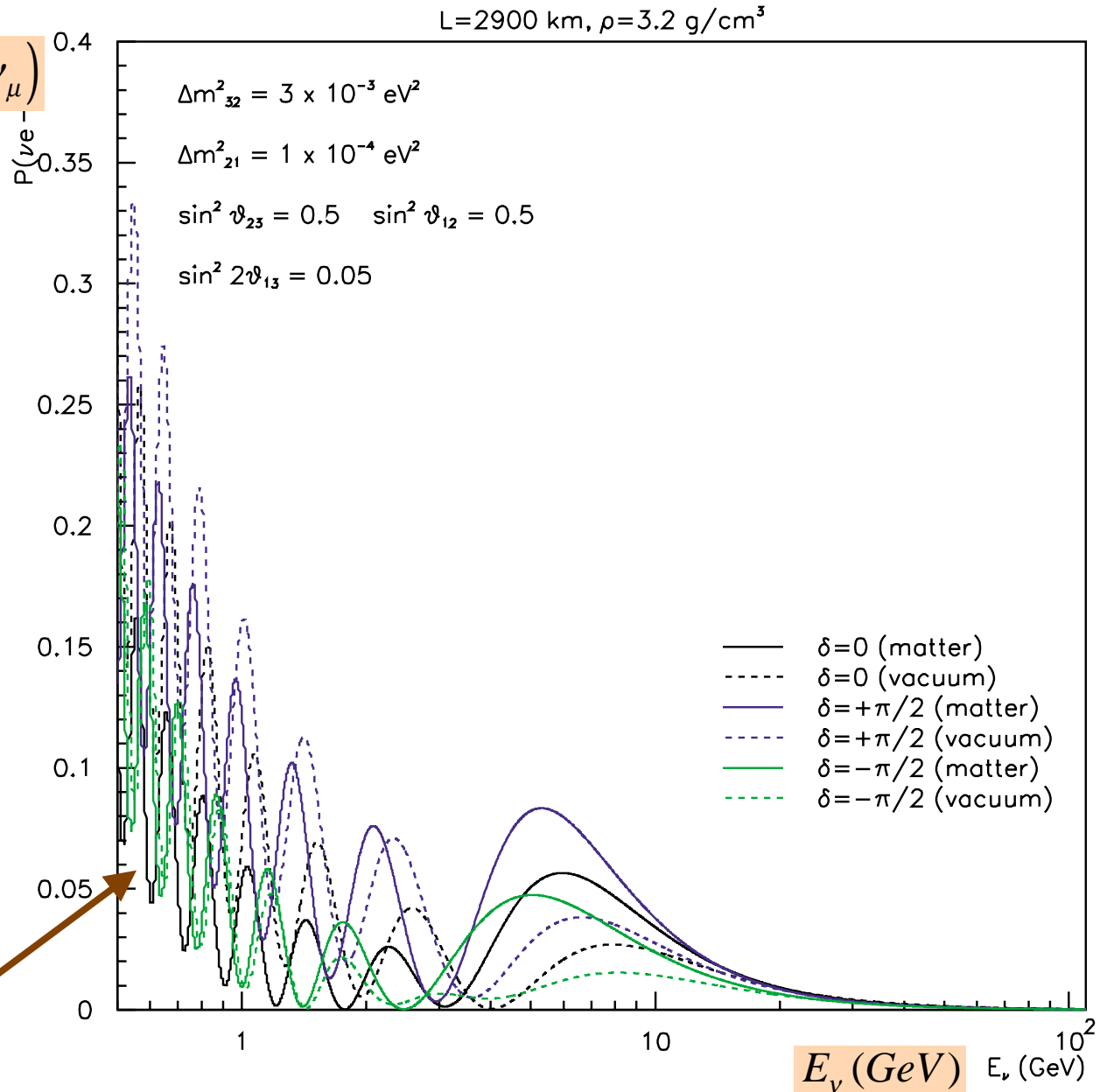
$$\Delta m^2_{21} (L/4E_\nu) > 1$$

&

$$\Delta m^2_{32} (L/4E_\nu) > 1$$

$\Rightarrow L/E$ of “solar” !

$$P(\nu_e \rightarrow \nu_\mu)$$



The “CP-odd” term

$$\Delta CP = \frac{P(\nu_e \rightarrow \nu_\mu) - P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu)}{2} =$$

⁻¹ Complex term in matrix Need LA MSW Oscillation P goes like $\sin^2\theta_{13}$
 hence, $\Delta CP/\sqrt{P}$ independent of θ_{13}

$$\propto \cos\theta_{13} \sin\delta \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \times$$

$$\sin(\Delta m_{12}^2 L/4E_\nu) \sin(\Delta m_{13}^2 L/4E_\nu) \sin(\Delta m_{23}^2 L/4E_\nu)$$

$$\approx f \times \Delta m_{12}^2 (L/4E_\nu) \times \sin^2(\Delta m_{23}^2 L/4E_\nu)$$

for $\Delta m_{21}^2 (L/4E_\nu) \ll 1$

$$\approx f \times \Delta m_{12}^2 (\Delta m_{23}^2)^2 (L/4E_\nu)^3$$

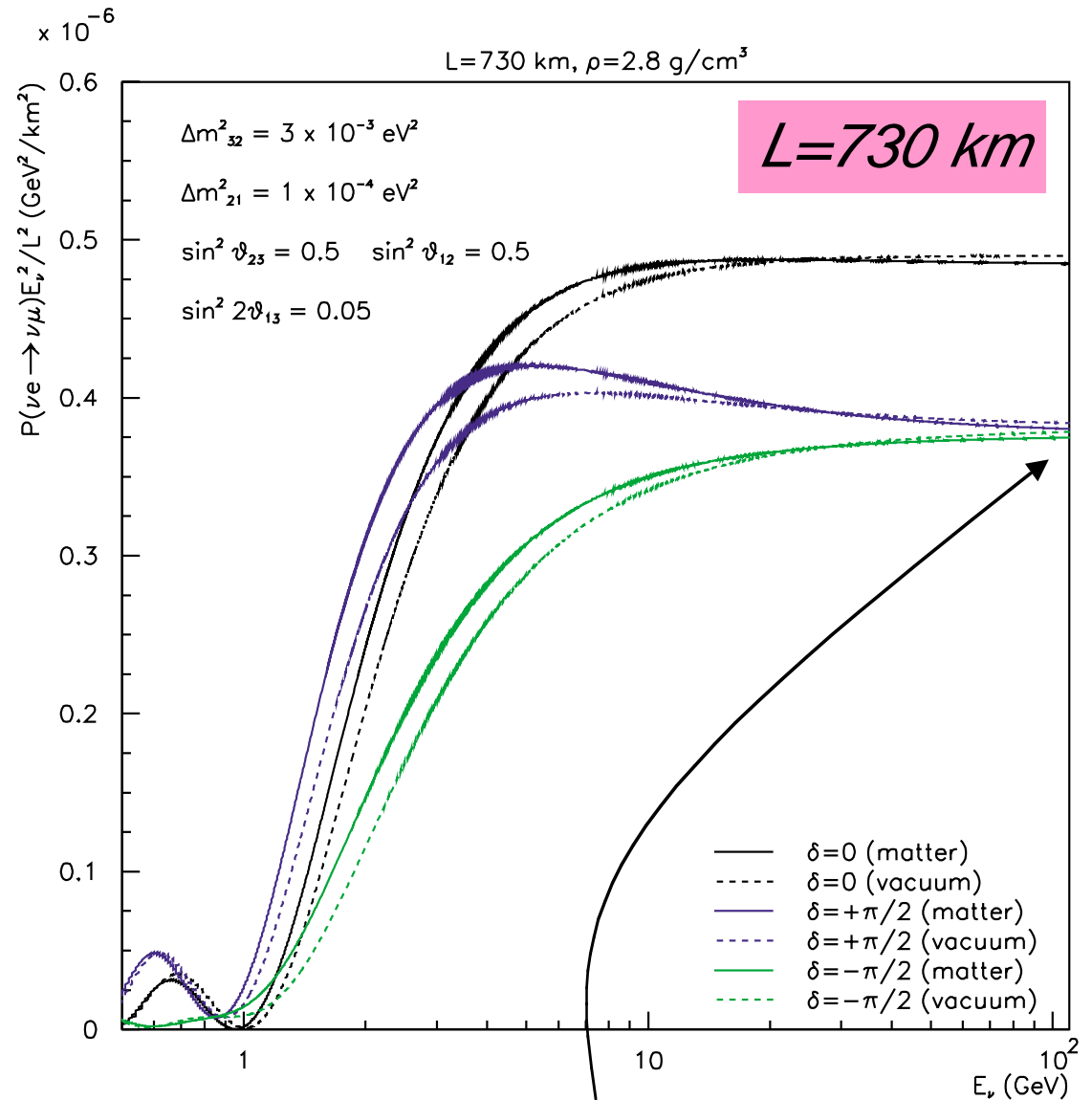
for $\Delta m_{21}^2 (L/4E_\nu) \ll \Delta m_{32}^2 (L/4E_\nu) \ll 1$

So what to do at high energy?

See also P. Lipari, hep-ph/0102046

$$P(\nu_e \rightarrow \nu_\mu) \times E_\nu^2 / L^2$$

1. The E_ν^2 term takes into account that the NF likes to go to high energy \Rightarrow damps the part $\Delta m_{21}^2 (L/4E_\nu) \approx 1$
2. At “high energy”, i.e. $\Delta m_{21}^2 (L/4E_\nu) \ll 1$ & $\Delta m_{32}^2 (L/4E_\nu) \ll 1$, there is no more oscillation \Rightarrow **change of $\delta =$ change of θ_{13} !!!**
3. At “high energy”, the CP-effect goes like $\cos\delta$, as pointed out by Lipari \Rightarrow cannot measure sign of δ



$$P\left(\nu_e \rightarrow \nu_\mu, \delta = \frac{\pi}{2}\right) - P\left(\nu_e \rightarrow \nu_\mu, \delta = 0\right) \propto \cos\delta$$

So where is the compromise in L/E?

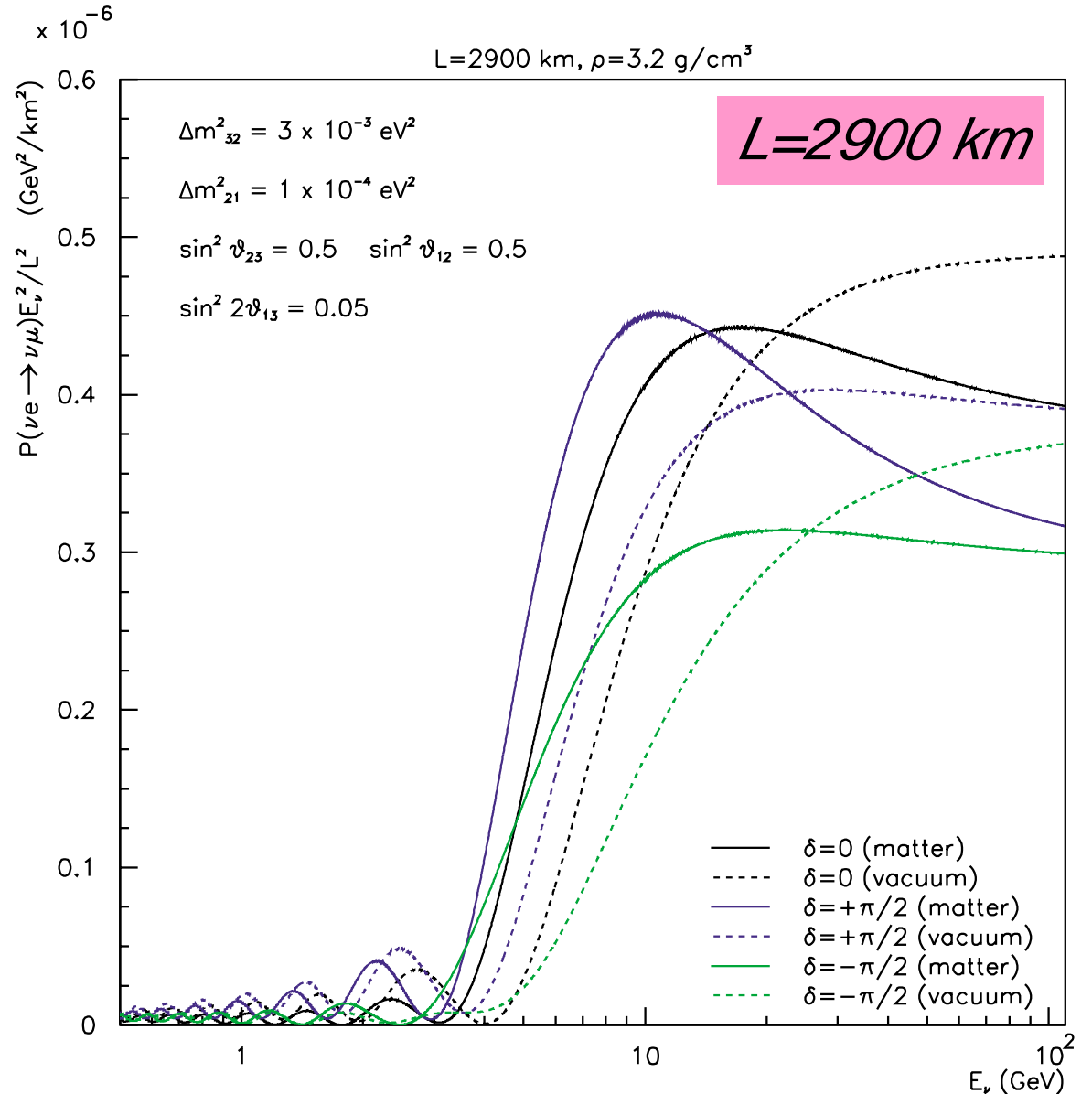
We must compromise at “medium” energy to

1. This means $\Delta m^2_{21}(L/4E_\nu) \ll 1$
& $\Delta m^2_{32}(L/4E_\nu) \approx 1$
2. To gain from the E_μ^3 behavior of the NF
3. To guarantee the possibility to disentangle δ from θ_{13}

➔
$$\frac{L}{E_\nu} \approx \frac{4\pi}{2\Delta m^2_{32}}$$

⚡ $E_{\nu, \text{MAX}} \sim 2 \text{ GeV}$ for $L=732 \text{ km}$

⚡ $E_{\nu, \text{MAX}} \sim 8 \text{ GeV}$ for $L=2900 \text{ km}$



If L/E_ν is fixed, what should be L and E_ν ?

The magnitude of the CP effect (given by J) is known to be unaffected by matter

$$J = \cos\theta_{13} \sin\delta \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} / 8$$

Our “choice-point” for CP is at the fixed $L/E_{\nu,\max}$ given by:
$$E_{\nu,\max} = \frac{2 \times 1.27 \times \Delta m^2 L}{\pi}$$

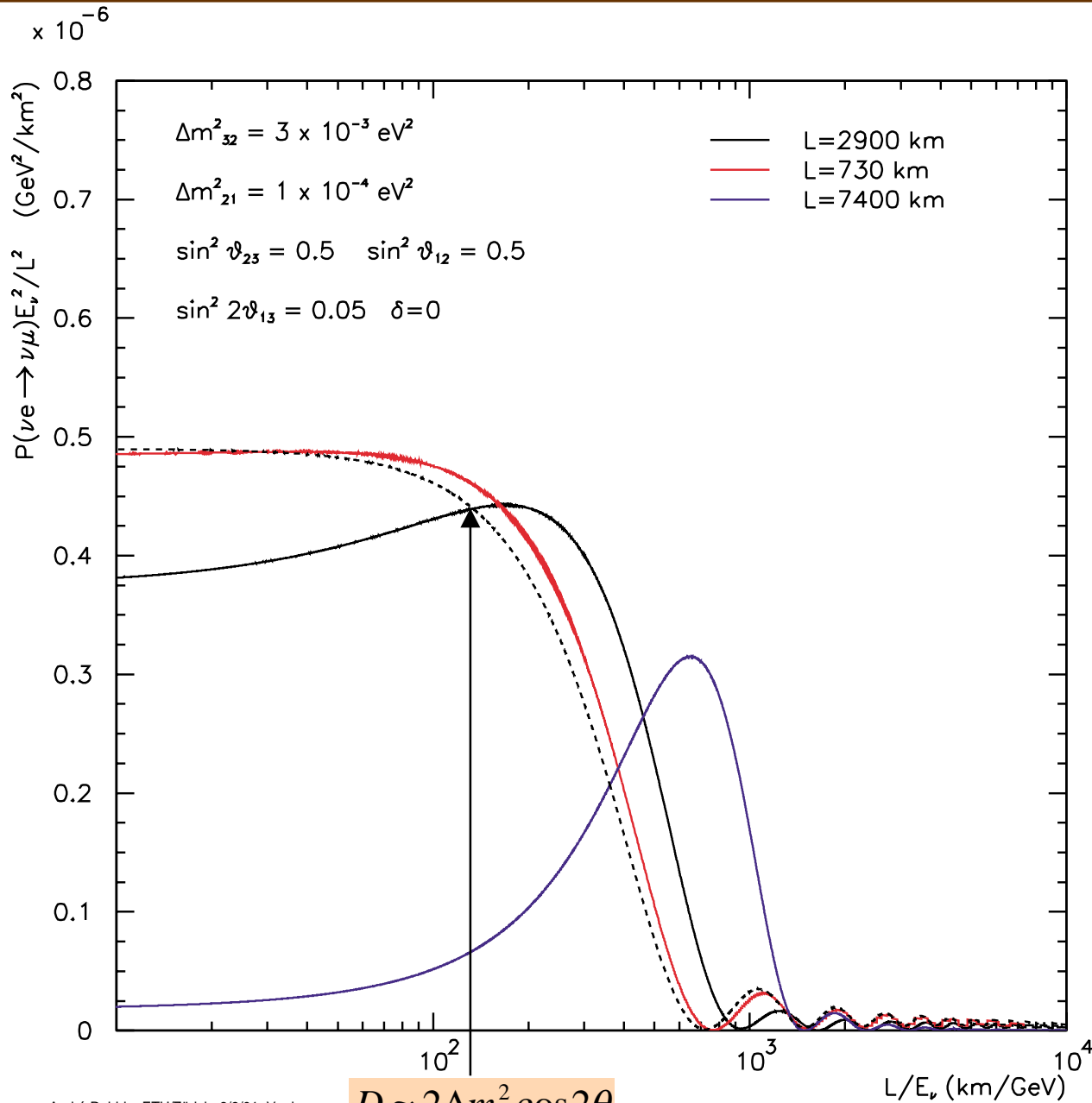
When the neutrino energy becomes close to the MSW resonance, the effective oscillation wavelength increases, hence the CP effect at a fixed distance L becomes less visible.

Hence, we gain until the MSW resonance region and then loose

$$2\sqrt{2}G_F n_e E_\nu < \Delta m^2 \cos 2\theta \quad \Rightarrow \quad 2\sqrt{2}G_F n_e \frac{2 \times 1.27 \Delta m^2 L}{\pi} < \Delta m^2 \cos 2\theta$$

$$L < \frac{\pi \cos 2\theta}{2 \times 1.27 \times 7.56 \times 10^{-5} \text{ eV}^2 \left(\frac{\rho}{\text{gcm}^{-3}} \right)} \approx \frac{1.5 \times 10^4 \text{ km}}{\left(\frac{\rho}{\text{gcm}^{-3}} \right)} \approx 5000 \text{ km}$$

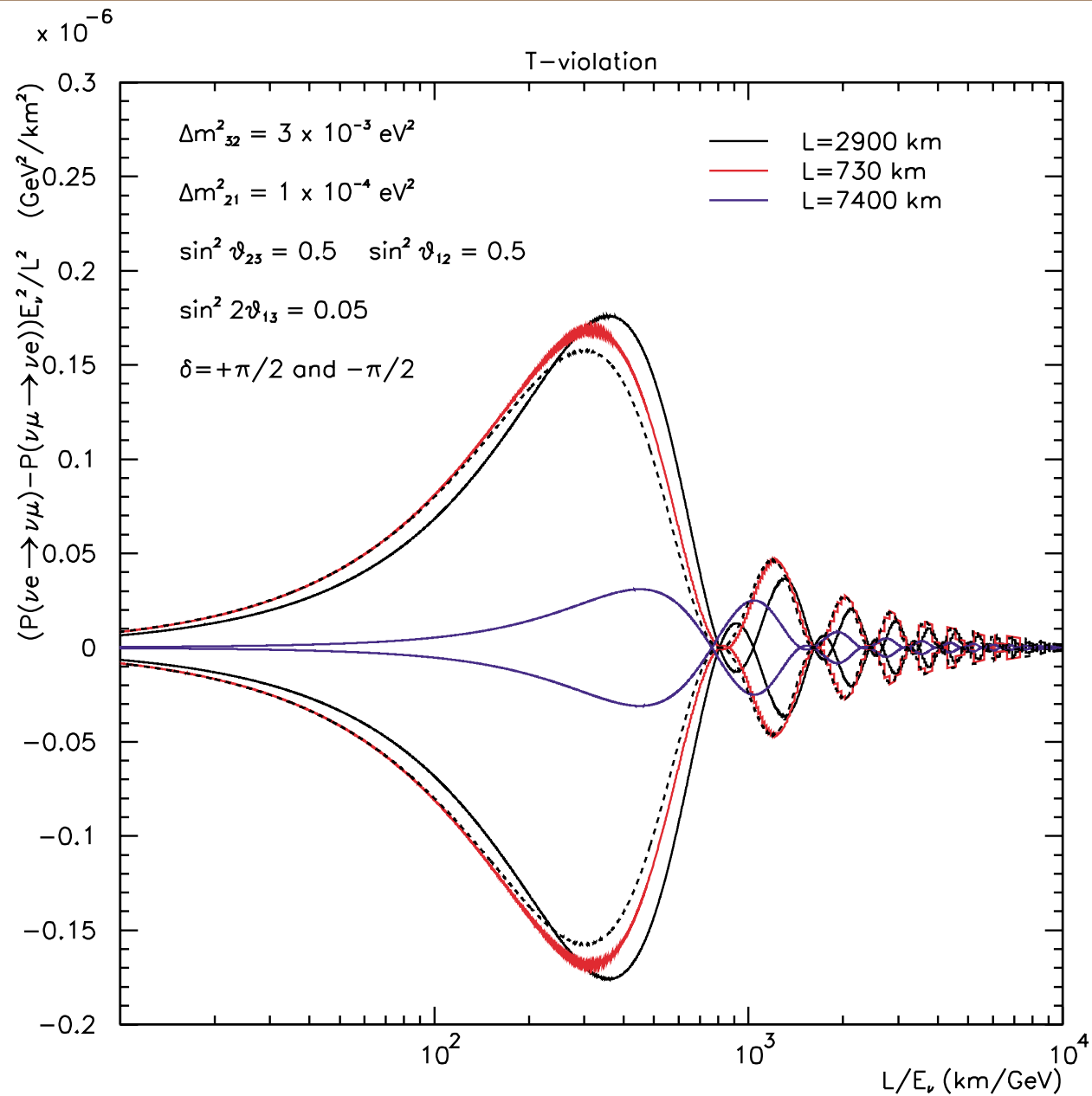
Dependence of probability on L/E_ν



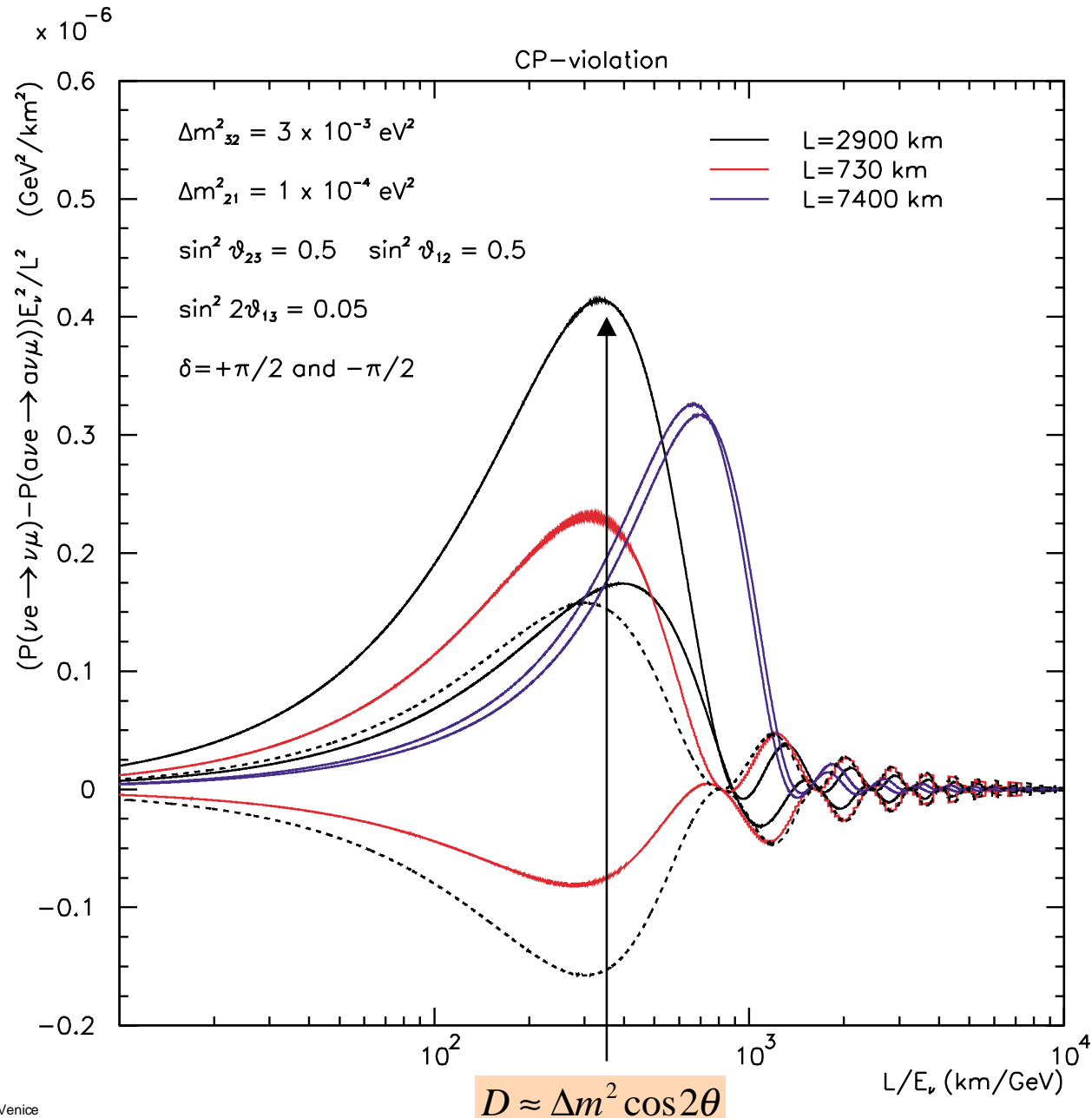
The “scaling” with L/E_ν of the probabilities is destroyed when $E_{\nu, \max} > E_{\nu, \text{resonance}}$ due to matter effects.

$$D \approx 2\Delta m^2 \cos 2\theta$$

The T-violation term dependence on L/E_ν



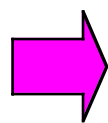
The CP-violation term dependence on L/E_ν



Fit with constant L/E

As long as $L < \approx 5000$ km, the effects scale with L/E_ν

However, keeping L/E_μ constant, we gain linearly with E_μ because of the NF flux dependence E_μ^3/L^2



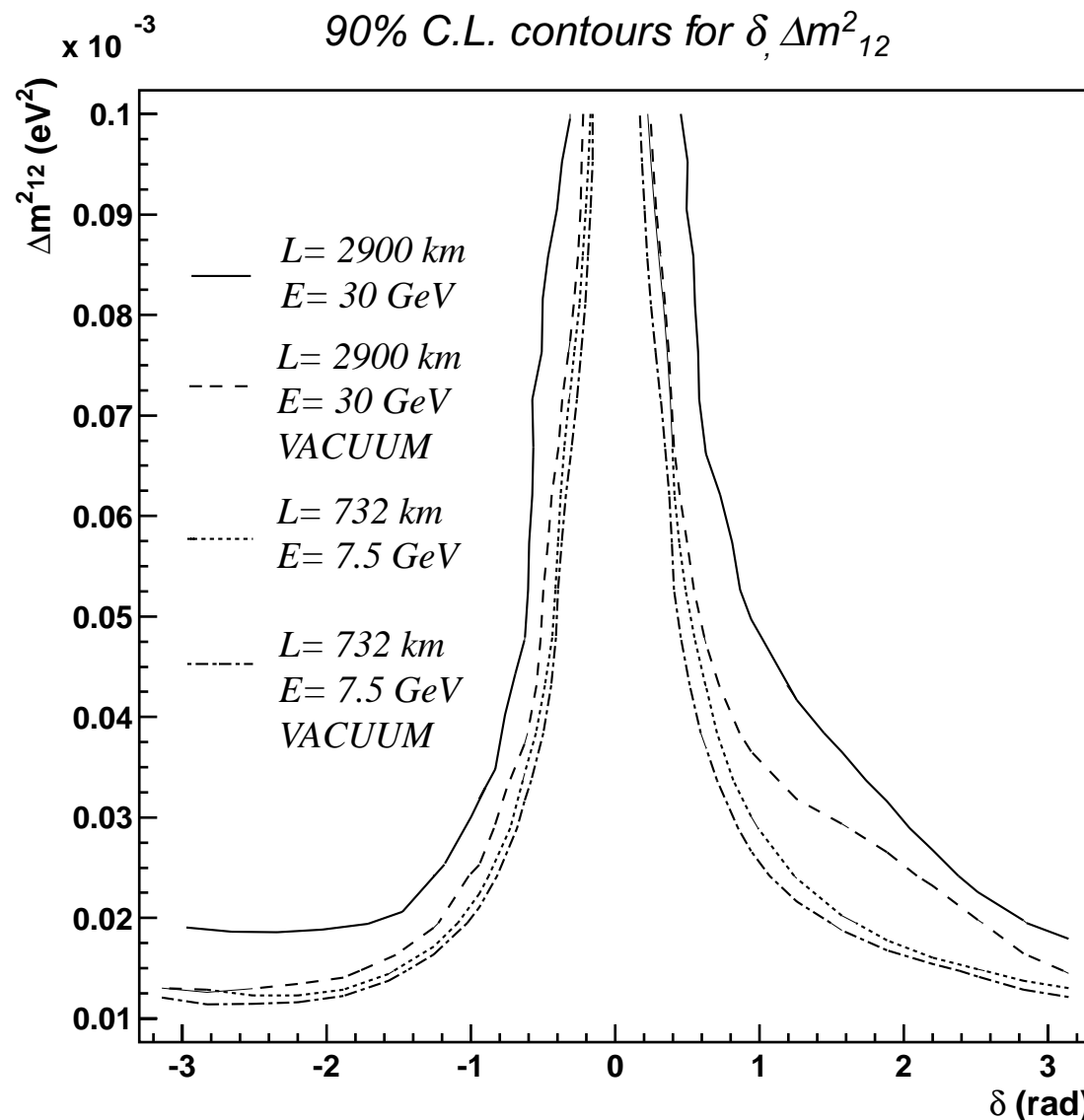
At lower E_μ must compensate with higher muon intensities

$E_\mu = 30$ GeV & $L = 2900$ km

→ $2.5 \cdot 10^{20}$ decays

$E_\mu = 7.5$ GeV & $L = 732$ km

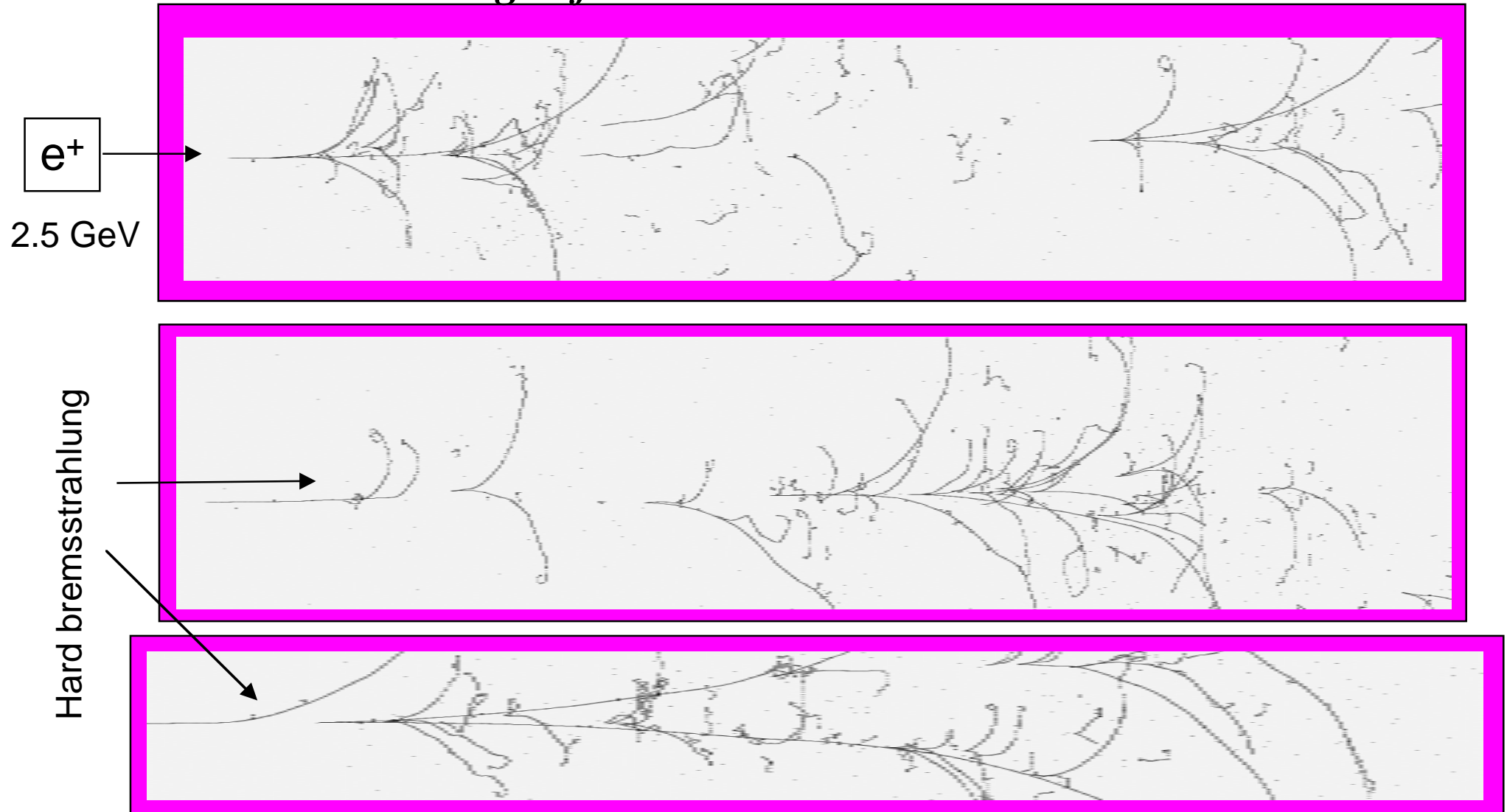
→ 10^{21} decays



On the possibility to measure the electron charge

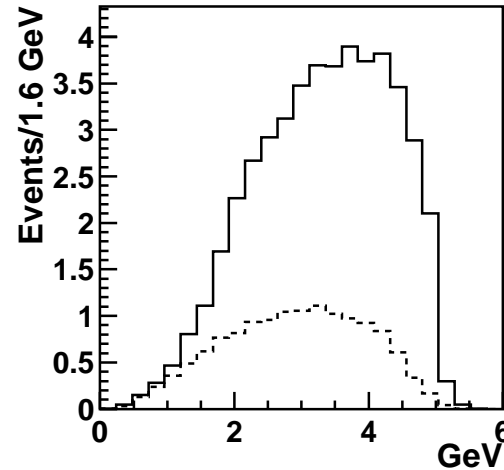
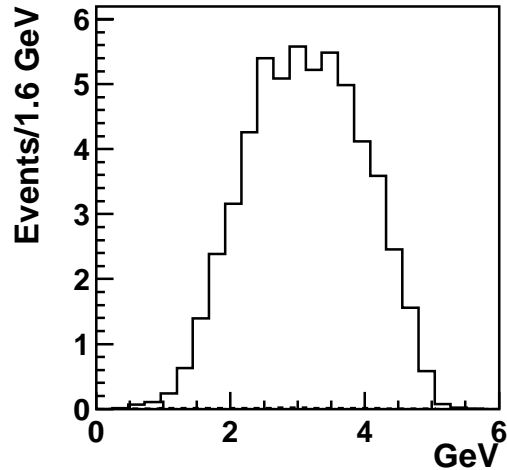
The presence of a magnetic field surrounding the LAr should allow to even determine the charge of electrons

B=1T

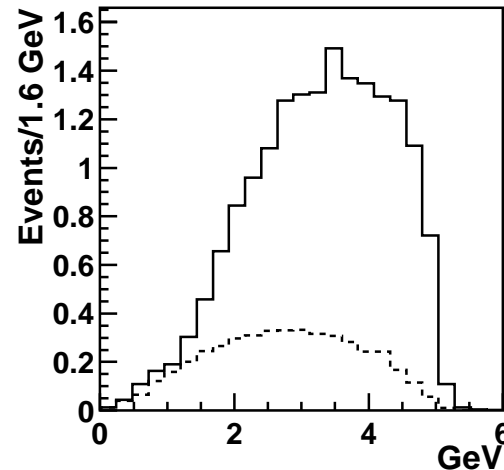
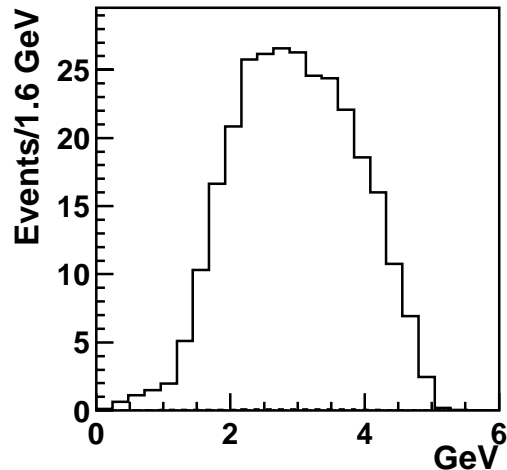


Wrong-sign lepton spectra

μ^-
decays



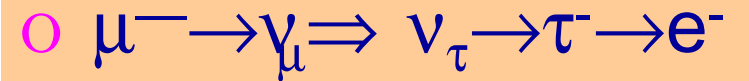
μ^+
decays



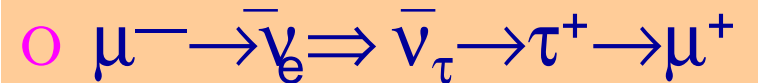
WS ν_μ CC

WS ν_e CC

★ Main background for WSE:



○ Wrong-sign electrons from τ decays not suppressed as in the WSM case:



$\propto \sin^2 2\theta_{13}$

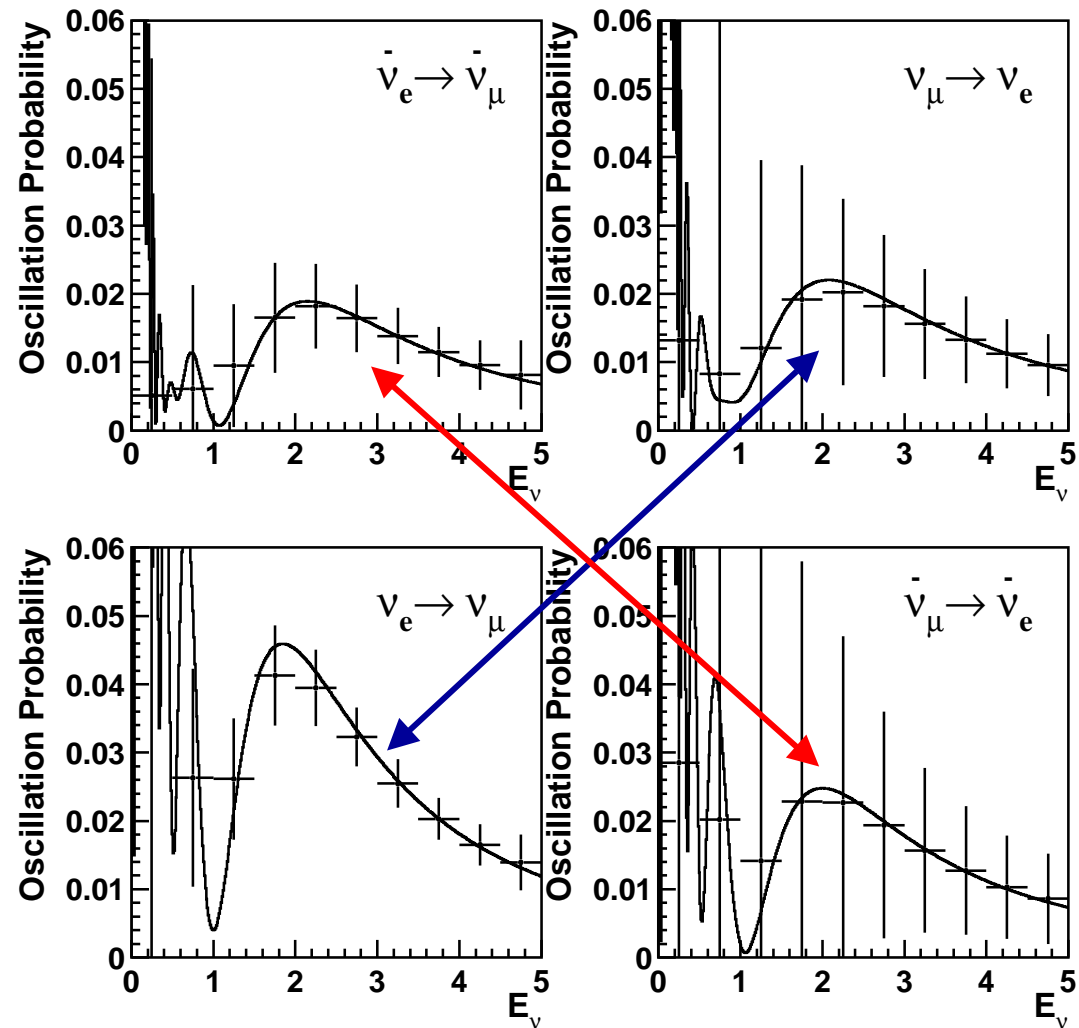
One more reason to go to lower energies where τ production is suppressed

Measured oscillation probabilities

- $\Delta m_{23}^2 = 3.5 \times 10^{-3} \text{ eV}^2$
- $\Delta m_{12}^2 = 1. \times 10^{-4} \text{ eV}^2$
- $\sin^2 2\theta_{13} = 0.05$
- $\sin^2 2\theta_{23} = 1.$
- $\sin^2 2\theta_{12} = 1.$
- $\delta_{13} = \pi/2$
- $10^{21} \mu$ decays $E_\mu = 5 \text{ GeV}$
- 10 kton detector

Direct comparison of oscillation probabilities for neutrinos and antineutrinos

$\varepsilon = 20\%$, *w.s. background* 10^{-3}



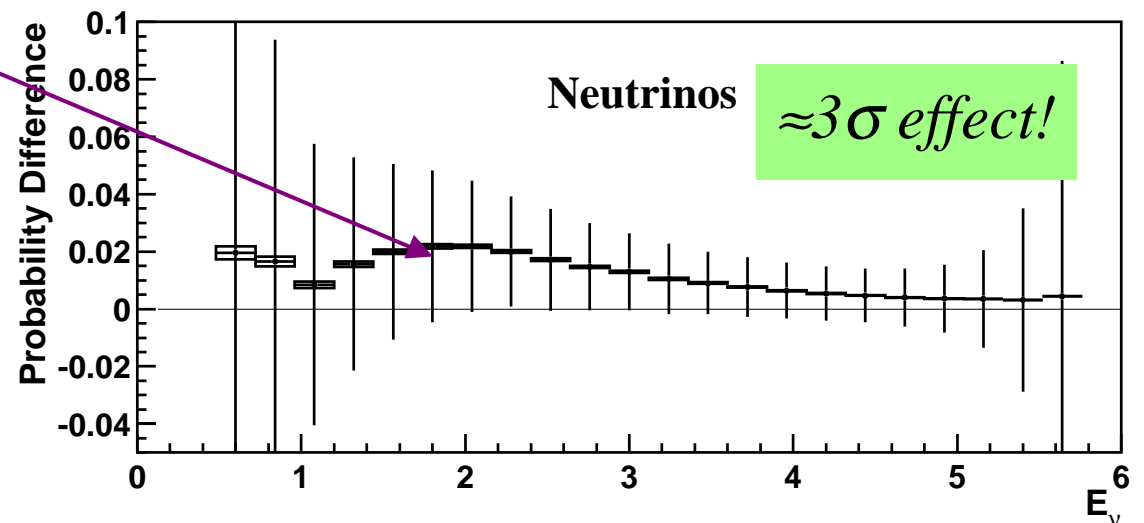
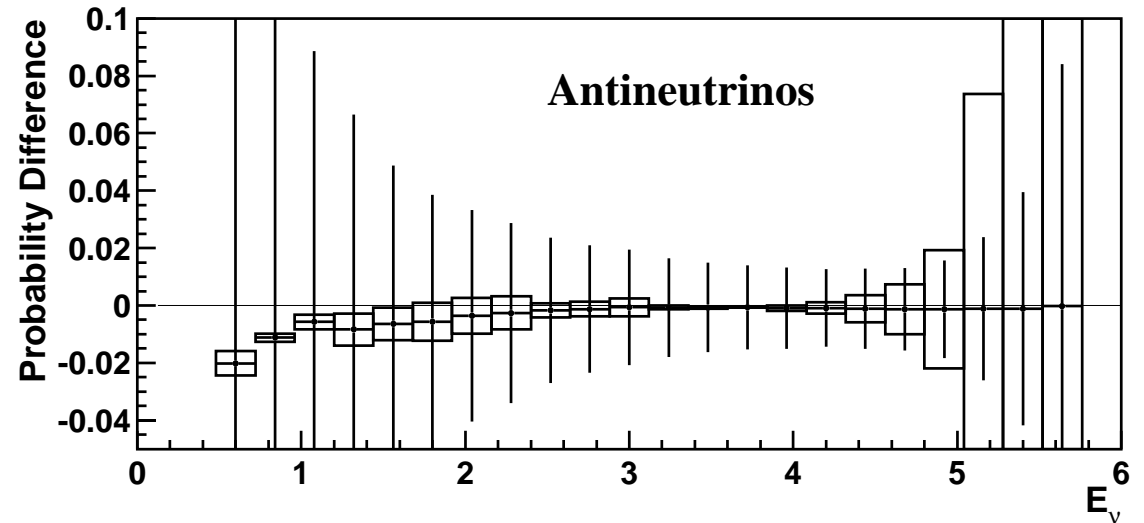
Probability difference

Difference is significant for neutrinos (antineutrinos are matter-suppressed) after evaluation of statistical and systematic errors (5% variation in τ contribution)

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$$

★ Direct measurement of the CP-odd component.

$$P(\nu_e \rightarrow \nu_\mu) - P(\nu_\mu \rightarrow \nu_e)$$



Conclusion

- ★ The richness of the NF oscillation physics calls for *multikton detectors capable of measuring all leptons and their charges!*
- ★ **A difficult task**, leading to different optimizations, but
 - **MINOS**-like (20kt?) & **ICARUS**-like (10kt?) (they would cost \approx the same!) are more than just detector concepts
 - They could be envisaged, but they will be expensive and to build either of them will be a great enterprise !
- ★ The physics output will depend on the detector, different optimizations? e.g.
 - For the **best θ_{13} sensitivity** (10^{-5} ? the small mixing angle syndrome) it would suffice to have large mass (statistics!) and good muon capabilities at the highest energies
 - For more subtle effects like **the study of δ -phase**, a detector with more redundancy and excellent detection in the range 1–15 GeV is favored
 - If **energy cannot be afforded**, must go closer, keeping L/E constant (at the price of increasing the muon intensity!)
 - Charge discrimination for electrons could provide **a “direct” proof** of T-violation !